

# SWARM

DTU

YEAR ANNIVERSARY SCIENCE CONFERENCE

#### Canada's Role in Swarm and the Thermal Ion Imagers

David Knudsen and Johnathan Burchill University of Calgary

Swarm 10 Year Anniversary & Science Conference 2024





## Why Canada?



- Canada is an ESA "Cooperating State"
- Heritage w/ low energy ion measurements

# $\Sigma_{\rm P} = \delta B_{\perp} / (\mu_0 E_{\perp})$ $\vec{S} = \vec{E} \times \delta \vec{H}$

 $\emptyset = \int \vec{E} \cdot \mathrm{d}\vec{l}$ 

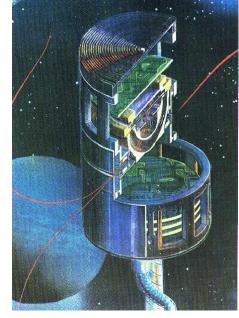
Why E?

## Why ions & TII ?

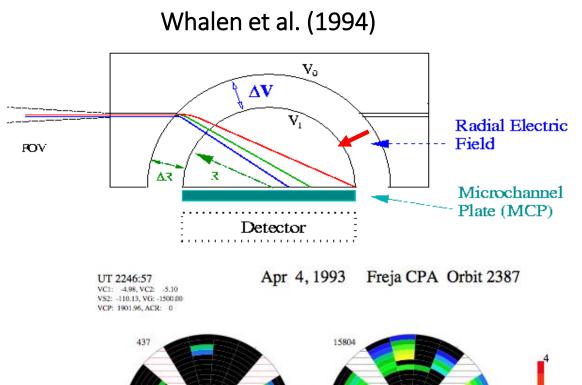
- $\vec{E} = -\vec{v}_i \times \vec{B}$
- Plasma dynamics



## Ion Distribution Imaging







46:57.505 0.545 s

-2.10--21.0 eV Clear need for improved detector:

DTU

· e e sa

- Sounding rockets (2000-2007)
  - ePOP
  - Swarm

Swarm 10 Year Anniversary & Science Conference 2024, 08 – 12 April 2024, CPH Conference, Copenhagen, Denmark

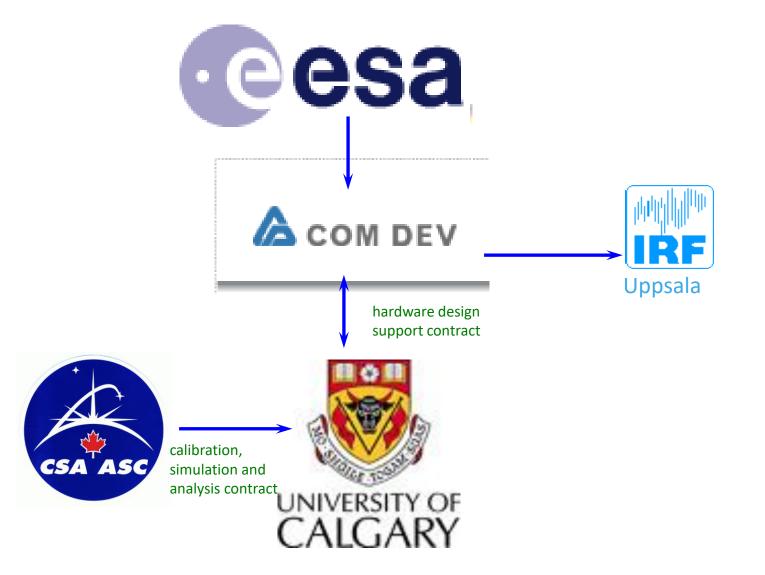
37:13.821 1.186 s



# Swarm EFI (2004+)

DTU

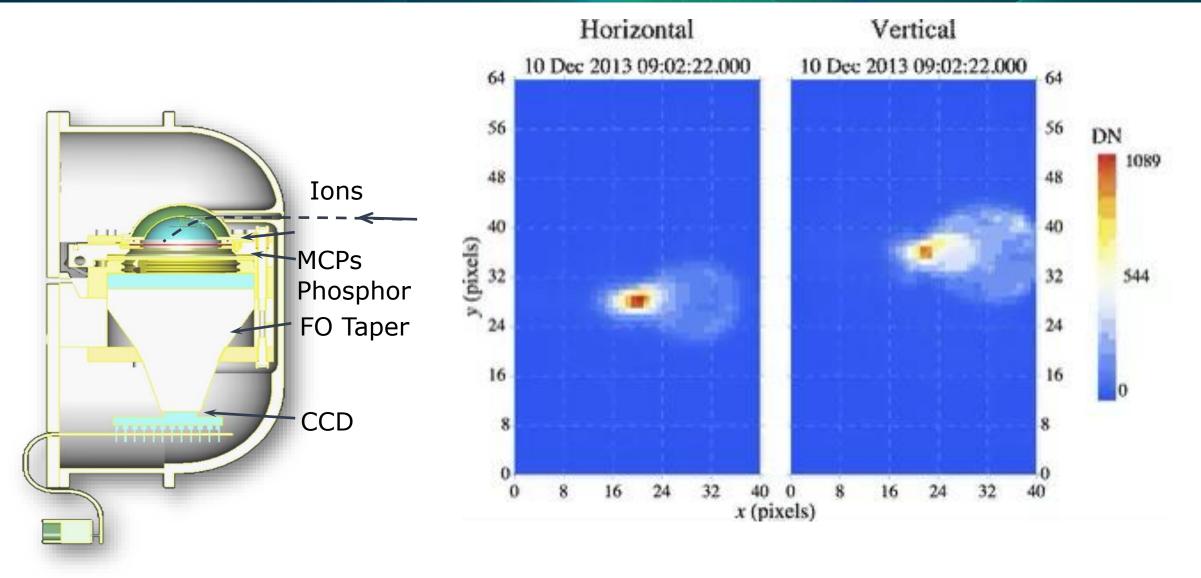
· e e sa



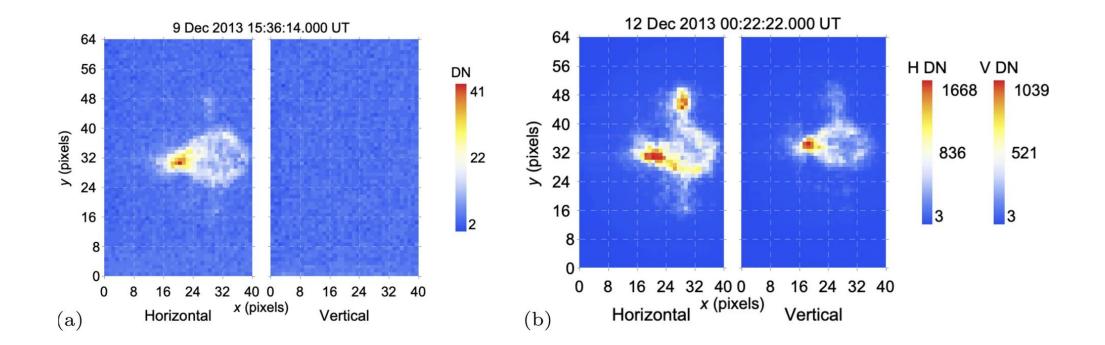


### CCD-based imaging









DTU

· e e sa

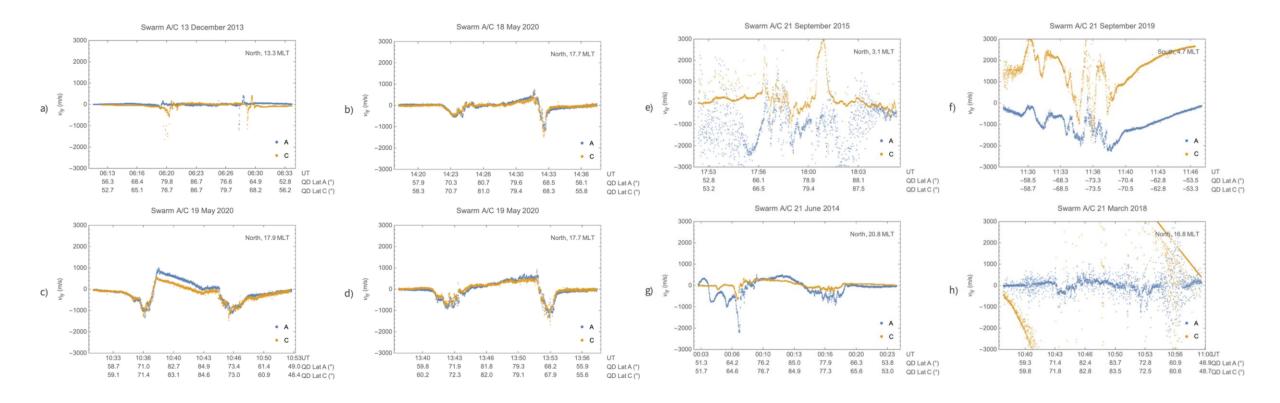
#### Burchill and Knudsen (2002)



## Swarm A/C v<sub>i</sub> comparisons

DTU

· e esa



Burchill and Knudsen (2022)



Science evolution

# Event studies

# Statistical Studies

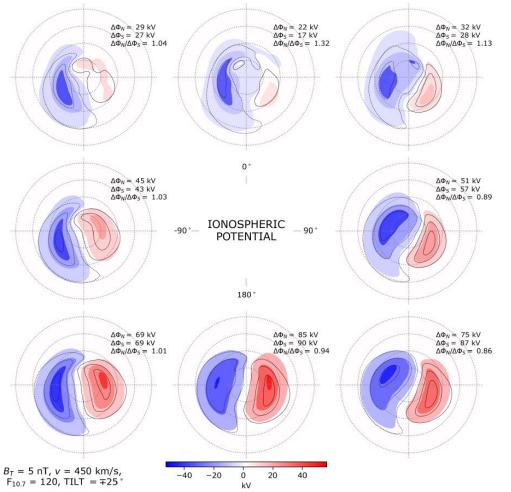
# Models (Data & Physics-based)

DTU



#### 1 - Models

#### Electric potential versus IMF clock angle:



#### Hi-C Potential Model

Hatch et al., ANGEO, in press

• Based on 9 years of Swarm data

DTU

- Spherical harmonic fits
- Can be combined with  $\delta \mathbf{B}$ ...



1.1 – Hi-C + AMPS (Hatch et al., 2024)



J.E Work:

Hi-C (E - Swarm) + AMPS ( $\delta$ B - CHAMP) = "Swipe"

0

Hall conductance

180

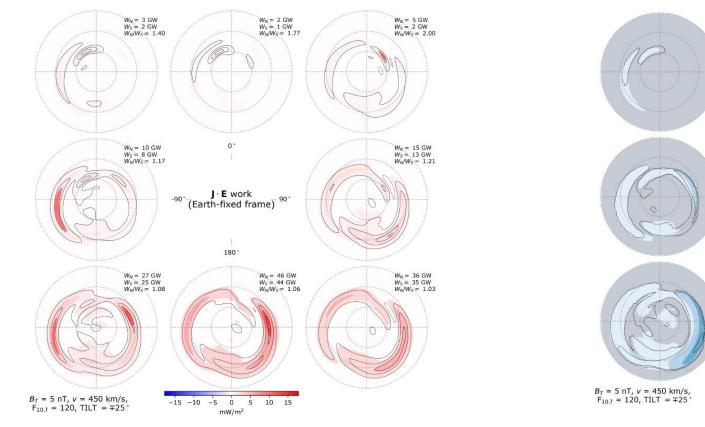
8 12

mho

16 20

-90°

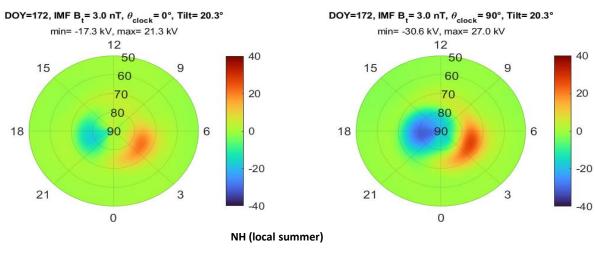
0 4



Hatch et al., *ANGEO*, in press, 2024 Python Package "pyswipe"

### 1.2) Swarm TII-ANN Electric Potential Model

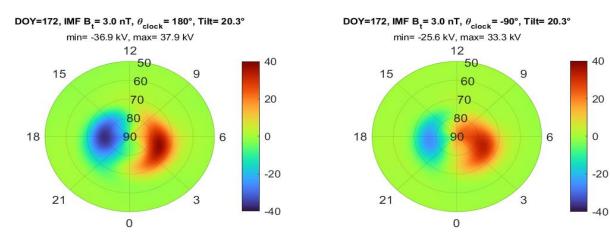
40



UNIVERSITY OF

LGARY

Swarm TII - ANN



Lomidze et al. (Poster #100)

DTU

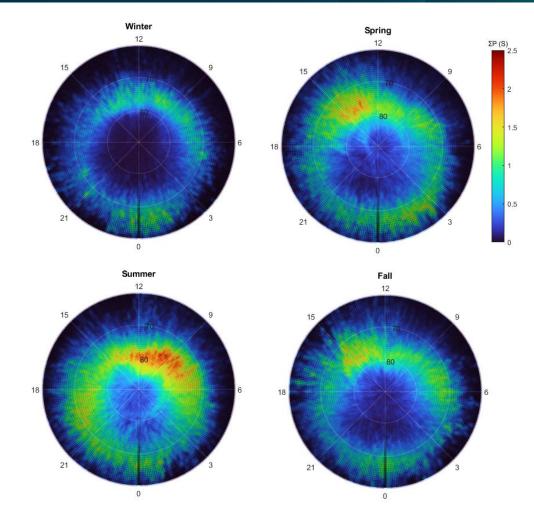
esa

· C

- 9 years of Swarm EFI data (and counting)
- Dependencies on: ۰
  - Season
  - Magnetic Activity
  - Solar Wind
  - Solar Flux (F10.7)
  - Hemisphere ٠



#### 2.1 – Conductance



**Pourkarim**, MSc thesis, 2023 (Poster #63)

DTU

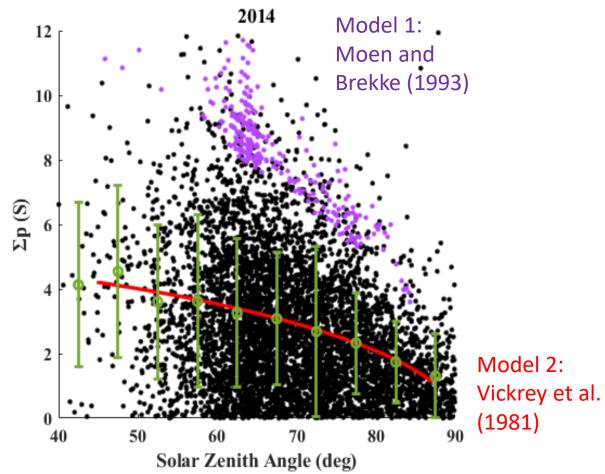
•  $\sum_{P} = \delta B_{V} / \mu_{0} E_{\chi}$ 

- ~500 km altitude
- 9 years
- 10-s bins (76 km)



#### 2.1a – Conductance vs SZA

**Pourkarim**, MSc thesis, 2023 (Poster #63)



• 
$$\sum_{P} = \delta B_{y} / \mu_{0} E_{x}$$

• Huge scatter, but

DTU

· eesa

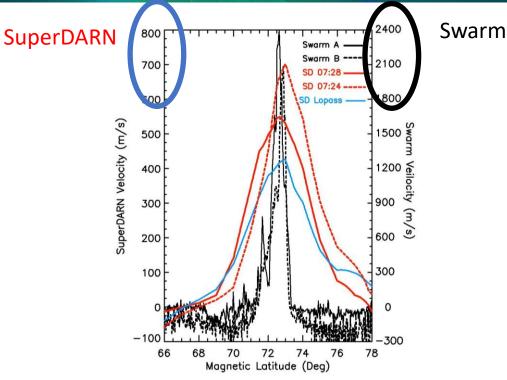
- Reasonable mean
- → Why?
- Assumes:
  - Static fields
  - Sheet-like
  - No **J**<sub>Hall</sub> contrib





3.6) Fenrich, et al., 2021: **Birkeland current boundary flows associated with field line resonances**. J. Geophys. Res., 2021.

3.5) Archer and Knudsen, 2018: Distinguishing Subauroral Ion Drifts From Birkeland Current Boundary Flows, JGR.



3.4) Aikio et al., 2018: Swarm satellite and EISCAT radar observations of a plasma flow channel in the auroral oval near magnetic midnight, *JGR*.

3.3) Archer et al., 2017: Birkeland current boundary flows, JGR.

3.2) Juusola et al., 2016: Ionospheric Conductances and Currents of a Morning-Sector Auroral Arc From Swarm-A Electric and Magnetic Field Measurements, *GRL*.

3.1) Archer et al., Anisotropic core ion temperatures associated with strong zonal flows and upflows, GRL, 2015. Swarm 10 Year Anniversary & Science Conference 2024, 08 – 12 April 2024, CPH Conference, Copenhagen, Denmark



#### 4: Intense Flow Channels - STEVE

DTU

· eesa

2.010 Peak Te at: GLat: -43.39 QDLat: -52.80 Min Ne at: GLat: -43.01 QDLat: -52.36 1.510 Min Vel at: GLat: -43.43 ODLat: -52.37 5 1.010 5.010 1.510 1.010 Te(K) 5.010 UT 10:03:00 GLAT -37.74 GLON 158.04 10:04:00 10:05:00 10:06:00 -41.23 -44 73 48 23 156.44 154.59 152.46 QDLAT -46.38 -50.33 -54.34 -58.40 4.3) Martinis, C., et al., 2022: Rainbow of 10 Horizontal Ion Cross Track (m/ the Night: First Direct Observation of a -4000 km/s! -6000 **SAR arc evolving into STEVE**. *Geophys.* -8000 -10000 UT 10:03:00 GLAT -37.76 GLON 158.07 QDLAT -46.40 10:04:00 -41.24 156.46 -50.34 10:05:00 -44.74 154.62 -54.34 10:06:00 -48.24 152.48 -58.40 Res. Lett., 2022.

Sat. B Time Range: 10:03:00 UT - 10:06:44 UT DATE: 03/17/2015 SWARM B mapped from 520 NH to 425 SH

4.2.) Nishimura, Y. 2019: Magnetospheric signatures of STEVE: Implications for the magnetospheric energy source and interhemispheric conjugacy. *Geophys. Res. Lett*, 2019.

4.1) MacDonald et al., 2018: New Science in Plain Sight: Citizen Scientists Lead to Discovery of Optical Structure in the Upper Atmosphere, Science Advances.





5.6) Wu et al., 2020: Swarm survey of Alfvénic fluctuations and their relation to nightside field-aligned current and auroral arc systems. *JGR*.

5.5) Gillies et al., 2018: A statistical survey of the 630.0 nm optical signature of periodic auroral arcs resulting from magnetospheric field line resonances, GRL.

5.4) Wu et al., 2017: Swarm Observation of Field-Aligned Currents Associated With Multiple Auroral Arc Systems, JGR.

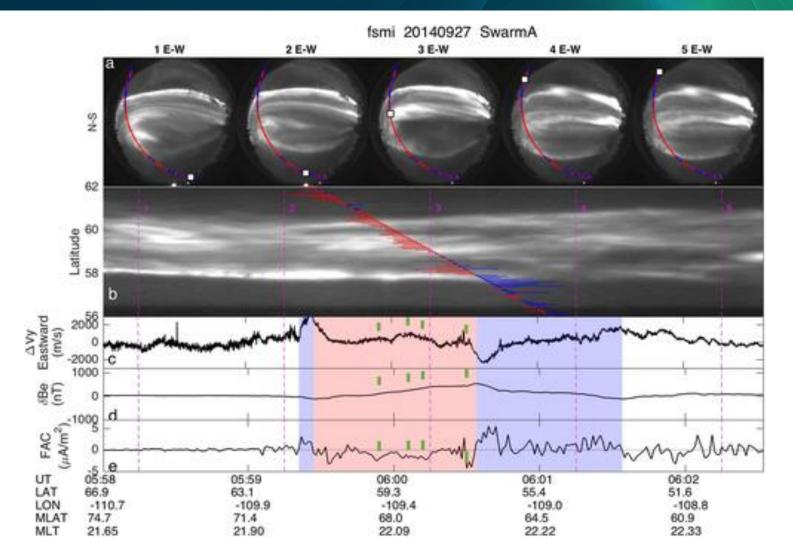
5.3) Gillies et al., 2015: Swarm observations of field-aligned currents associated with pulsating auroral patches, *JGR*.

5.2) Juusola et al., 2016: Ionospheric Conductances and Currents of a Morning-Sector Auroral Arc From Swarm-A Electric and Magnetic Field Measurements, *GRL*.

5.1) Aikio et al., 2018: Swarm satellite and EISCAT radar observations of a plasma flow channel in the auroral oval near magnetic midnight, *JGR*.



#### 5.4: Auroral Arcs - Example



5.4) Wu et al., 2020: **Swarm survey of Alfvénic fluctuations and their relation to nightside field-aligned current and auroral arc systems**. *JGR*.

DTU

· e e sa



# 6: Alfvén waves



6.6) Ghadjari, H. et al., Post-sunset field-line resonances at equatorial latitudes observed by Swarm, GRL, 2023.

6.5) Ghadjari, H. et al., 2022: **Standing Alfvén waves within equatorial plasma bubbles**. *Geophys. Res. Lett*.

6.4) Ivarsen et al., 2023: **Observational evidence for the role of Hall conductance in Alfvén wave reflection**, *JGR*.

6.3) Wu et al., 2020: Swarm survey of Alfvénic fluctuations and their relation to nightside field-aligned current and auroral arcs systems, JGR.

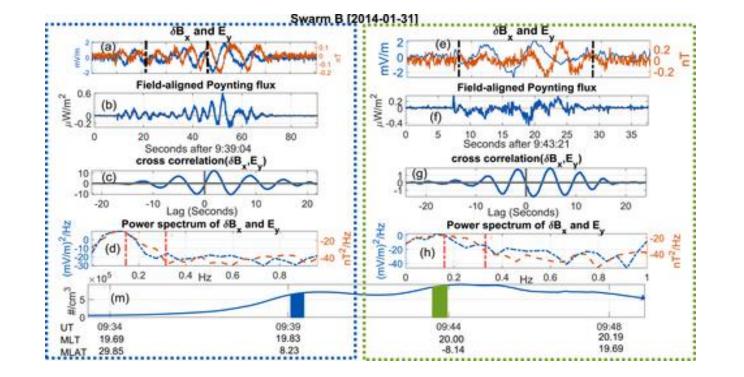
6.2) Miles et al., Alfvénic dynamics and fine structuring of discrete auroral arcs: Swarm and e-POP observations, *GRL*, 2018.

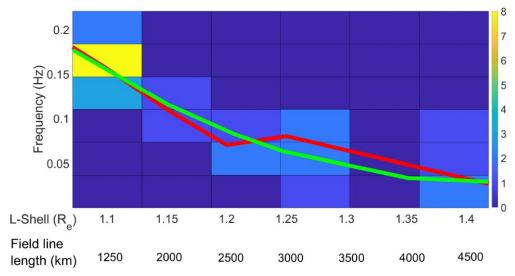
6.1) Pakhotin et al., Diagnosing the Role of Alfvén Waves in
Magnetosphere-Ionosphere Coupling: Swarm Observations of Large Amplitude
Nonstationary Magnetic Perturbations During an Interval of Northward IMF, JGR, 2018.



#### 6: Alfvénic field-line resonance - example







DTU



7.7) Billett et al., 2023: Multi-scale Ionospheric Poynting Fluxes Using Ground and Space-Based Observations, *JGR*.

7.6) Billett, D.D., et al. High-resolution Poynting flux statistics from the Swarm mission: How much is being underestimated at larger scales? *JGR*, 2022.

7.5) Rodríguez-Zuluaga, et al. 2022. **Topside equatorial spread F-related field-aligned Poynting flux: observations and simulations**. *Earth, Planets and Space*, *74*(1), 2022.

7.4) Pakhotin, I.P., Mann, I.R., Xie, K. *et al.* Northern preference for terrestrial electromagnetic energy input from space weather. *Nat Commun.*, (2021).

7.3) Rodríguez-Zuluaga et al, **On the direction of the Poynting flux associated with equatorial plasma depletions as derived from Swarm**, *GRL*, 2017.

7.2) Park et al., Alfvén waves in the auroral region, their Poynting flux, and reflection coefficient as estimated from Swarm observations, *JGR*, 2017

Swarm, GRL, 2017.

7.1) Park et al., Statistical survey of nighttime midlatitude magnetic fluctuations: Their source location and Poynting flux as derived from the Swarm constellation, *JGR*, 2016.





8.1) Spicher et al., Observation of polar cap patches and calculation of gradient drift instability growth times: A Swarm case study, *GRL*, 2015.

8.2) Goodwin et al., Swarm in situ observations of F-region polar cap patches created by cusp precipitation, *GRL*, 2015.

8.3) Zou et al., Localized field-aligned currents in the polar cap associated with airglow patches, *JGR*, 2016.





9.5) Burchill and Knudsen (2022): **Swarm Thermal Ion Imager measurement performance.** *Earth, Planets and Space.* 

9.4) Lomidze et al., 2019: Validity study of the Swarm horizontal cross-track ion drift velocities in the high-latitude ionosphere, *Earth and Space Science*. Ion drifts consistent with Weimer (2005)

9.3) Koustov et al., 2018: A comparison of cross-track ion drift measured by the Swarm satellites and plasma convection velocity measured by SuperDARN. *JGR*.

9.2) DJ Knudsen, JK Burchill, SC Buchert, AI Eriksson, Reine Gill, J-E Wahlund, Lennart Åhlén, M Smith, B Moffat., Thermal ion imagers and Langmuir probes in the Swarm electric field instruments, *JGR*, 2017.

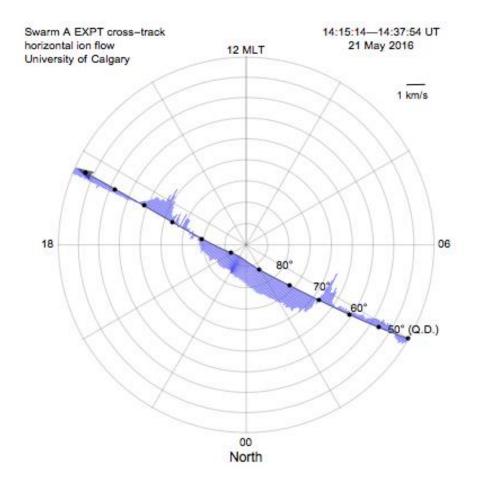
9.1) Fiori et al., 2016: Calibration and assessment of Swarm ion drift measurements using a comparison with a statistical convection model, *Earth, Planets and Space.* 



• Cross-track velocity available through January 2024. See esa.int/Swarm

#### Sincere thank you to:

- Roger, Rune, Anja, MAG
- ESA Project Team, FOS, PLSO, ARB board
- Swarm DISC & EFI Science Discussion Group
- Airbus, COM DEV/Honeywell
- Canadian Space Agency
- Alexei Kouznetsov, Levan Lomidze



DTU