

SWARM

10

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## A geomagnetic field model incorporating Earth's mantle conductivity

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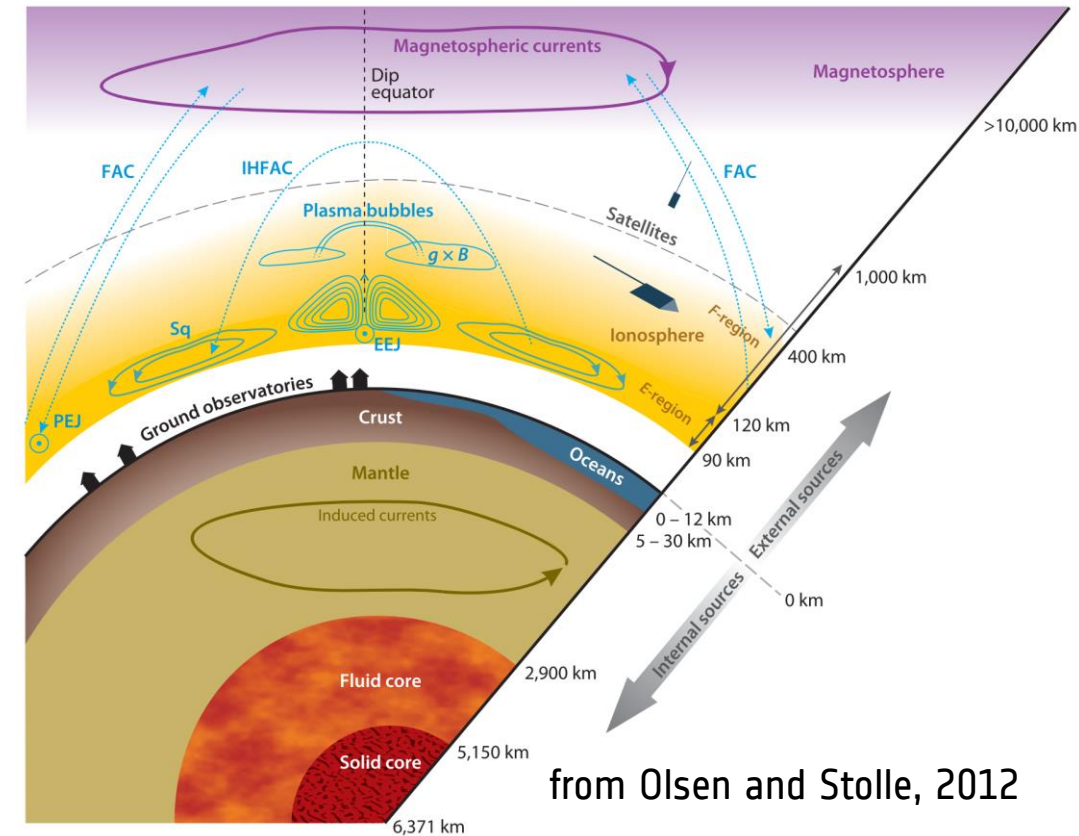
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# Swarm 10 Year Anniversary & Science Conference 2024



## Sources of near-Earth magnetic Field

- **Magnetic measurements:**
  - Core dynamics
  - Lithospheric magnetization
  - Ocean circulation and tides
  - Ionospheric currents and electromagnetic (EM) induction in the Earth
  - Magnetospheric currents and EM induction
  - Ionosphere-magnetosphere coupling currents
- **Magnetic field models:**
  - Separate different sources by geomagnetic field modeling
  - Those rapidly time-varying magnetospheric primary and Earth induced fields play important role in high-resolution geomagnetic field modeling



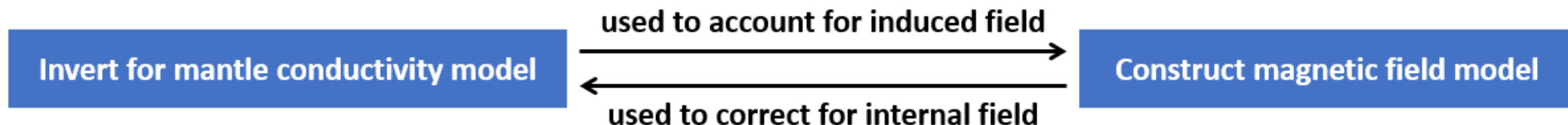
from Olsen and Stolle, 2012

Sources of near-Earth magnetic Field



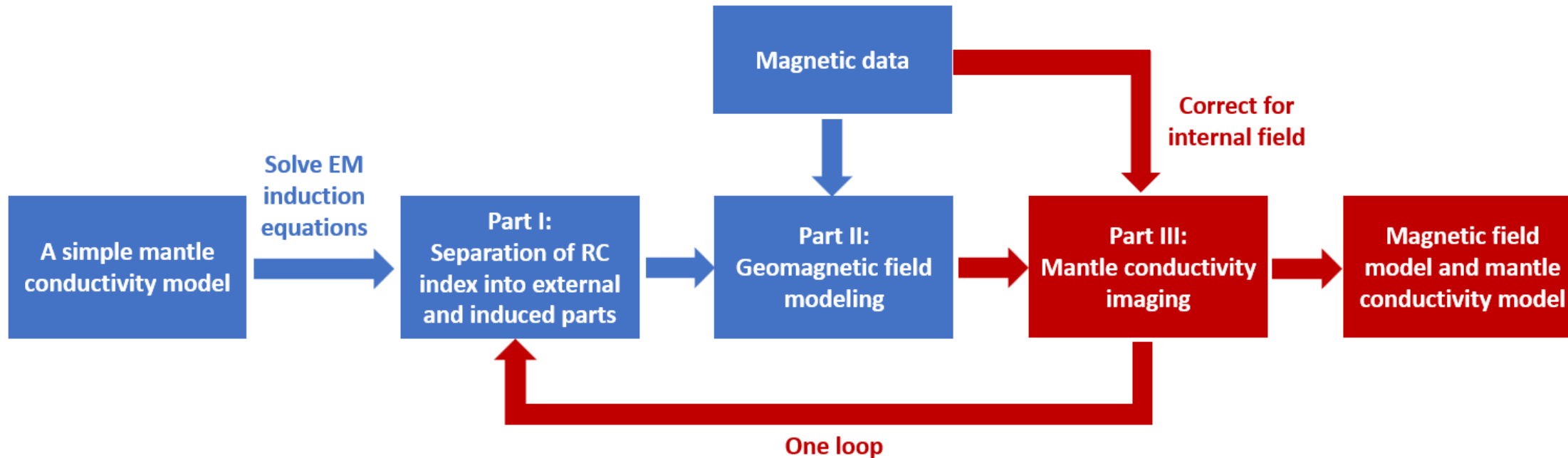
## How does CM and CHAOS models deal with magnetospheric primary and induced fields?

- Spatial: by spherical harmonic (SH) expansion
- Temporal:
  - CM model series (CM5, CM6): by a series of 1-hour time bins
    - EM induction equations are not invoked
    - Many parameters are needed (in CM6, magnetosphere/total =  $685110/734670 \approx 93\%$ )
  - CHAOS model series: by a linear function of the external and induced parts of RC index
    - RC index is separated into external and induced parts using a prior mantle conductivity model and EM induction equations
    - Less parameters are needed (in CHAOS-7, magnetosphere/total =  $613/31757 \approx 2\%$ )





## Our approach - incorporating Earth's mantle conductivity during geomagnetic field modeling



### Advantages:

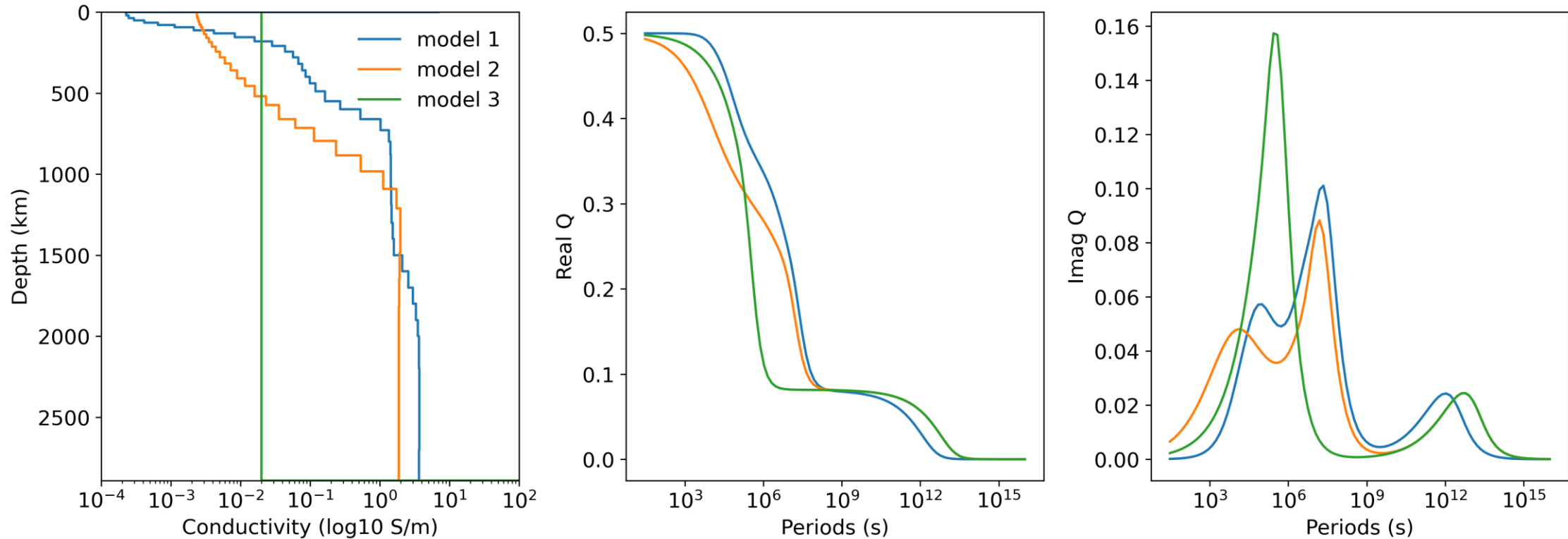
- Constrain geomagnetic field modeling with equation of physics
- Treat geomagnetic field modeling and mantle conductivity imaging in a more self-consistent manner



- Given a mantle conductivity model  $\sigma$ , compute EM responses  $Q_1(\omega, \sigma)$  by solving Maxwell's equations using a finite-element code (Yao et al., 2022).
- Separate the RC index into external and internal parts using computed EM responses (Maus and Weidelt, 2004; Olsen et al., 2005)

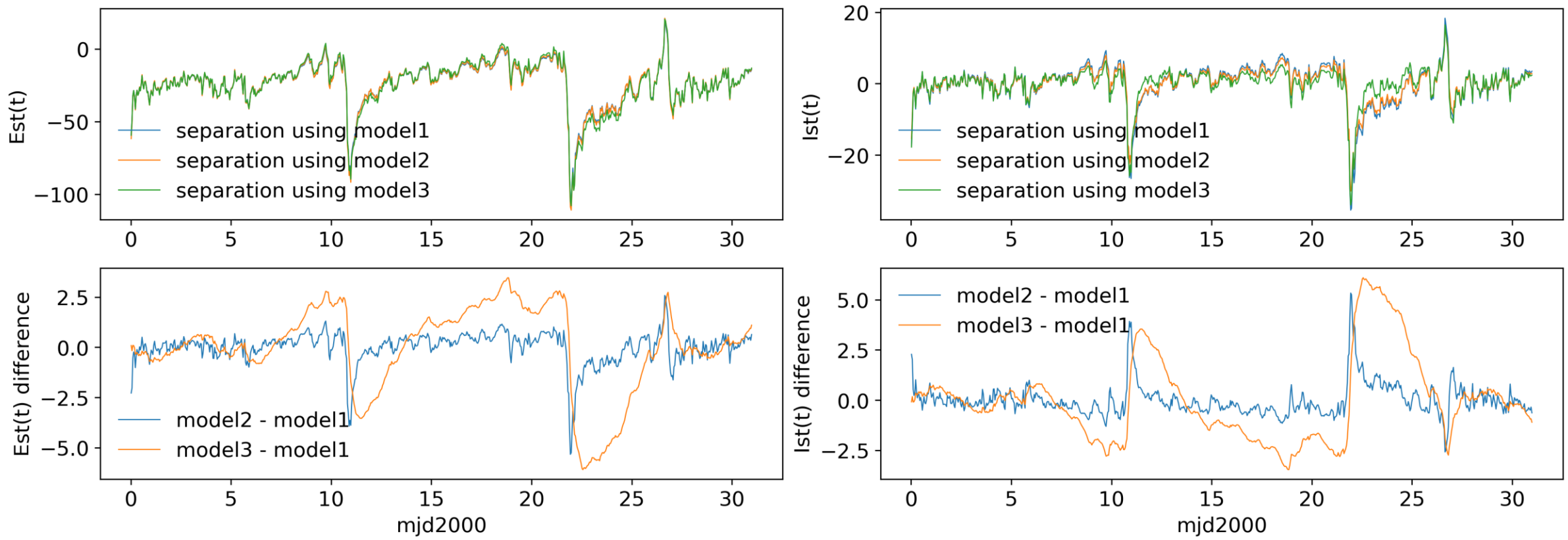
$$RC(t) = Est(t) + Ist(t, \sigma) = Est(t) + Q_1(t, \sigma) * Est(t)$$

- The time variation of magnetospheric primary and induced fields are expressed by a linear function of the external  $Est$  and induced  $Ist$  parts of RC index (Olsen et al., 2005).



**Illustration of EM responses for different Earth's conductivity models  
(model 1: Grayver et al., 2017; model 2: Yao et al., 2023a; model 3: half-space model)**

# Method – Part I: separation of RC index into external and induced parts



**Illustration of RC separation using different Earth's conductivity models**



- **Model parametrization:** core + lithosphere + magnetosphere (Olsen et al., 2006 ):

$$\mathbf{B} = -\nabla V, V = V^{\text{int}} + V^{\text{ext}}$$

$$V^{\text{int}} = a \sum_{n=1}^{N_{\text{int}}} \sum_{m=0}^n (g_n^m(t) \cos m\phi + h_n^m(t) \sin m\phi) \left(\frac{a}{r}\right)^{n+1} P_n^m(\cos \theta)$$

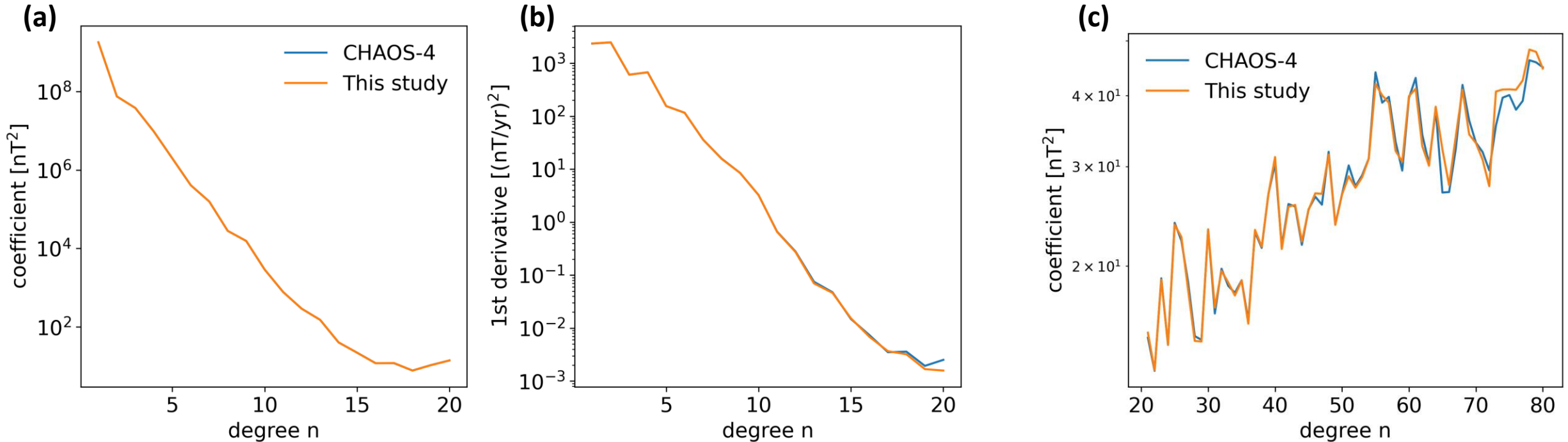
$$V^{\text{ext}} = a \sum_{n=1}^2 \sum_{m=0}^n (q_n^{m,SM}(t) \cos m\phi_{SM} + s_n^{m,SM}(t) \sin m\phi_{SM}) \left(\frac{r}{a}\right)^n P_n^m(\cos \theta_{SM}) + a \sum_{n=1}^2 q_n^{0,GSM} \left(\frac{r}{a}\right)^n P_n^0(\cos \theta_{GSM})$$

- **Model estimation:** we implemented a robust least squares method with Gauss-Newton iteration to minimize the objective function, which includes a data misfit term and two regularization terms (Olsen et al., 2014).





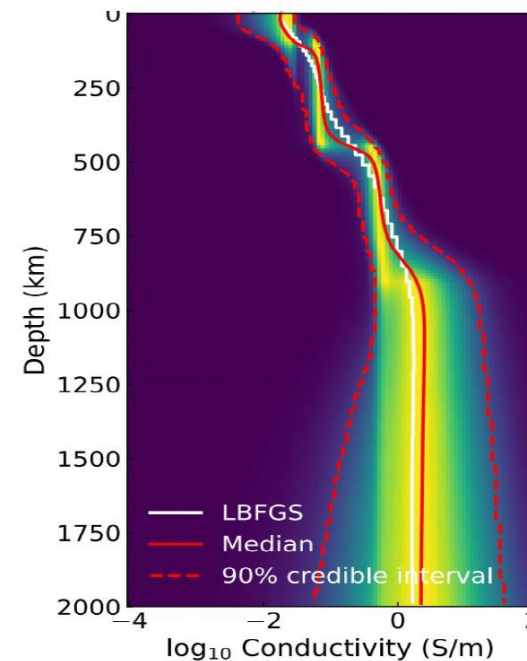
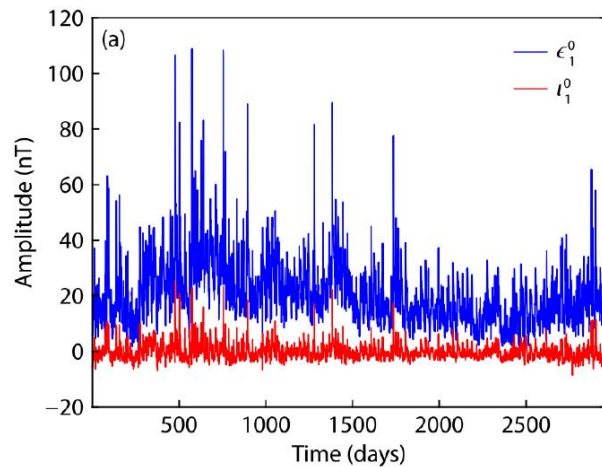
## Validation- reproducing CHAOS-4 model using a new Python code



**Mauersberger-Lowes spectrum of the (a) core field; (b) secular variation; and (c) lithospheric field of the CHAOS-4 model (Olsen et al., 2014) and the model derived with our code**



- Correct for the internal field using the derived geomagnetic field model.
- Invert/update for mantle conductivity model using a trans-dimensional Bayesian inversion code (Yao et al., 2023b).
- Update the magnetic field model using the updated mantle conductivity



**Examples of extracted magnetospheric SH coefficients and recovered mantle conductivity model**



- **Data**

- Swarm (2023.01-12) and MSS-1 (2023.05-12, launched on May 21, 2023) vector and scalar magnetic data.
- Only night-time data from geomagnetically quiet times were used
- 18505 scalar data and 80140 vector data

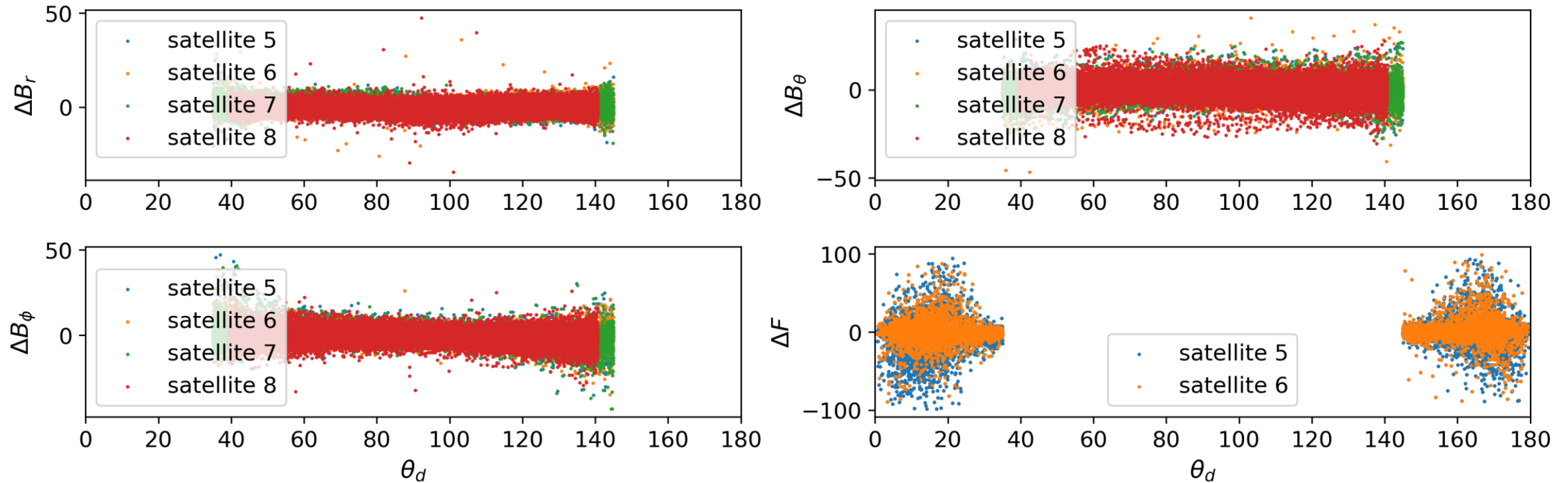
- **Modeling parameters:**

- A half-space conductivity model (0.2 S/m) was used for the initial RC separation.
- Internal fields were estimated up to degree  $n=30$  ( $n=1-15$  is time varying)
- 1835 model parameters (internal + external + induced)
- Loop convergence condition: the root-mean-square (rms) error of the magnetic field modeling between adjacent loops is less than  $10^{-2}$ .



## Magnetic data residuals

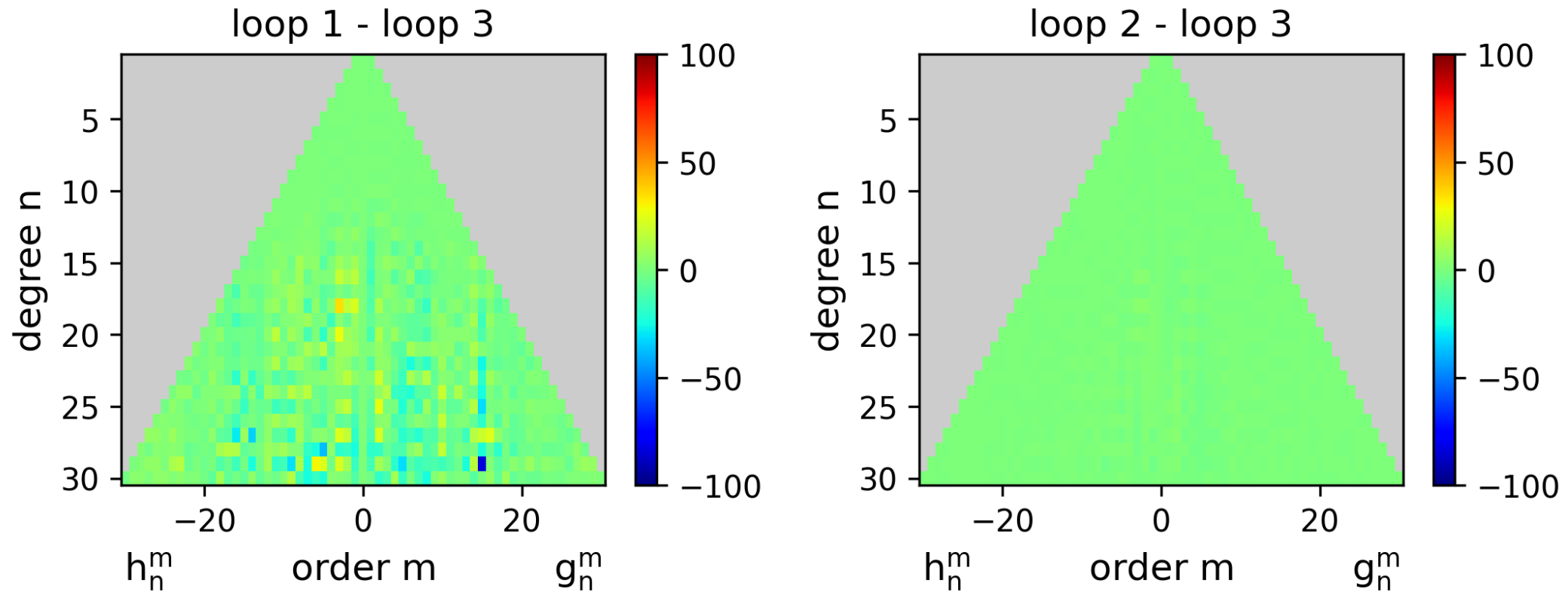
The iteration converges after 3 loops, here are the magnetic field residuals at the last loop:



**Magnetic data residuals of geomagnetic field modeling at the last loop  
(satellite 5-8 denote Swarm A, B, C, and MSS-1, the final Huber-weighted rms is 3.6)**



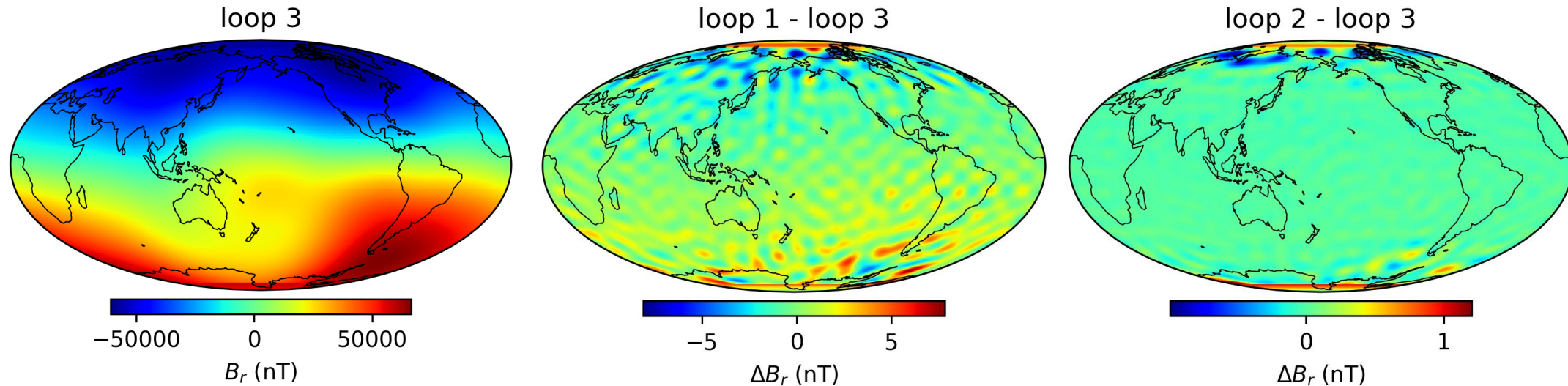
## Internal field for different loops



**Relative differences of the estimated internal field coefficients between different loops**



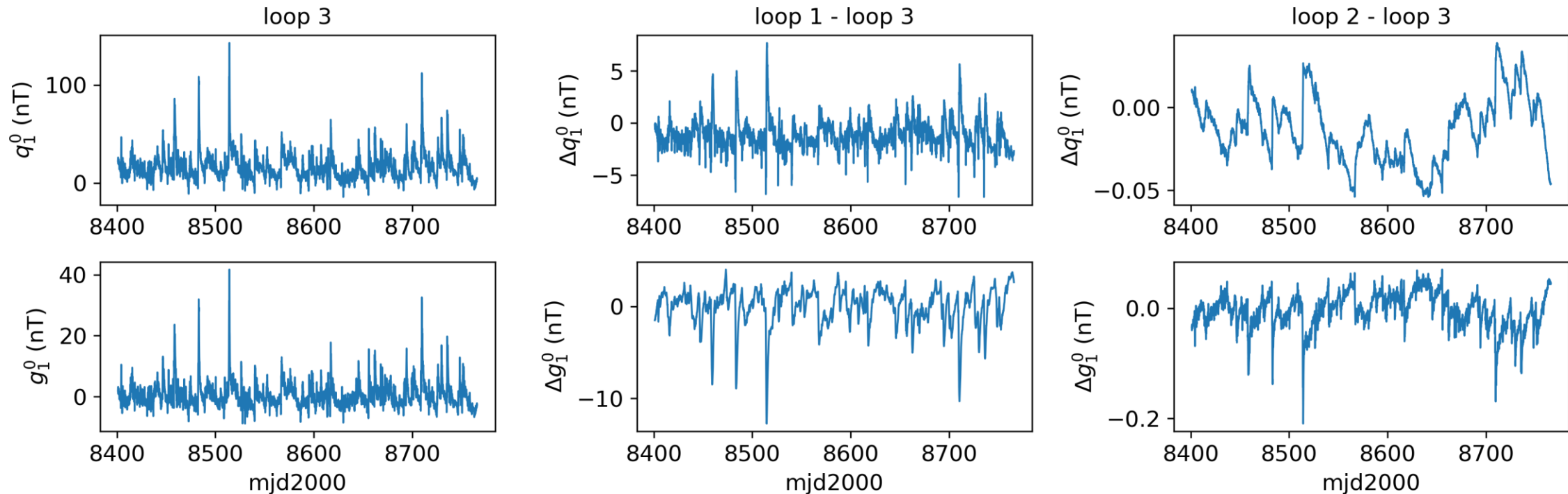
## Internal field for different loops



**Global maps of internal field ( $n=1-30$ , epoch=2023.75) at the Earth's surface and the corresponding differences between different loops**



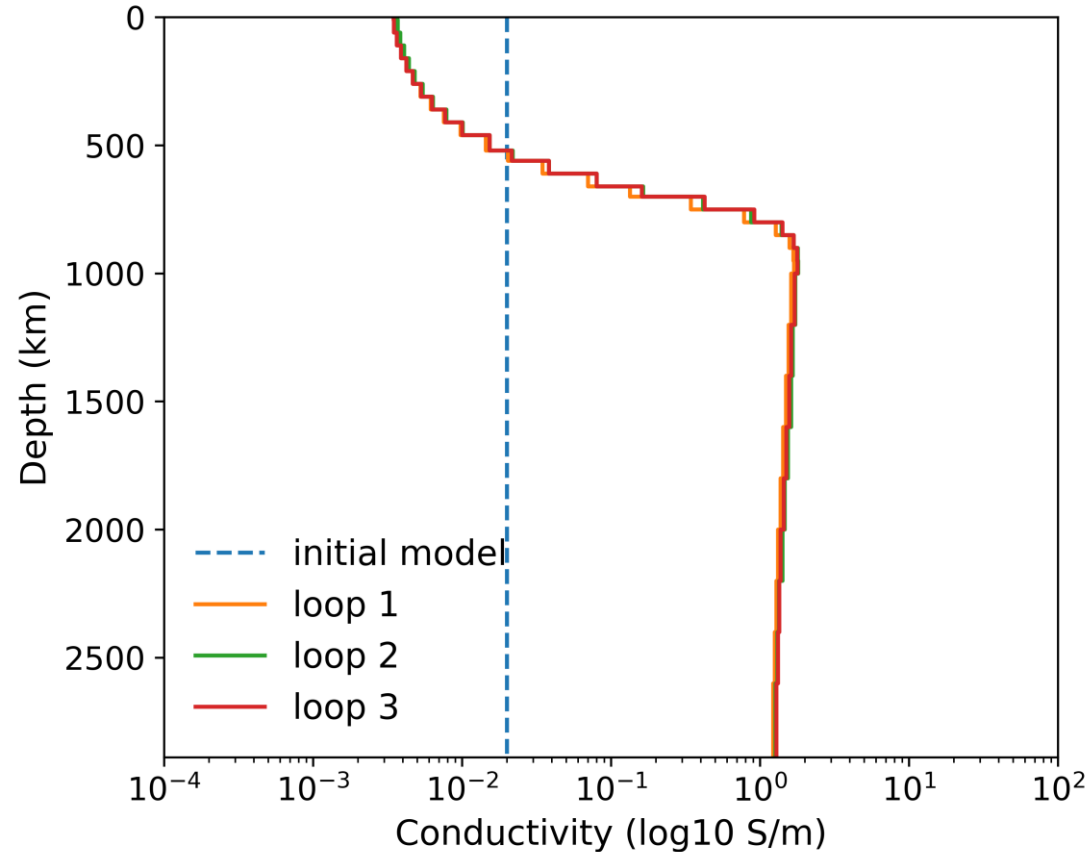
## External field for different loops



**Time series of external and induced coefficients describing large-scale magnetospheric ring current and the corresponding differences between different loops**



## Mantle conductivity models for different loops



**Initial conductivity model used for RC external and induced separation and the recovered mantle conductivity profiles for different loops**





- We presented a new method for geomagnetic field modeling that incorporates Earth's mantle conductivity as model parameter.
- Initial results show that inappropriate treatment of induced field may cause error more than 5 nT in both internal and external field models, while our method can iteratively reduce these errors.
- This method may be extended to incorporate more physical processes into geomagnetic field modeling, such as the large-scale ocean circulation and tidal flows.



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