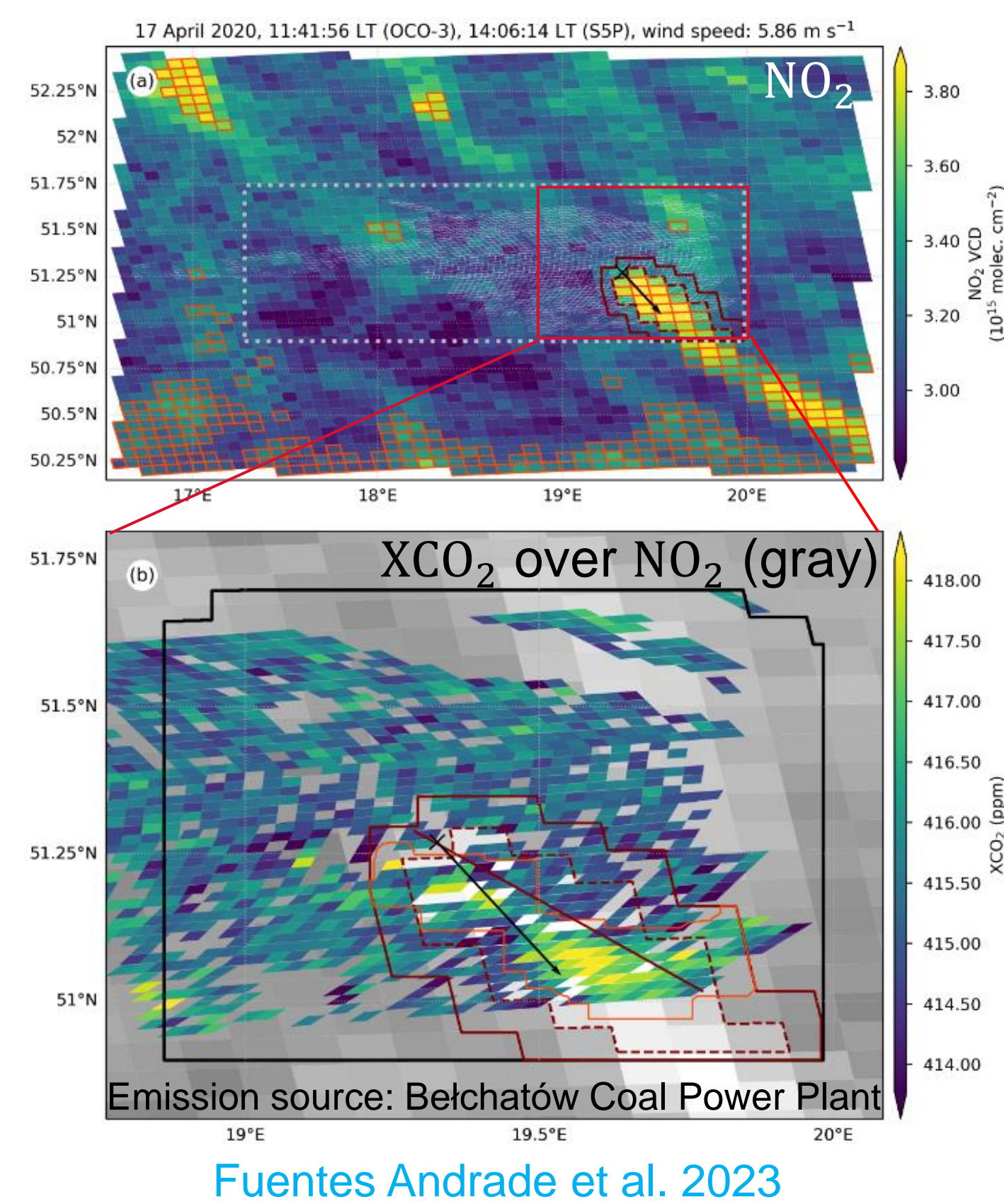


# Improved tropospheric NO<sub>2</sub> columns retrieved from TROPOMI on board the Sentinel-5P satellite

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## 1. Motivation

CO<sub>2</sub> is the most important anthropogenic greenhouse gas. Detecting and quantifying CO<sub>2</sub> emission sources is important but difficult using satellite data. NO<sub>2</sub> is co-emitted with CO<sub>2</sub> during combustion processes and is much easier to detect. Therefore, it can be used as a proxy for CO<sub>2</sub> emissions. The objective is to establish a high resolution TROPOMI NO<sub>2</sub> for Europe, optimised for small scale processes, like emission plumes.



## 2. NO<sub>2</sub> columns and AMF

- Spectra measured by TROPOMI
- Analysed using DOAS → slant columns (SC)
- Stratospheric SC subtracted to obtain tropospheric SC
- Transformed into vertical columns using air mass factor (AMF)
- $AMF \stackrel{def}{=} \frac{SC}{VC}$ , a priori information for calculation used e.g. surface reflectivity
- Has influence on BAMF
- Idea: replace operational DLER reflectivity product from TROPOMI with BRDF product from MODIS
- MODIS dataset has higher temporal and spatial resolution

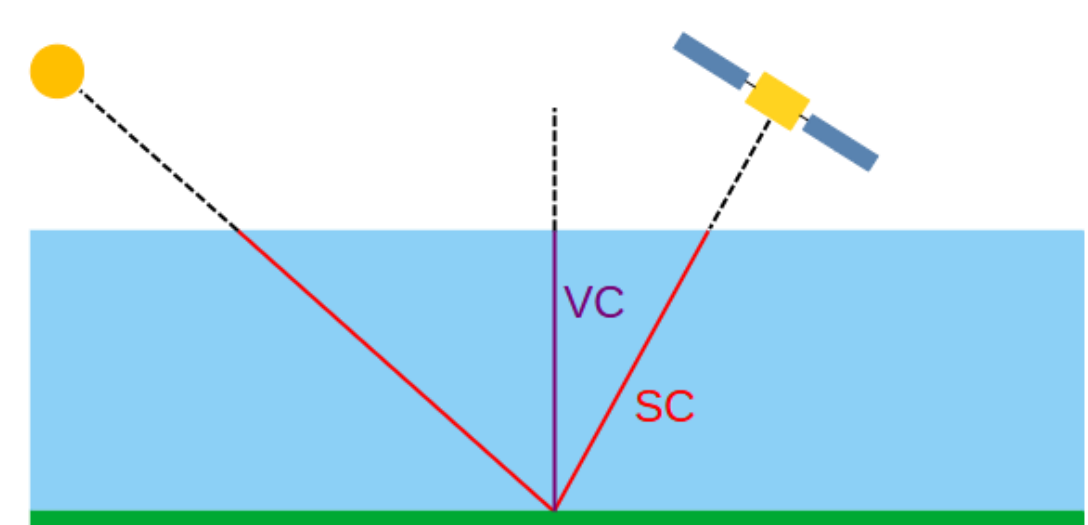
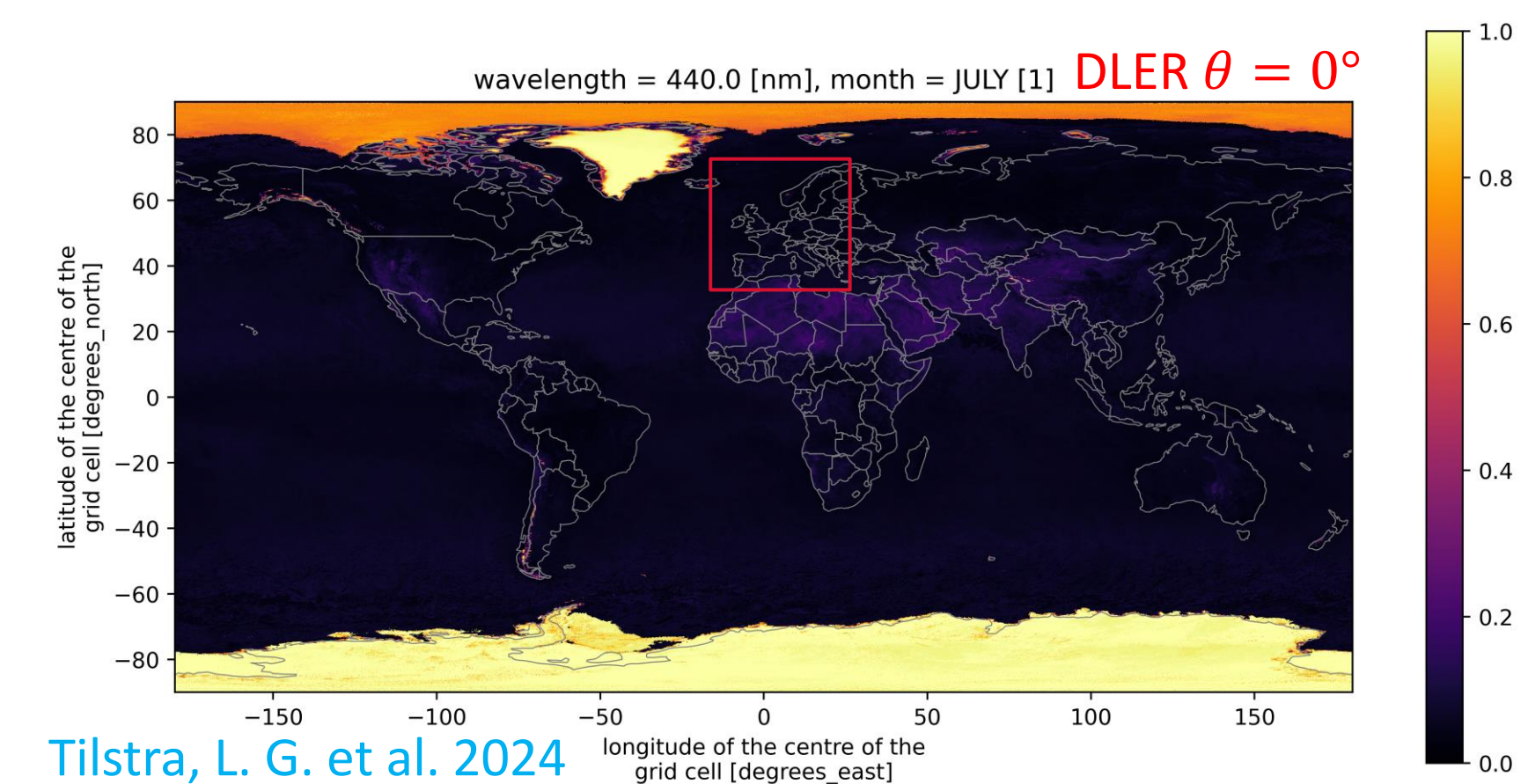
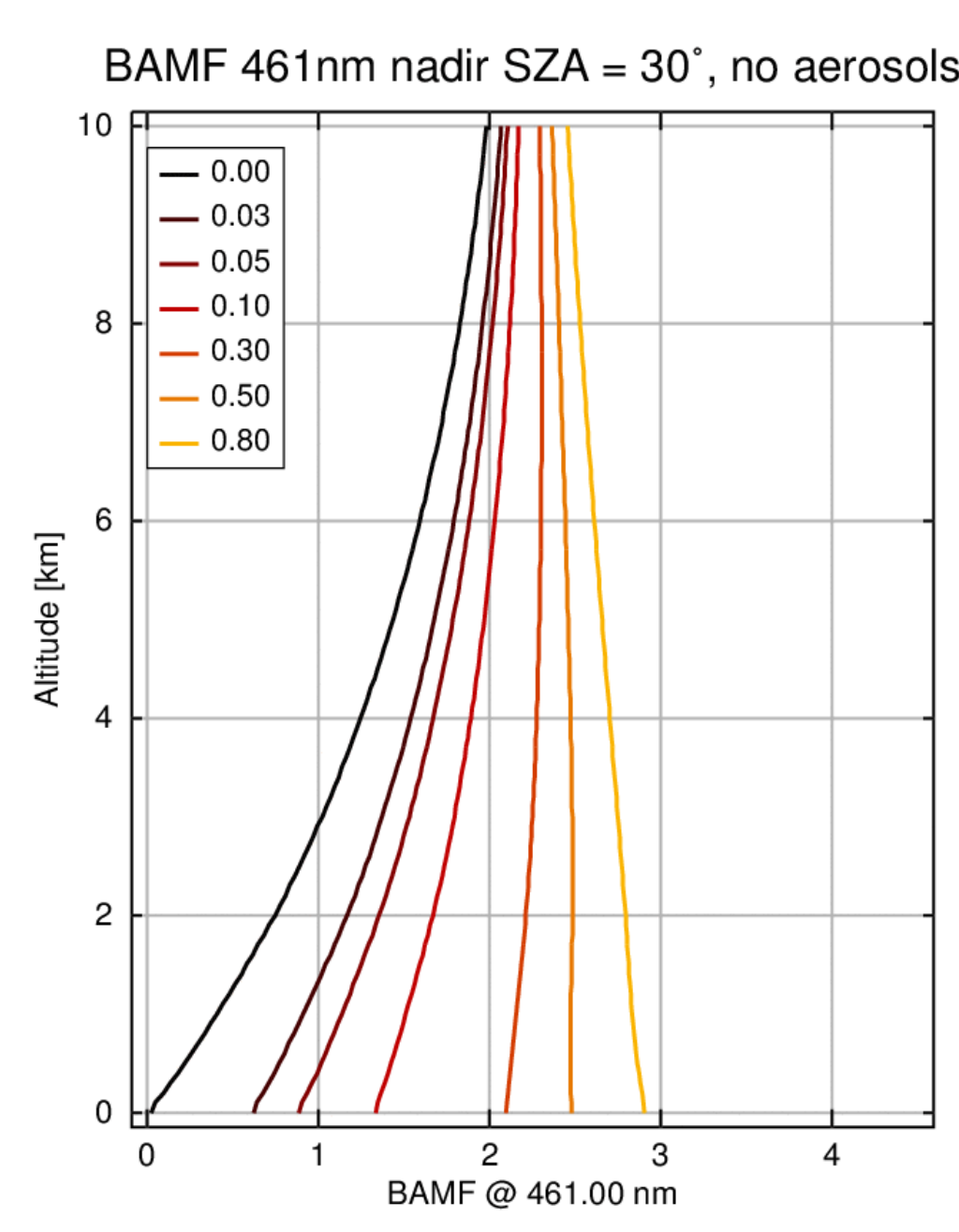


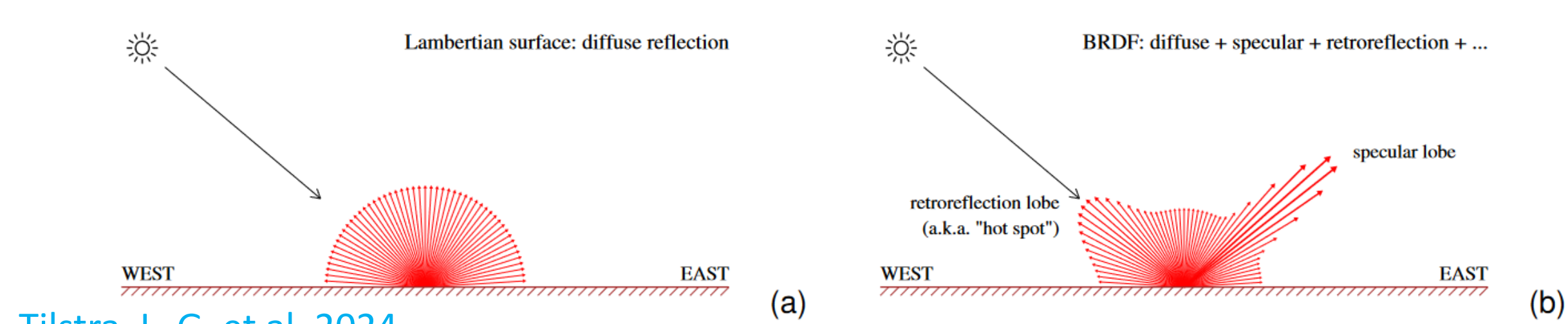
Illustration of slant and vertical columns.



Surface reflectivity of DLER product shows low reflectivities in target area (red)

$$BRDF(\lambda, \theta, \theta_0, \varphi - \varphi_0) = f_{iso}(\lambda) + f_{vol}(\lambda) * k_{vol}(\theta, \theta_0, \varphi - \varphi_0) + f_{geo}(\lambda) * k_{geo}(\theta, \theta_0, \varphi - \varphi_0)$$

$f_{iso}$ ,  $f_{geo}$  and  $f_{vol}$  obtained from MODIS



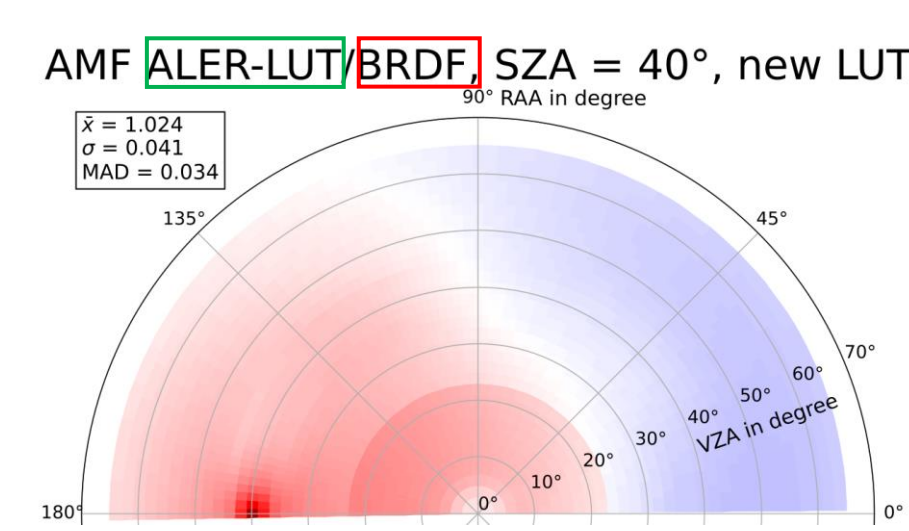
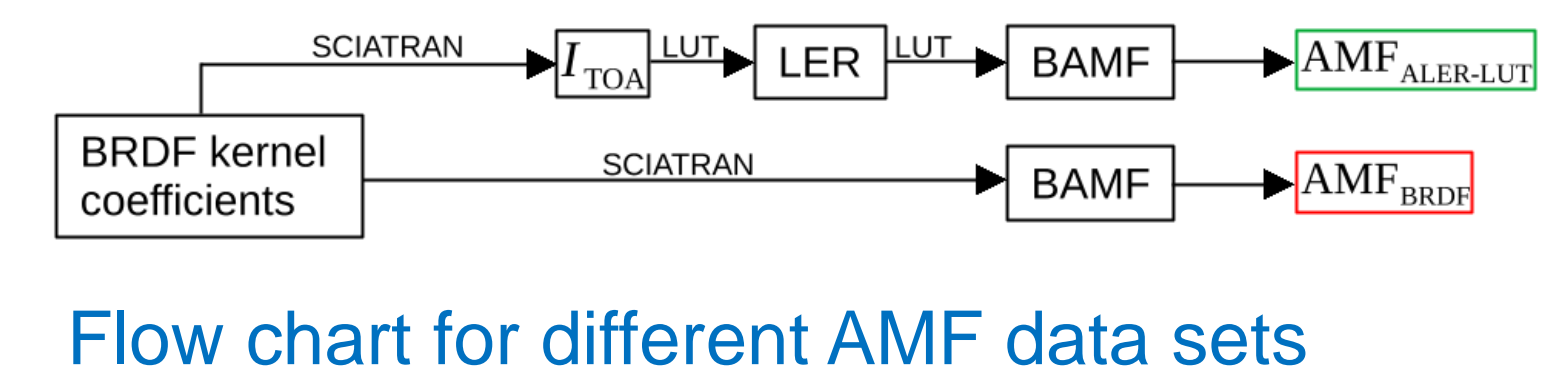
Tilstra, L. G. et al. 2024

a) Lambertian equivalent reflectivity (LER) amount of reflected light in all directions the same

b) Bi-directional reflectance distribution function (BRDF), describes angle-dependent reflectivity of surfaces

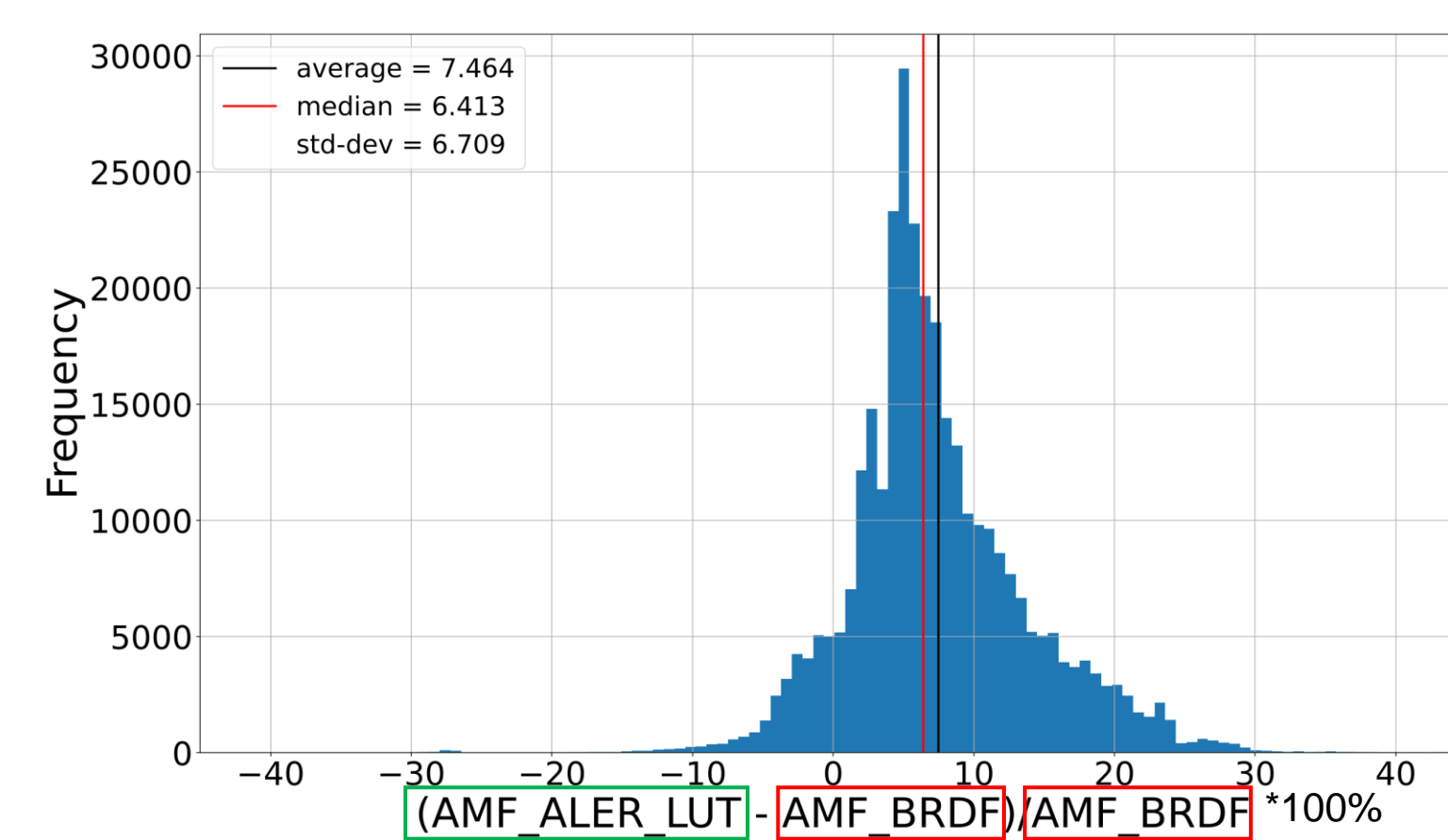
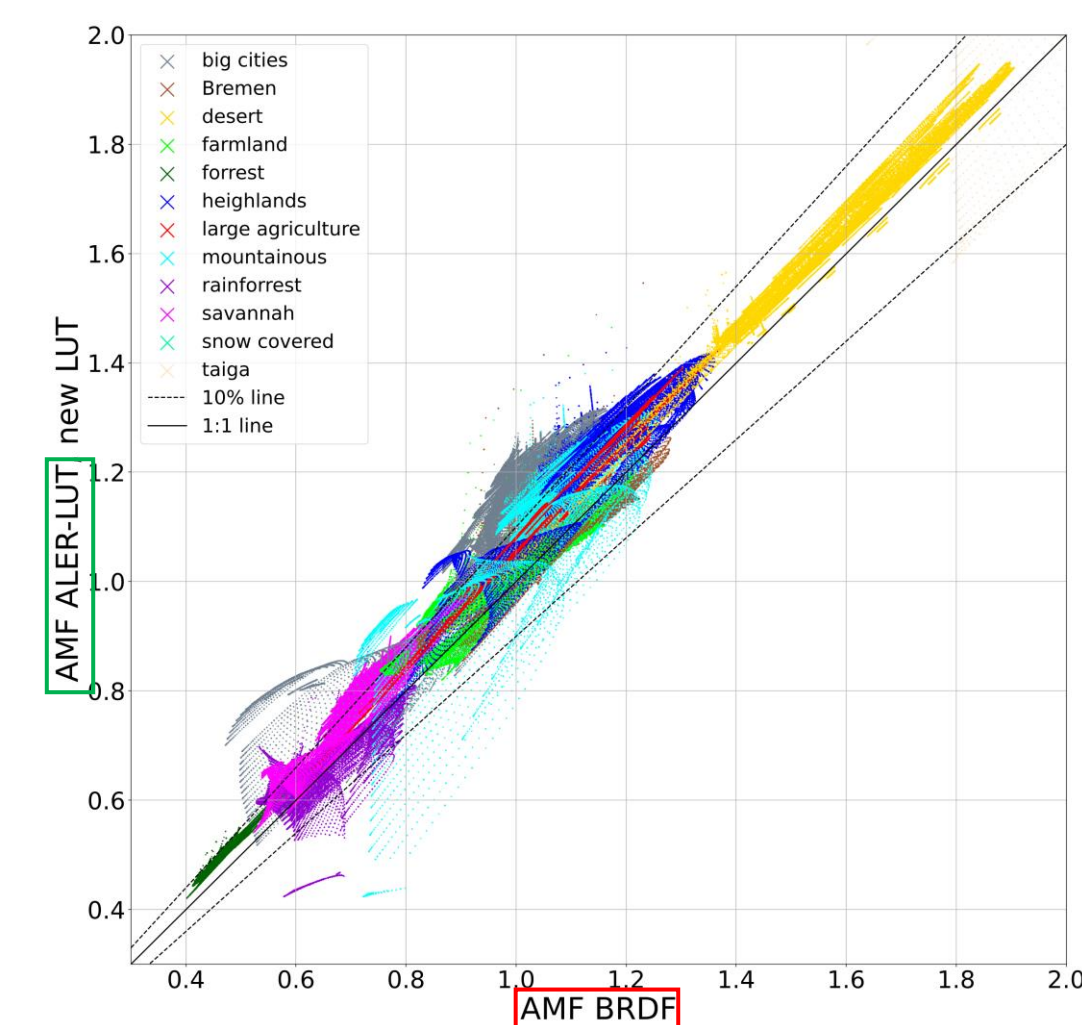
## 3. Results

- AMF calculation using SCIATRAN for the AMF calculation using the BRDF for all TROPOMI pixels and orbits takes too long
- Shortcut: use LUT and interpolation, as described by Vasilkov et al. (2017)



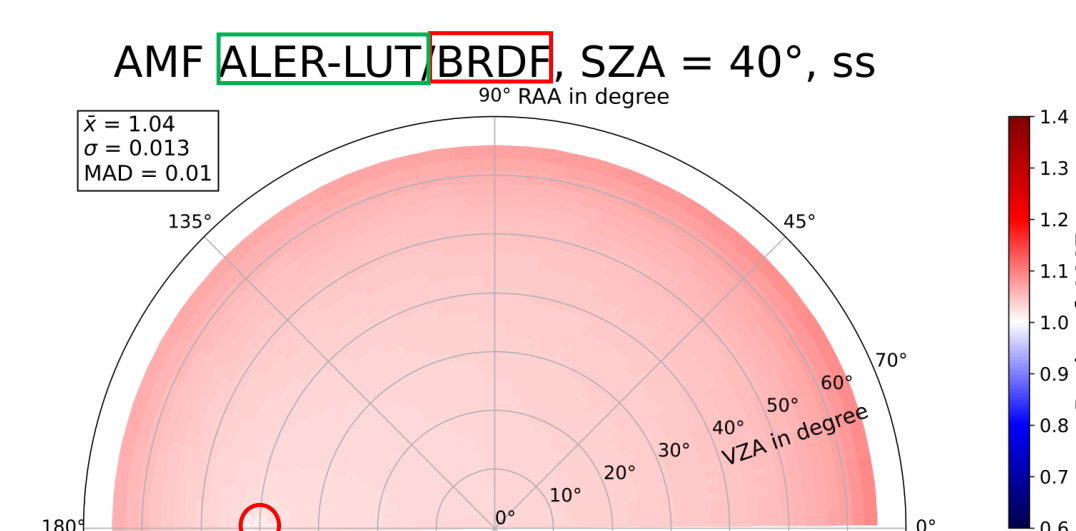
Pixel by pixel ratio of the AMF for Vasilkov's method and using the BRDF.

- AMF\_BRDF and AMF\_ALER-LUT comparison shows problems:
  - Ring structure around VZA ~4° - ~22°
  - Hotspot at retroreflection
  - Ratio is RAA dependent
- Check more than one scene in a scatter plot

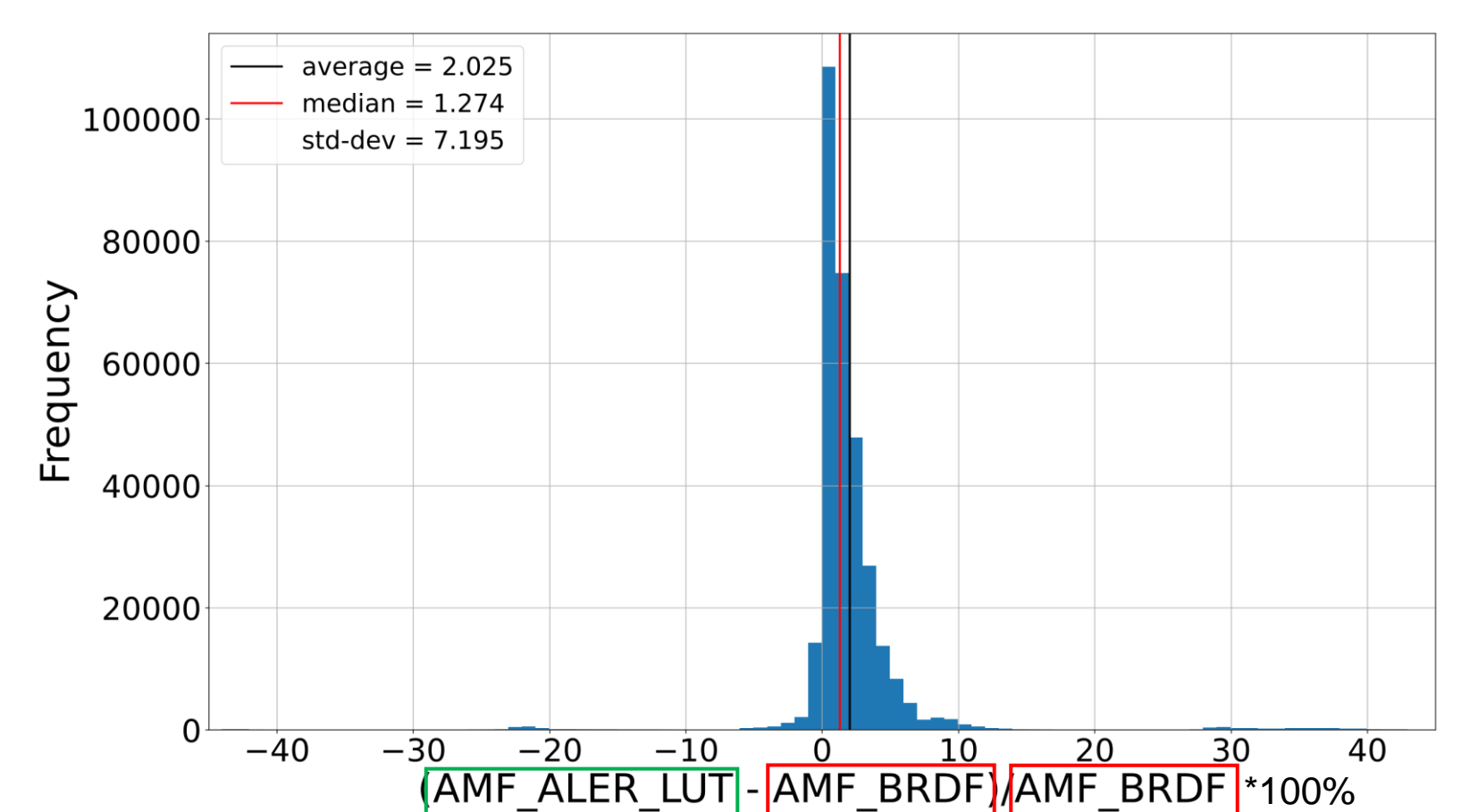
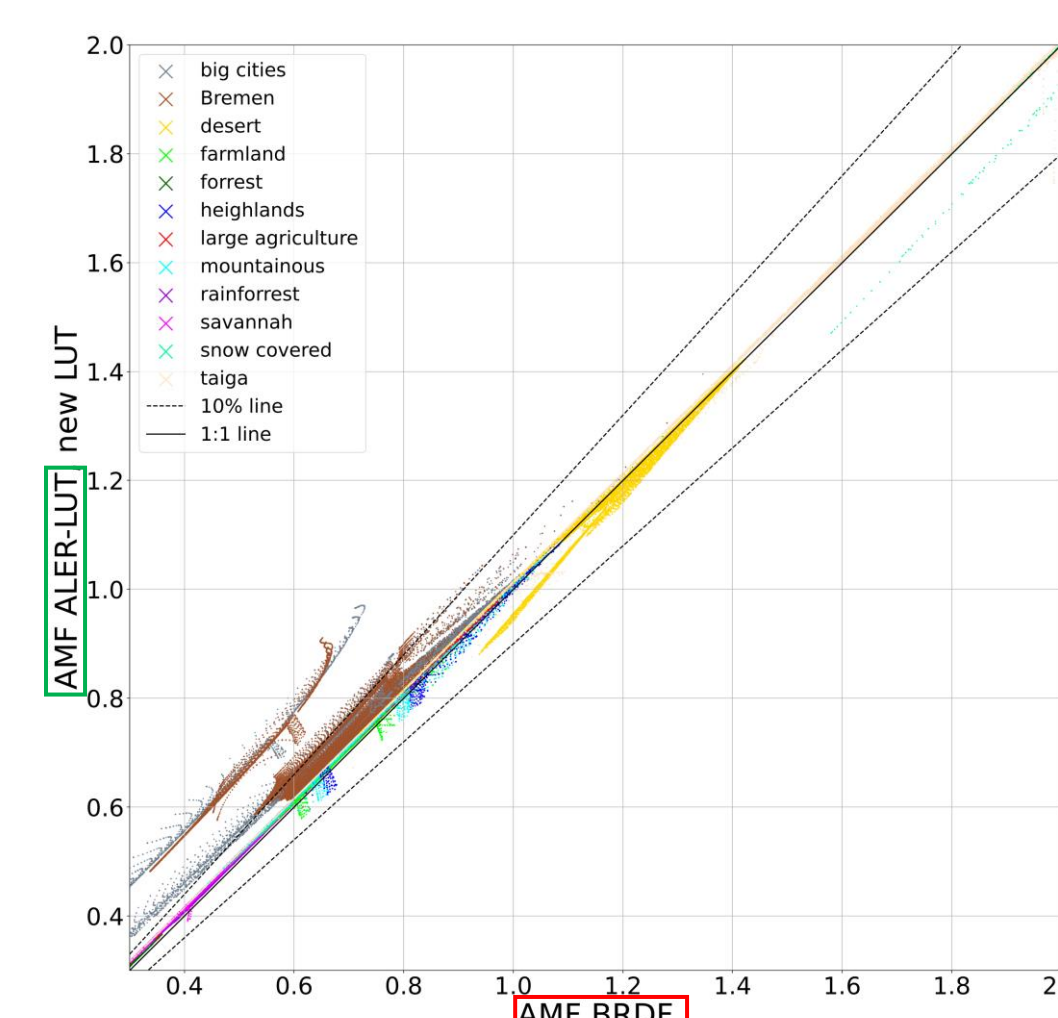


Quantifying the scatter shows spread up to ~40% and an offset of ~7.5%

- Testing single scattering (ss) instead of multiple scattering



- Only a slight hotspot is visible
- statistics is closer to what is hoped for



- Statistics is better, but problems at low AMF (not shown, reason not known)
- ss is not realistic in the atmosphere
- Shows that ms introduced problems

## 4. Conclusion

- BRDF has significant influence on the AMF
- Comparison of BRDF and method of Vasilkov showed problems:
  - Offset between AMF datasets
  - Ring structure around VZA ~4° - ~22°
  - Hotspot at retroreflection
- Better results obtained using single scattering, but problems occurred for some geometries

## References

Fuentes Andrade et al. (2023), A method for estimating localized CO<sub>2</sub> emissions from co-located satellite XCO<sub>2</sub> and NO<sub>2</sub> images, <https://doi.org/10.5194/egusphere-2023-2085> (Preprint)

Tilstra, L.G. et al. (2024), TROPOMI ATBD of the directionally dependent surface Lambertian-equivalent reflectivity