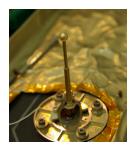
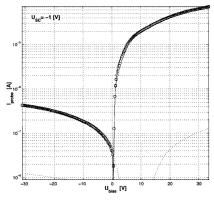
Status of the Swarm Langmuir Probes

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Bias sweep (scan) of the Cassini TiN LP probe

- Titanium has low mass density and is strong,
- commonly used for Langmuir probes (LP) in space,
- but must not oxidice at the surface to TiO₂,
- which is non-conducting.
- Therefore pure titanium is sublimed in a Nitrogen atmosphere producing a TiN surface.

- ➤ Swarm with LPs is at IABG Munich, (almost) ready for transport to Plesetsk in August 2013
- July 2013: "J-P Lebreton (PI LP of the Demeter satellite) observes an hysteresis effect, attributed to an RC time constant on TiN coated LPs"
- ... likely because of a non-conducting surface layer;
- possibly from aggressive atomic O in the Earth's thermosphere?
- X-ray Photoelectron and Raman Spectroscopy analysis of a (spare) Langmuir probe is done at ESTEC:

ABSTRACT

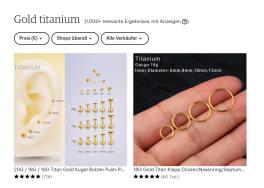
XPS and Raman spectroscopy analyses indicate that titanium nitride coating of the sample contains impurities/defects. High concentration of contaminants (more than 50 atomic%) is registered in the top layer (~5 mm). These contaminants include oxygen-, nitrogen- and fluoro-containing organic compounds, silicon oxides, silicon oxy-carbides, and small amounts calcium, magnesium (likely present in forms of their salts, e.g., in chlorides, carbonates, phosphates). Total 22.4±2.0 atomic% of titanium was measured by XPS in the surface layer of ~5 mm. Pure stoichlometric titanium nitride (TIN) was not registered by XPS in this surface layer. Titanium nitride in the top layer of ~5 mm (XPS data) is present in a partially oxidised state and in oxy-nitrides TIN,O, (63±3 % of the total titanium content). Titanium bonded to fluorine (9±2 % of the total titanium content) is also boserved in the surface layer. Titanium is also present in forms of its oxides (TiO, and TiO, 28±3 % of the total titanium content) and possibly in titanium oxy-carbides. Copper and zinc are also found and could be present in one of their alloys (e.g. foxas).

- ▶ alternative probe material: brass, electroplated with Au
- ▶ heavier than Ti! (test shaking etc cannot be done in time);
- electroplating gold on Ti? From https://www.sharrettsplating.com/base-materials/titanium

ISSUES WITH PLATING ON TITANIUM

Plating on titanium has long been considered an extremely difficult, if not impossible, process to master. The biggest issue is that titanium is a highly reactive metal. Specifically, titanium reacts with the oxygen that is produced by many plating processes

But: Jewelry



lt cannot be that difficult, go to the experts:



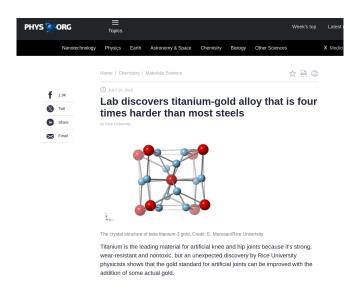
Kugel / IRF

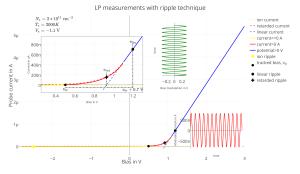
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Nb.	Ti (%)	d (um)	Ni (%)	d (um)	Au (%)			
1	100,0	0,0	100,0	3,08	100,0			
2	100,0	0,0	100,0	3,13	100,0			
3	100,0	0,0	100,0	2,36	100,0			
4	100,0	0,0	100,0	3,11	100,0			
5	100,0	0,0	100,0	3,25	100,0			
6	100,0	0,0	100,0	3,10	100,0			
7	100,0	0,0	100,0	2,85	100,0			
8	100,0	0,0	100,0	2,78	100,0			
Stat	istik							
		d (µm)	Ti (%) d (µr	n) Ni (9	6)	d (µm)	Au (%)
Mittelwert			100,	0 0	,0 100	0,0	2,96	100,0
StdAbw.			0,0	0 0	,0 0	,0	0,28	0,0
Minimum			100	0 0	0 100	20	2 36	100.0

- ► Lennart Åhlen (†) and I went to IABG, Munich, to replace the +Y TiN probes with gold-plated ones,
- on August 29, 2013.

July 20, 2016





Probe	Gain ^a	Surface	Position
1	high -2018 low 2019-	TiN	left
2	low -2018 high 2019-	Au	right

- ► There is no firm indication that the efforts made any difference,
- both probes, TiN and gold-plated Ti, work fine;
- perhaps the "ripple" mode is less sensitive to thin surface contamination?
- ▶ there are no clear sign of LP aging after 10 years in orbit :-)

Future?

So far unique features of the Swarm LPs:

- "ripple" mode instead of bias sweeps/scans;
- the plasma density is derived from the ion admittance (at negative bias),
- which is independent of the spacecraft potential;
- the operations do not need any "maintenance";
- we suggest to equip a constellation (\sim StarLink) with simple LPs,
- ▶ 100-1000 satellites, "big data" for ionosphere research.