

Swarm FAST data processing chain

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Abstract:

After 10 years in Space, Swarm ESA's Earth Explorer mission is still in excellent shape and contribute to a wide range of scientific studies, from the core of our planet, via the mantle and the lithosphere, to the ionosphere and interactions with Solar wind. Its highly accurate observations of electromagnetic and atmospheric parameters of the near-Earth space environment, and the peculiar mission constellation design, make Swarm eligible for developing novel Space Weather products and applications. In April 2023 a "FAST" processing chain has been transferred to operations, providing Swarm L1B products with a minimum delay respect to the acquisition. These Fast data products add significant value in monitoring present Space Weather phenomena and help modelling and nowcasting the evolution of several geomagnetic and ionospheric events. This work presents the set-up of the Swarm "FAST" data processing chain, current status and plans for future improvements and applications.

ESA

SMOS

Swarm L1B FAST chain setup

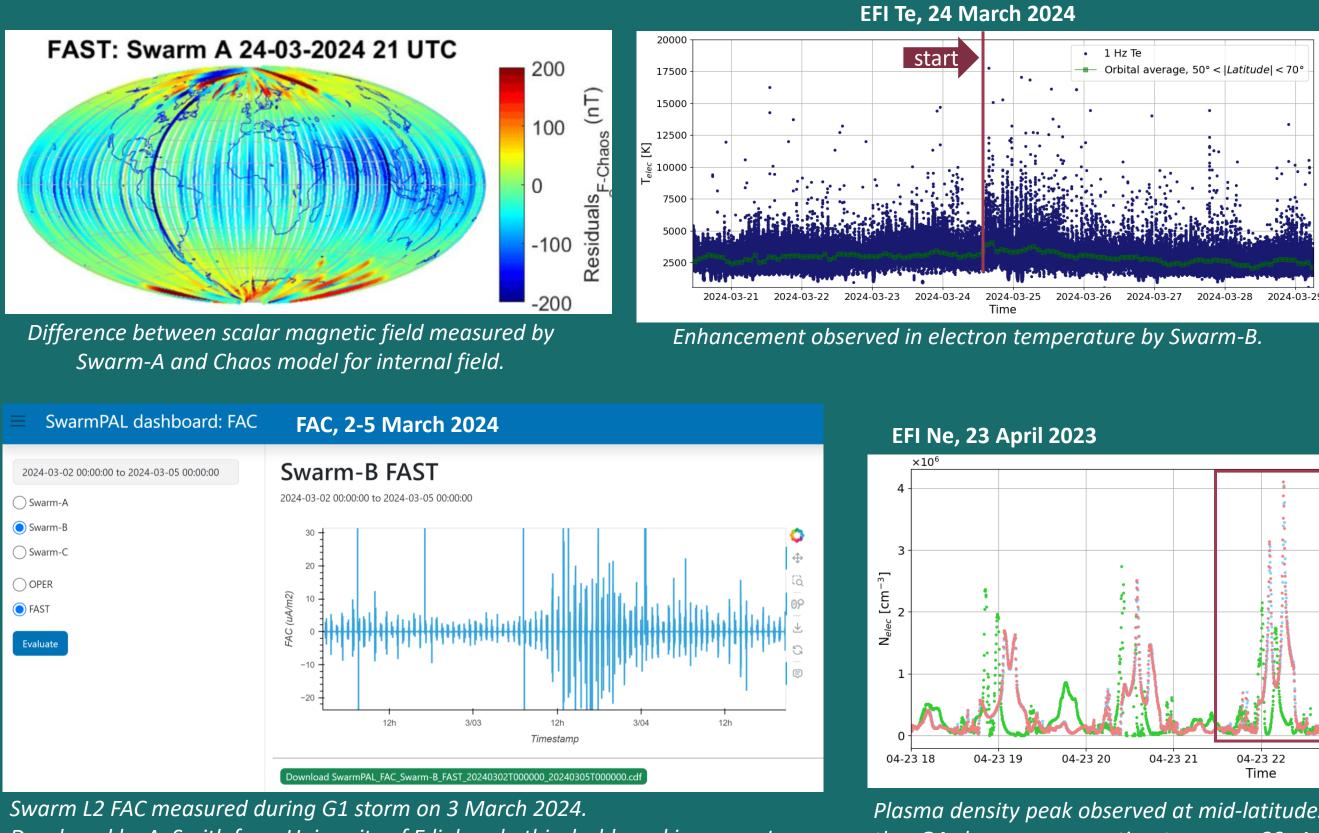
Implementation plan:				
PHASE I - Analysis	PHASE II - Feasibility	PHASE II	l - Demo	PHASE IV - Routine
 Retrieve alternative Auxiliary Data 3-days L1B TDS, QC & validation Characterise possible errors Identify potential users 	 Integrate modified IPF into a parallel chain 1 year L1B TDS, QC & validation Provide TDS to selected users and collect feedback 	 Demonstra Datasets: T S/Cs disseminar selected u Collect feet 	TDS for all te to sers	 Systematic Fast L1b Data production for all S/Cs Revised dump approach Extend to selected L2 products
Activities timeline:				
study pro	y 2022 Apr. 2023 June 20 totype TTO FAST Pass stra ployed v101 optimiza 2023 2023	itegy data	v. 2023 a access all users 2023	L2 FAST chain 2024
2021 20	March May	Aug		March
Sept. 2022May 2023Oct. 2023March 2024TDS distributeddata access forTTO FASTDownlink approachto Cal ValMission usersv102optimization				
FAST Data Processing	g approach:			
FAST Processor	Changes:		Main differ	ences introduced in
version:	The only change for FAST		orbit data determination:	
Aligned with most	chain with respect to OF			R accuracy is 7 cm;
updated version of L1B data	alternative input AUX files used for ORBATT processing.		MOD FAST accuracy is about 16 cm in terms of 3D RMS	
algorithm	\rightarrow No changes for the c	J		gradation, with the

Swarm FAST data for Space Weather

Space Weather events observed with Swarm FAST data:

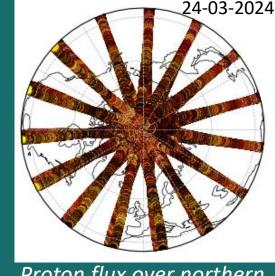
With a minimum delay with respect to the sensing time, Swarm FAST data are suitable for Space Weather applications. Several events of Solar Cycle #25 have been already observed and analysed with Swarm FAST data, allowing the scientific community to promptly monitor Magnetic and Electric field variations.

As a representative example, the strongest event of Cycle #25 so far (a severe G4 class storm) and its impact on external magnetic field and plasma environment has been analysed with Swarm few hours after the interaction:

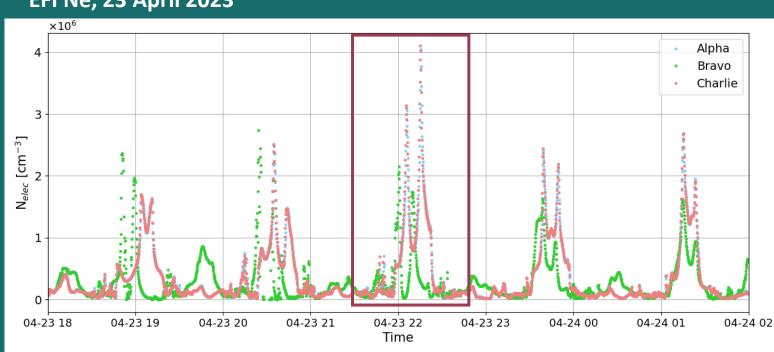


Developed by A. Smith from University of Edinburgh, this dashboard incorporates a

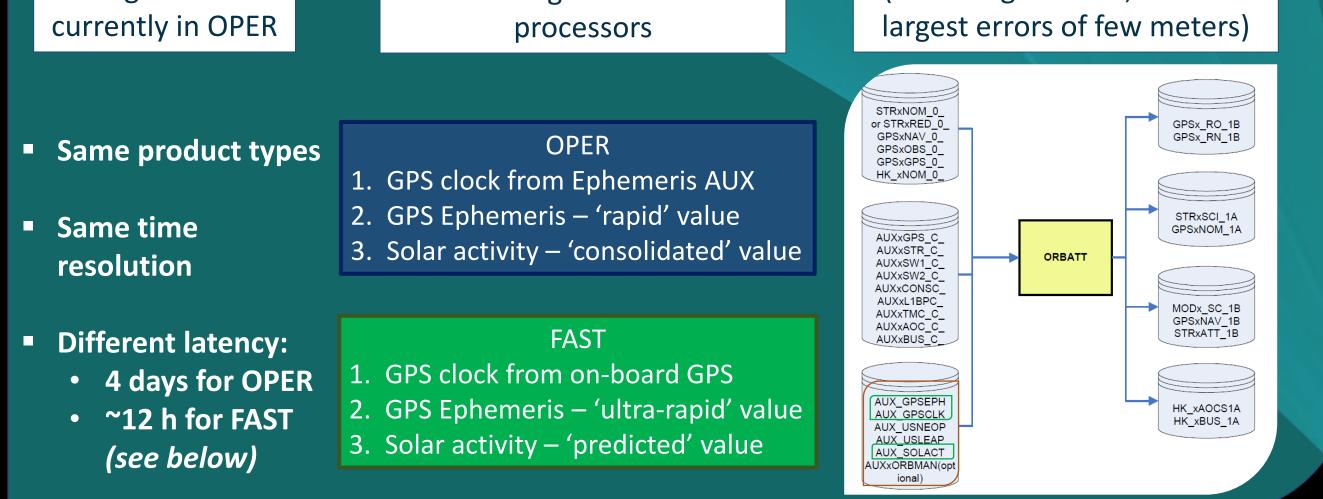
5 10 15 20 25 Swarm Bravo Proton flux [cm⁻² s⁻¹]



Proton flux over northern hemisphere, computed by the prototype processor under development for the next baseline



Plasma density peak observed at mid-latitudes by Swarm in correspondence of the G4-class geomagnetic storm on 23 April 2023. This event has been



Latency and Coverage improvements



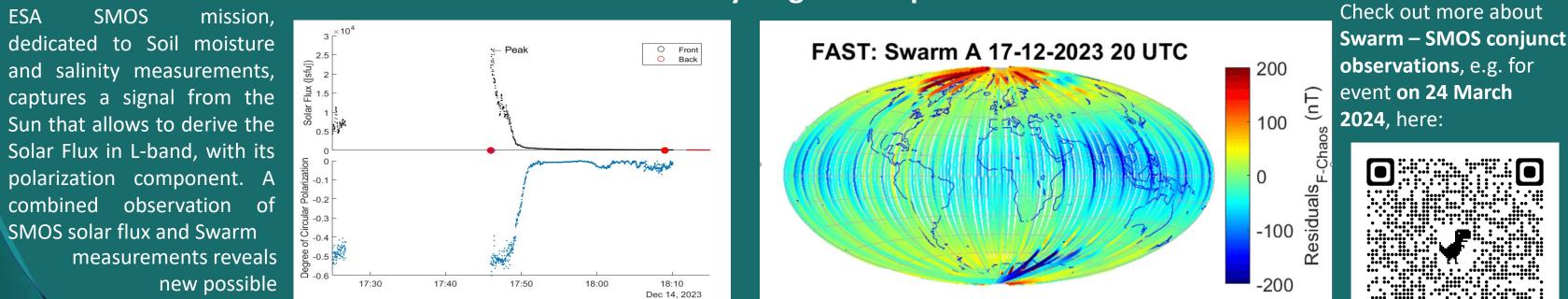
Latency after downlink and distribution is mainly archival to and processing time. ORBATT, ACCELE and PLASMA data are available after ~1h 20m. MAGNET processor is the most time-consuming, taking ~30min more in average. Processor optimization, to

further improve the latency, can be applied in the future.

Users' community recommended to improve pass strategy to provide a more uniform pass distribution and a more efficient utilisation of the FAST data. The optimization has been defined, reflecting requirements and operational constraints, and applied: current strategy foresees 2 passes per day every ~12h (with a standard deviation of ~1:30-2h).

prototype FAST processing allowing a quick look of the FAC for all Swarm spacecraft characterized by auroras observed at lower latitudes (e.g. in central Europe). from both the operational and FAST production.



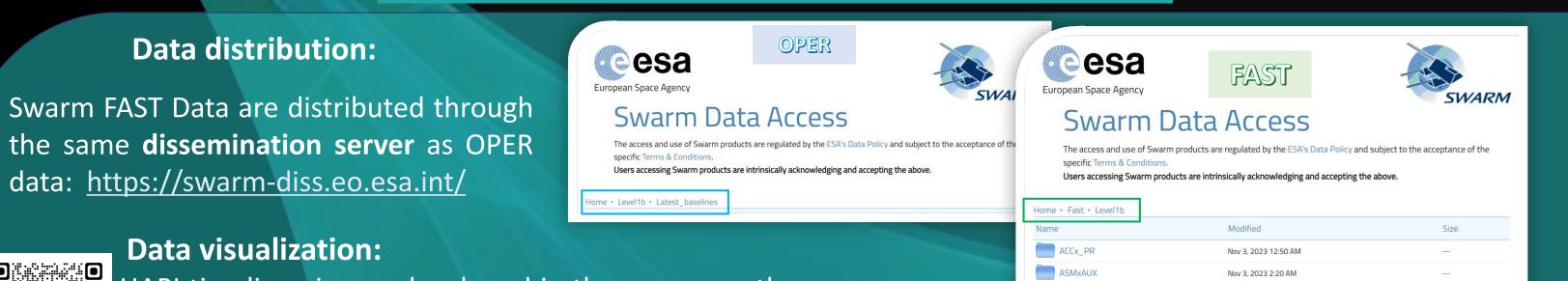


applications! Solar Radio Burst measured by SMOS MIRAS sensor in terms of derived solar flux and Degree of Circular Polarization component on 14-12-2023 (left); variations of external magnetic field measured by Swarm few days later (right).

What's coming next?

- FAST processing data quality is continuously under monitoring and frequently used by Swarm community. Feedbacks and recommendations are collected and applied to the future improvements plan. Short-term upgrades includes:
- Align L1B FAST processor with the upcoming L1B OPER processor version 3.25 (see poster #9, N. Comparetti et al.)
- Implement L2 FAST production chain

FAST data distribution and visualization



EFIX_LP

EFIxL

Nov 3, 2023 12:50 AM

Nov 3, 2023 12:50 AM

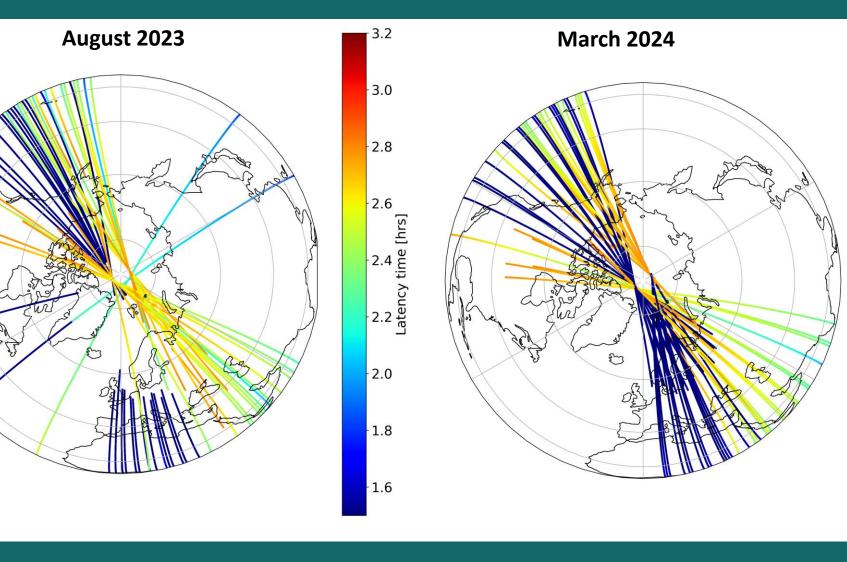
Config + Add plot

B_C_res_Model (Alpha B_C_res_Model (Bravo)

B_C_res_Model (Charlie)

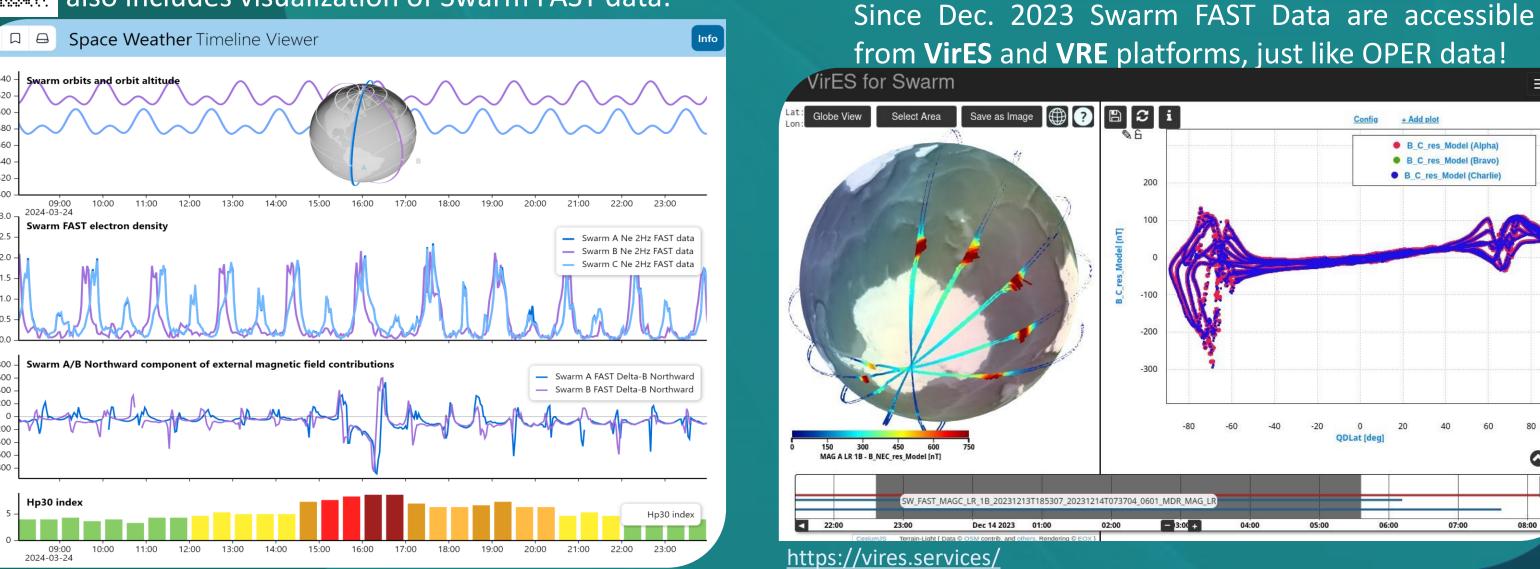
A further improvement involving the downlink approach, set up in March 2024, guarantees a better coverage and reduces latency over northern Europe.

Latency here is defined as difference between data acquisition time and data generation + distribution time.



Impact of the change in the downlink approach for Swarm Alpha during a 15 days period.

HAPI timeline viewer, developed in the space weather group at KNMI as part of the Swarm SWITCH project, also includes visualization of Swarm FAST data:



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