



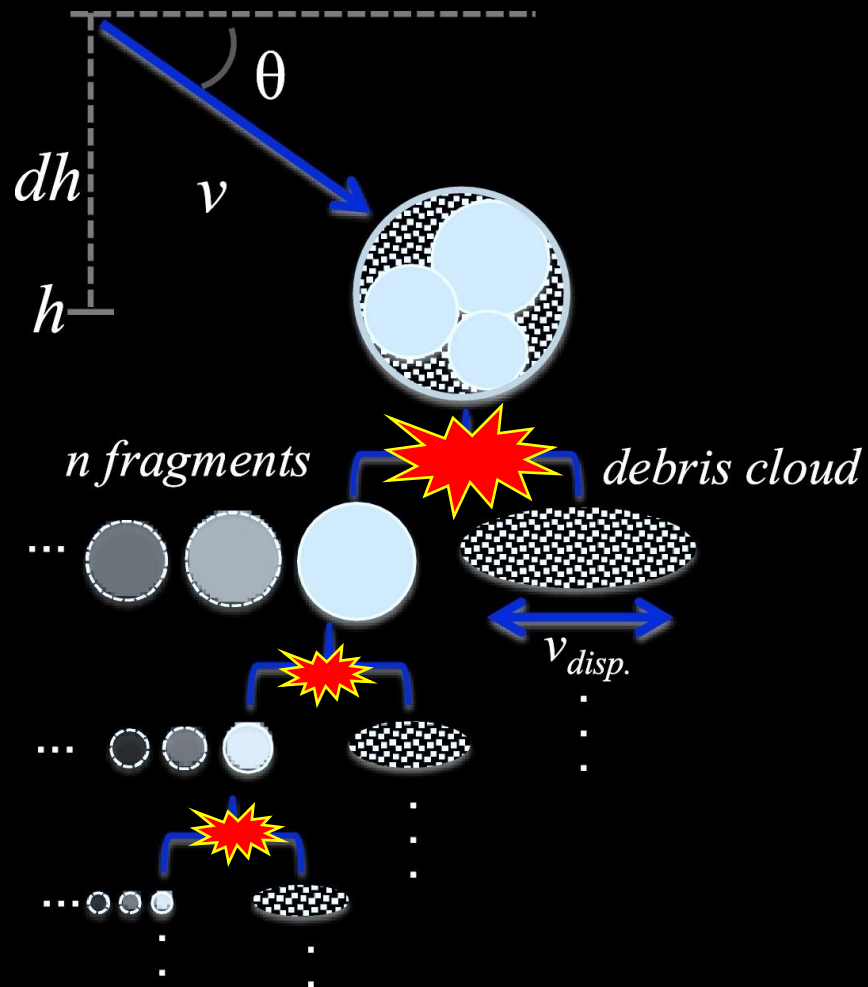
Accurate characterization of metre-sized impactors
through casual bolide observations:

Novo Mesto superbolide as evidence for a new class of high-risk objects

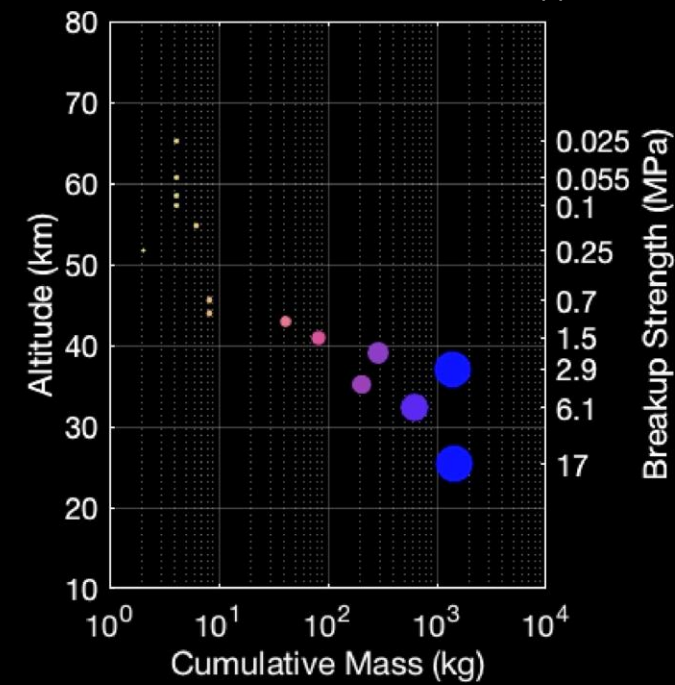
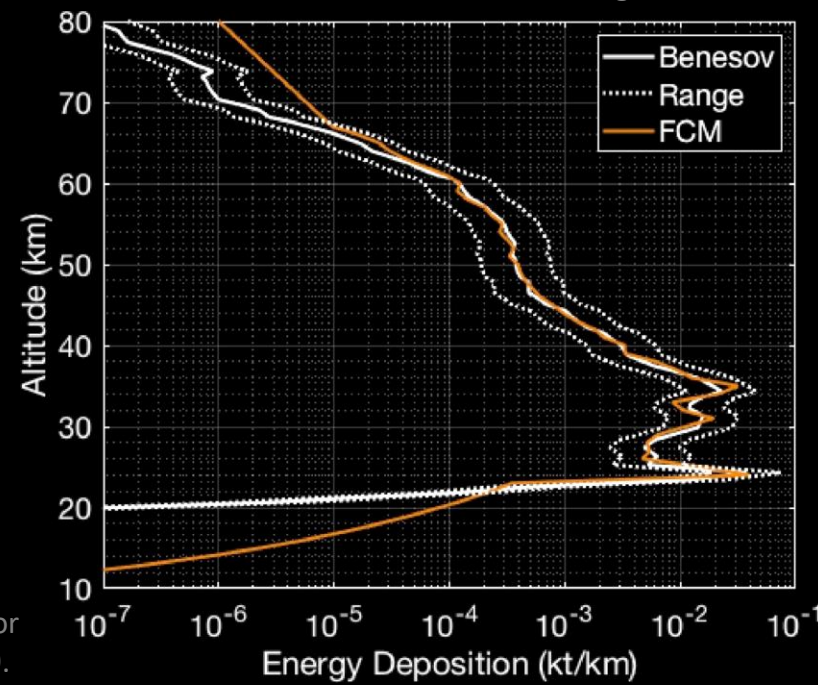
Denis Vida, Damir Šegon, Marko Šegon, Jure Atanackov, Bojan Ambrožič, Mikael Granvik, Luke McFadden, Ludovic Ferrière, Peter Brown, Barbara Malečić, Maja Telišman Prtenjak, Javor Kac, Gregor Kladnik, Mladen Živčić, Aleksandar Merlak, Ivica Skokić, Lovro Pavletić, Gojko Vinčić, Ivica Ciković, Zsolt Perkó, Martino Ilari, Mirjana Malarić, Igor Macuka

Motivation: Modelling airburst ground effects

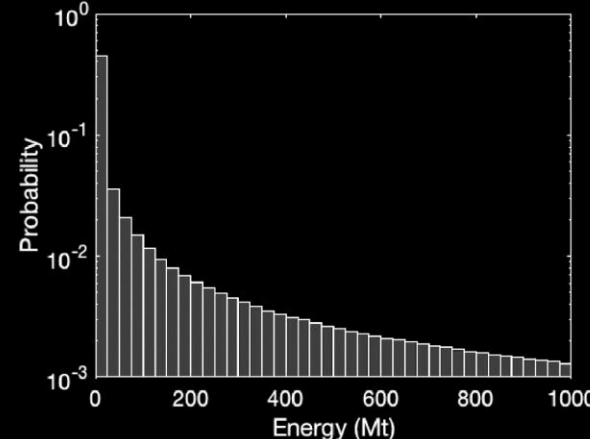
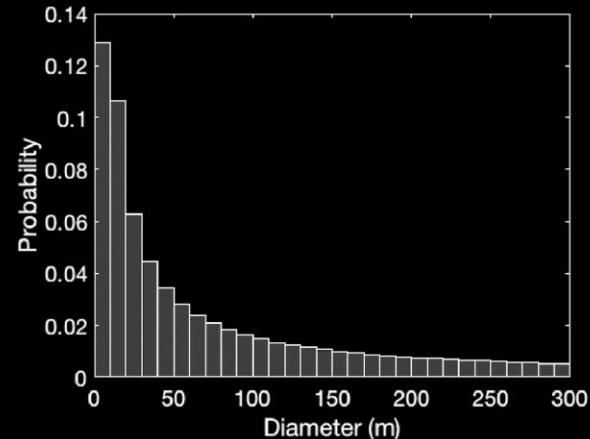
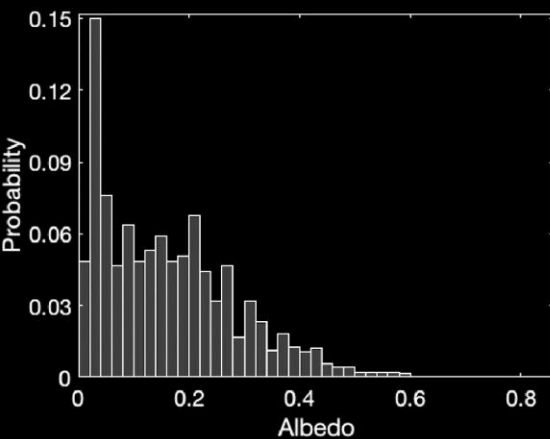
- Assumed strength distribution and fragmentation behavior (Weibull)
- Blast overpressure the main damage source for $\lesssim 100$ m objects



Wheeler, L.F., Mathias, D.L., Stokan, E. and Brown, P.G., 2018. Atmospheric energy deposition modeling and inference for varied meteoroid structures. *Icarus*, 315, pp.79-91.

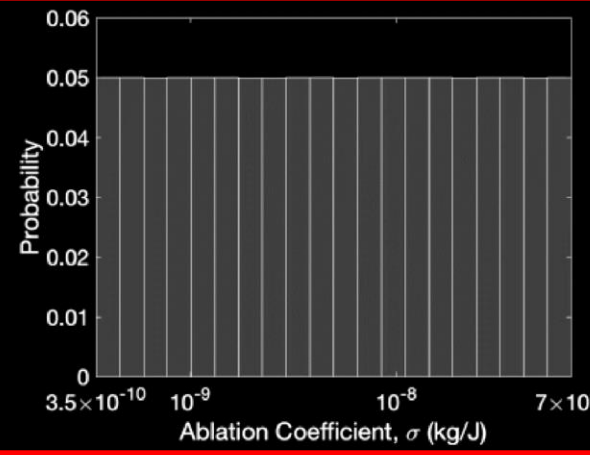
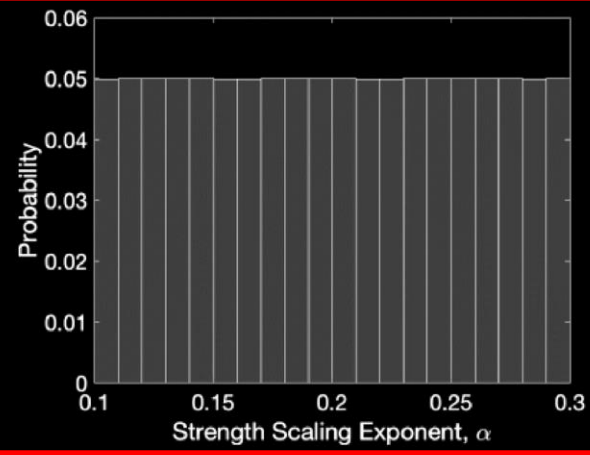
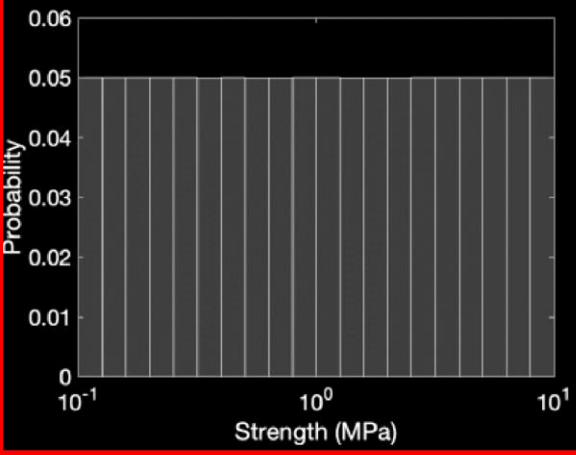
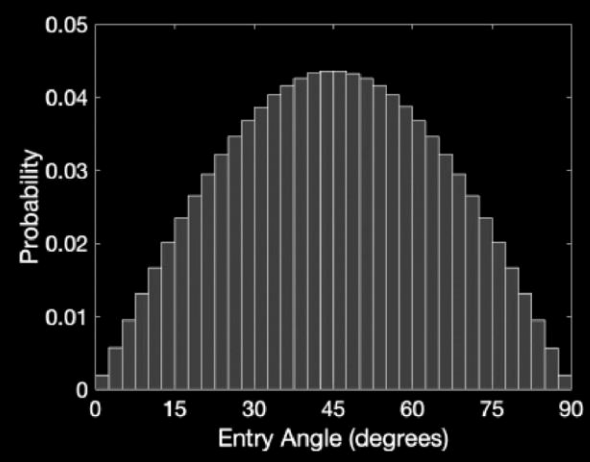
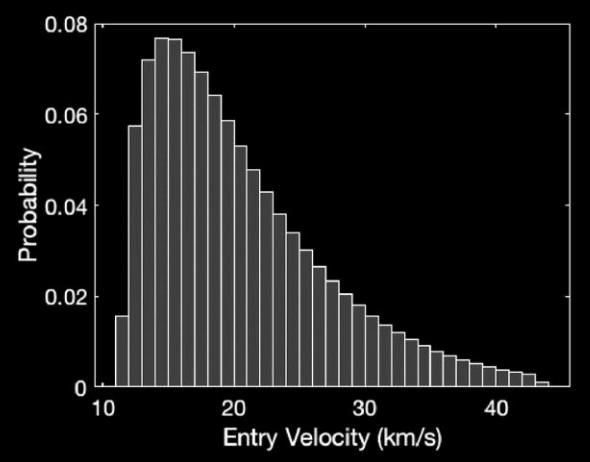
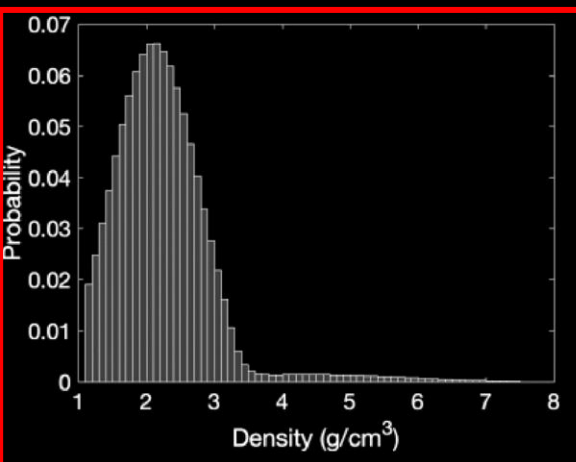


Wheeler, L.F., Register, P.J. and Mathias, D.L., 2017. A fragment-cloud model for asteroid breakup and atmospheric energy deposition. *Icarus*, 295, pp.149-169.



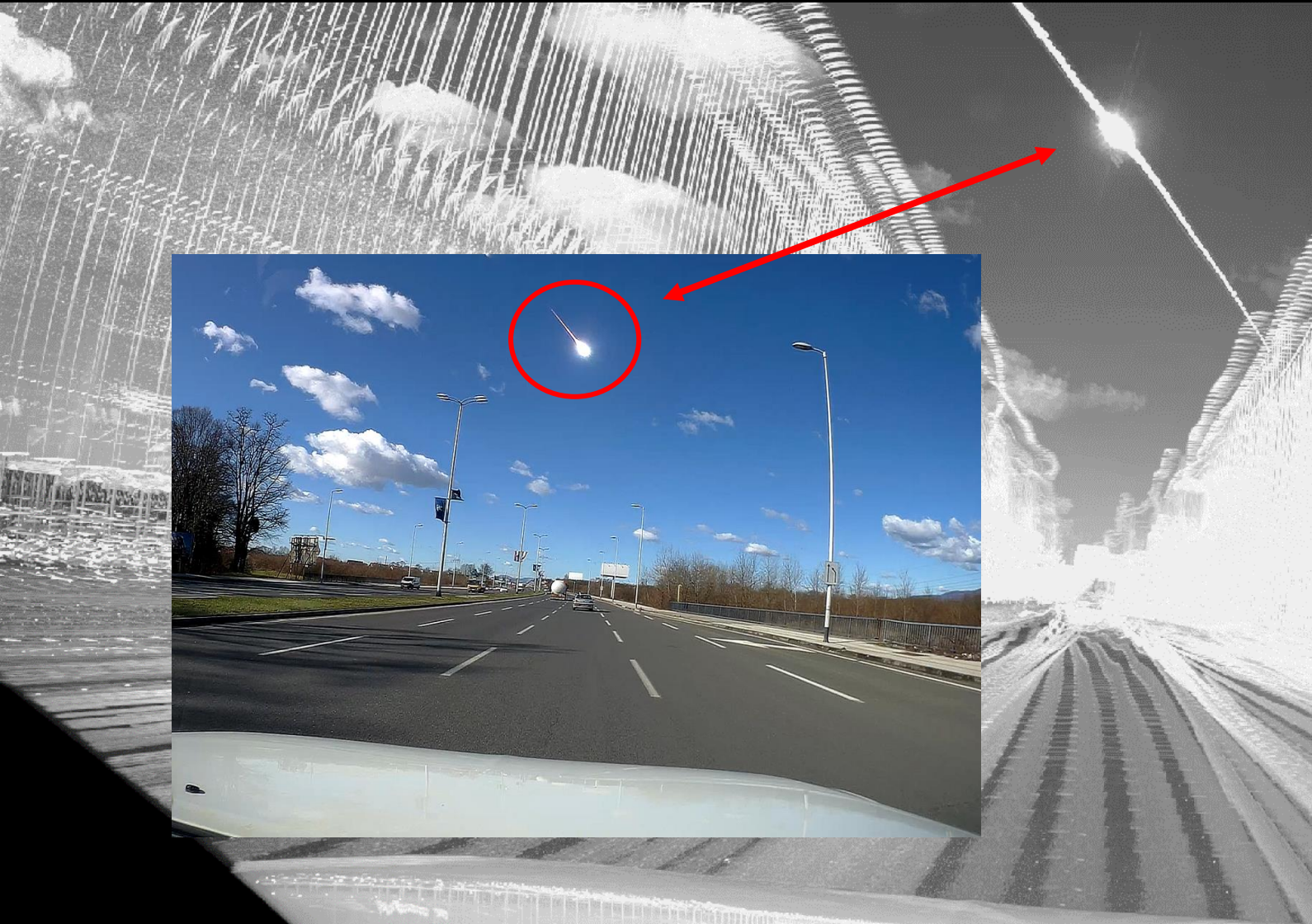
Need observational constraints:

- Dependence of these on orbital class
- Object bulk density
- Strength distribution
- Ablation characteristics
- Fragment mass distribution (small dust vs large fragments)



Novo Mesto (Slovenia) superbolide

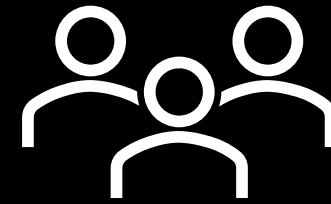
- February 28, 2020, 10:30 AM local time



Up to 150 km away:

“loud explosion”

“ground shaking”



“rattling windows”



Three L5 chondrite meteorites (450, 200, 50g)
recovered in the following days

Station: HRIVOA
2020-02-28 09:30:30.666667

avepixel
Ref Az = 262.181°
Ref Alt = 2.427°
Rot horiz = 5.307°
Rot eq = 311.074°
Pix scale = -2.616/px
Lim mag = 4.0
Increment = 1.00
Img Gamma = 1.00
Camera Gamma = 1.00
Refraction corr = False
Distortion type = poly3+radial
Extinction Scale = 1.00

RA centre = 15h 35m 08.24s
Dec centre = -3.849°

Vehicle location inverted to within ~2 cm
Some travelling over 70 km/h!

Astrometric accuracy ~3 arcmin

Keys:

F1 - Hide/show this text
Left/Right - Previous/next image
CTRL + Left/Right - +/- 10 images
A/D - Azimuth
S/W - Altitude
Q/E - Position angle
Up/Down - Scale
T - Toggle refraction correction
1/2 - X offset
3/4 - Y offset
5/6 - X 1st dist. coeff.
7/8 - Y 1st dist. coeff.
9/0 - extinction scale
CTRL + 1 - poly3+radial distortion
CTRL + 2 - poly3+radial3 distortion
CTRL + 3 - radial3 distortion
CTRL + 4 - radial5 distortion
CTRL + 5 - radial7 distortion
CTRL + 6 - radial9 distortion

CTRL + R - Pick stars

Scroll - zoom in/out
R/F - Lim mag
+/- - Increment adjust

M - Toggle maxpixel/avepixel
H - Hide/show catalog stars
C - Hide/show detected stars
CTRL + I - Show/hide distortion
U/J - Img Gamma
I - Invert colors
V - FOV centre

CTRL + A - Auto levels
CTRL + D - Load dark
CTRL + F - Load flat
CTRL + X - astrometry.net img upload
CTRL + SHIFT + X - astrometry.net XY
only
SHIFT + Z - Show zoomed window
CTRL + N - New platepar
CTRL + S - Save platepar state

STAR PICKING MODE
LEFT CLICK - Centroid star
CTRL + LEFT CLICK - Manual star position
RIGHT CLICK - Remove pair
CTRL + SCROLL - Aperture radius adjust
CTRL + Z - Fit stars
CTRL + SHIFT + Z - Fit with initial distortion params set to 0
L - Astrometry fit details
P - Photometry fit

Station: HRIVOA
 2020-02-28 09:30:30.666667

avepixel
 Ref Az = 262.141°
 Ref Alt = 2.477°
 Rot horiz = 5.297°
 Rot eq = 311.089°
 Pix scale = -2.616/px
 Lim mag = 4.0
 Increment = 1.00
 Img Gamma = 1.00
 Camera Gamma = 1.00
 Refraction corr. = False
 Distortion type = poly3+radial
 Extinction Scale = 1.00

RA centre = 15h 35m 23.49s
 Dec centre = -3.941°

Vehicle location inverted to within ~2 cm
 Some travelling over 70 km/h!

Astrometric accuracy ~3 arcmin

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 CTRL + 4 - radial5 distortion
 CTRL + 5 - radial7 distortion
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 P - Photometry fit

Station: HRIVOA
 2020-02-28 09:30:30.666667

avepixel
 Ref Az = 262.131°
 Ref Alt = 2.457°
 Rot horiz = 5.127°
 Rot eq = 311.260°
 Pix scale = -2.616/px
 Lim mag = 4.0
 Increment = 1.00
 Img Gamma = 1.00
 Camera Gamma = 1.00
 Refraction corr = False
 Distortion type = poly3+radial
 Extinction Scale = 1.00

RA centre = 15h 35m 22.03s
 Dec centre = -3.961°

Vehicle location inverted to within ~2 cm
 Some travelling over 70 km/h!

Astrometric accuracy ~3 arcmin

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 CTRL + 4 - radial5 distortion
 CTRL + 5 - radial7 distortion
 CTRL + 6 - radial9 distortion

CTRL + R - Pick stars

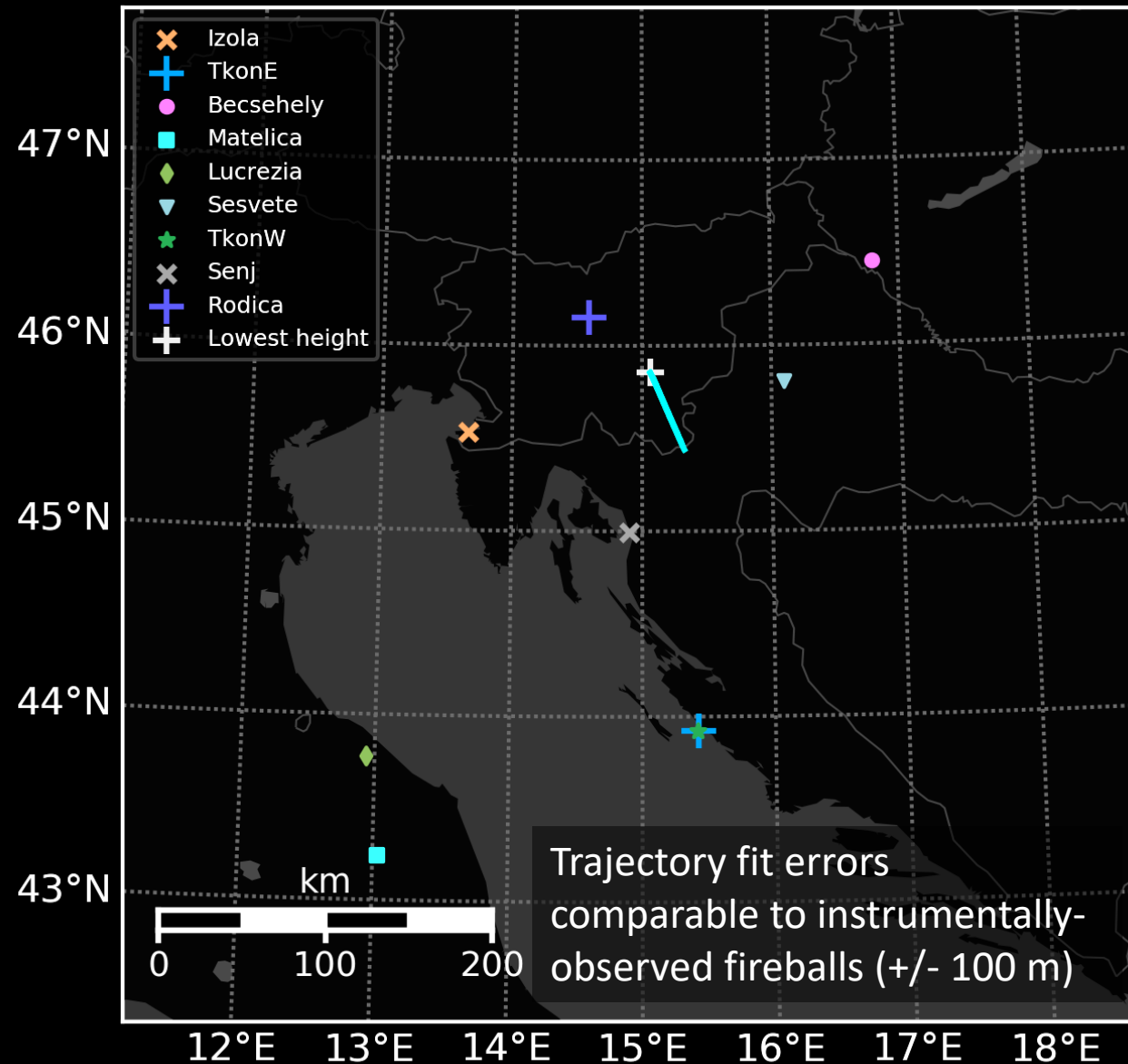
Scroll - zoom in/out
 R/F - Lim mag
 +/- - Increment adjust

M - Toggle maxpixel/avepixel
 H - Hide/show catalog stars
 C - Hide/show detected stars
 CTRL + I - Show/hide distortion
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 CTRL + F - Load flat
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 CTRL + SCROLL - Aperture radius adjust
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 CTRL + SHIFT + Z - Fit with initial distortion params set to 0
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Reconstructed trajectory and orbit



Orbit:

$$q = 0.5682 \pm 0.0005 \text{ AU}$$

$$e = 0.6095 \pm 0.0007$$

$$i = 8.777 \pm 0.021^\circ$$

$$\omega = 82.76 \pm 0.11^\circ$$

$$\Omega = 338.993019$$

Source region: inner-main belt

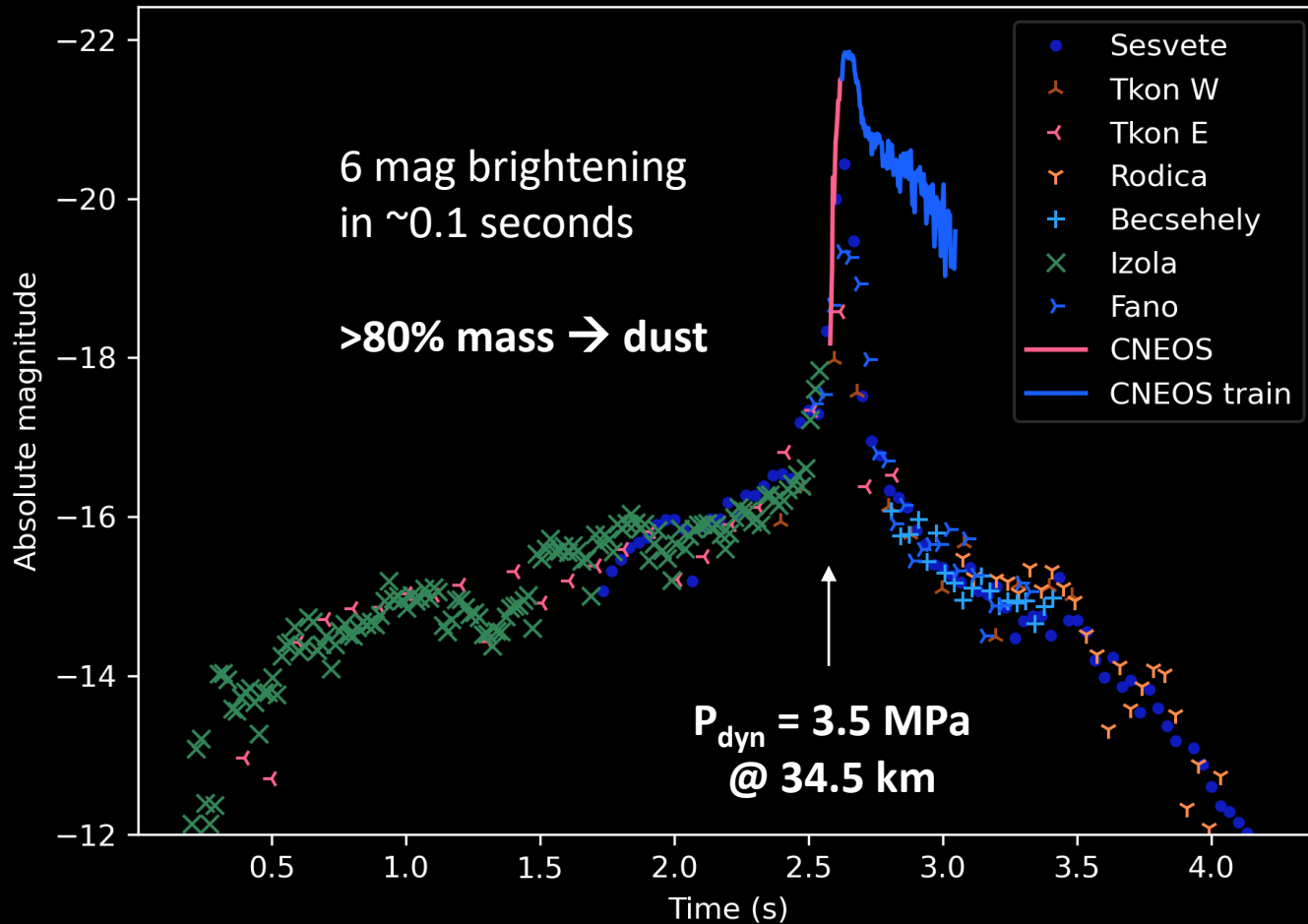
(definitely not from the Gefion family)

u_6 resonance	75–89%
3:1 resonance	4–16%
Hungaria	2–12%
5:2 resonance	1–4%

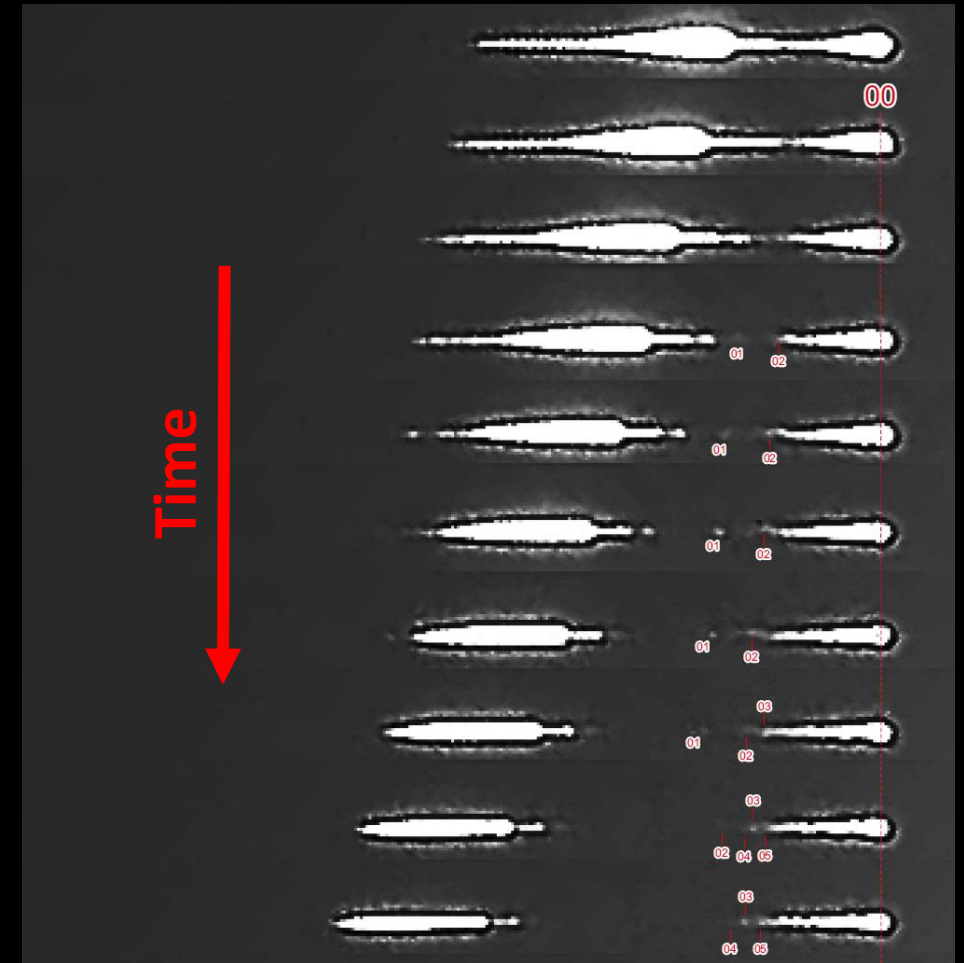
Energy deposition for Novo Mesto

- USG/CNEOS data includes the luminous train (0.34 kt)

Light curve after 2020/02/28 09:30:32.0 UTC



Luminous dust >> meteoroid



Fragmentation model fit

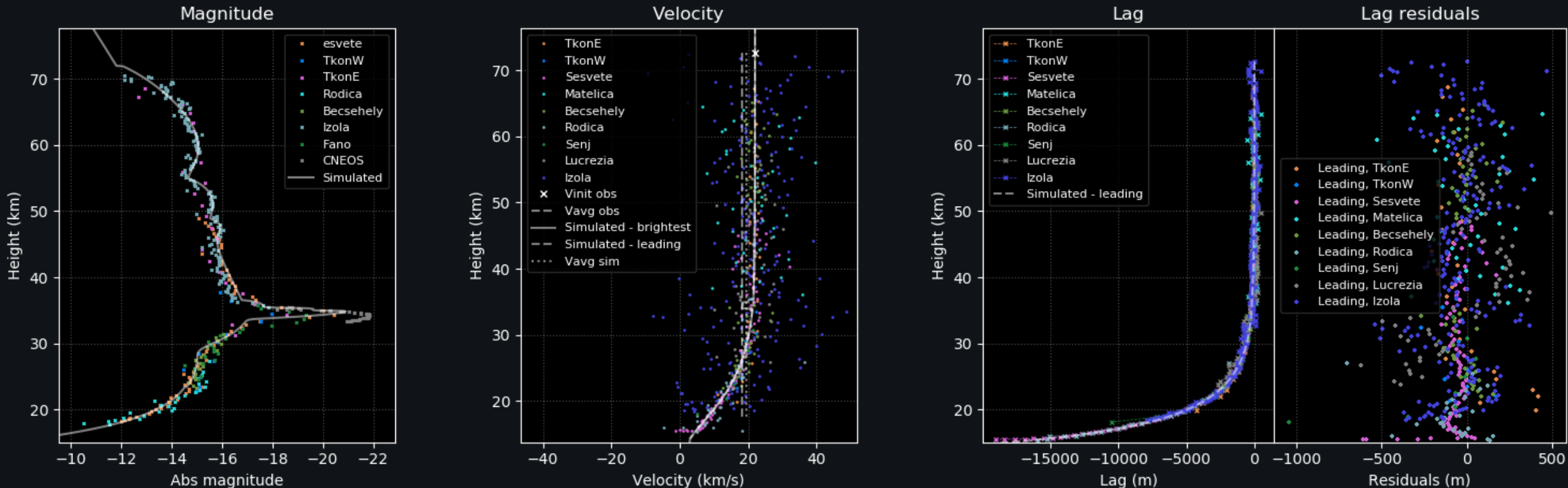
- Dynamics → accurate inversion - no deceleration until ~35 km

$m_0 \approx 3000$ kg, ~1 m diameter

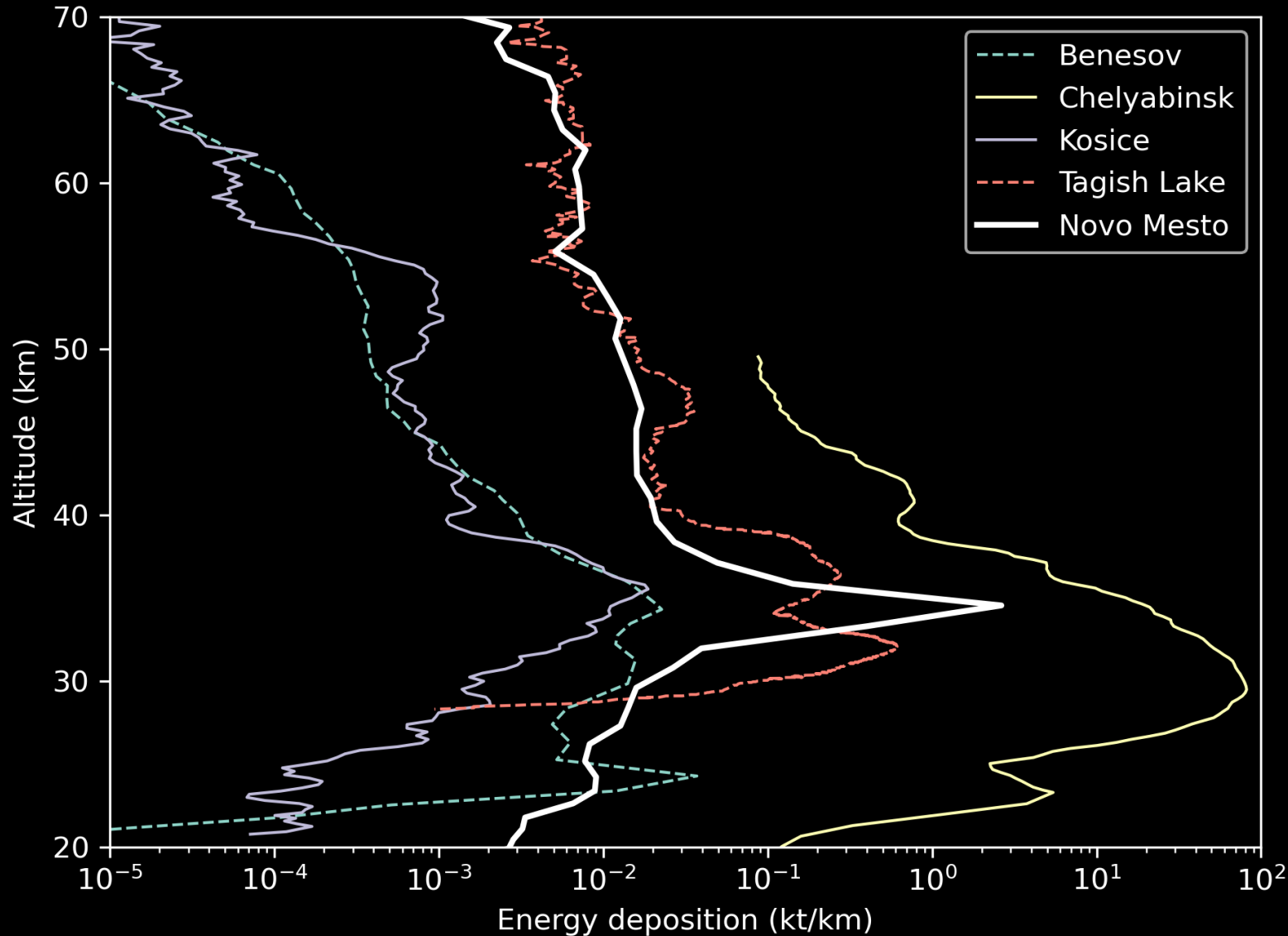
Bulk density = 3400 kg/m³

Main fragmentation: ~3.5 MPa (2500 kg into mm-sized dust)

Minor fragmentations: 0.2 – 0.6 MPa (20 – 70 kg fragments)



New object class: “Bunker buster” fireball



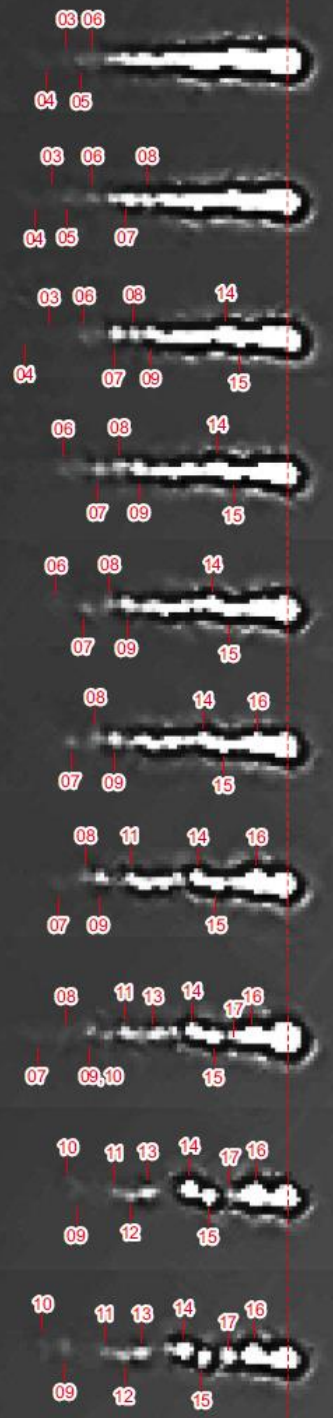
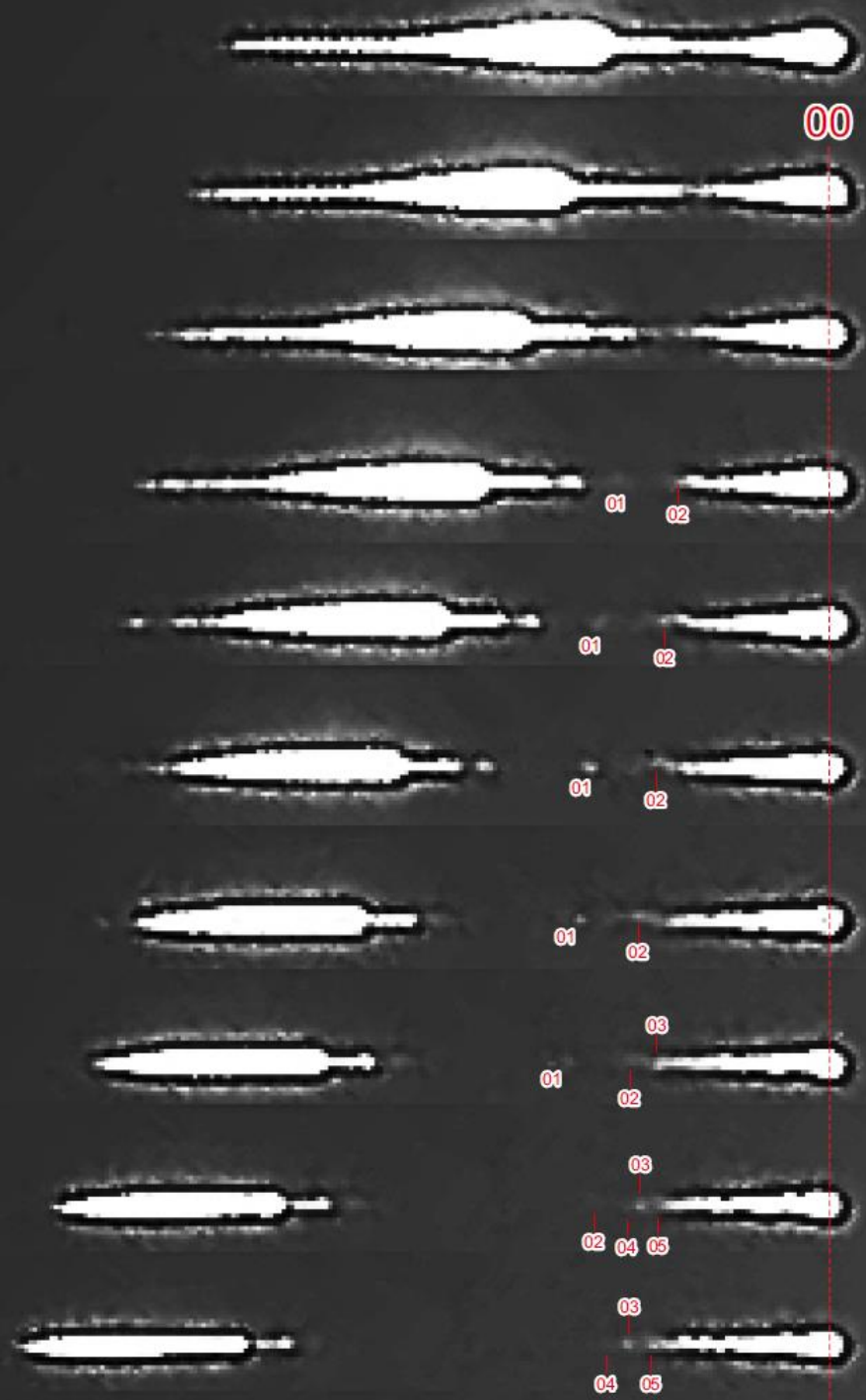
Peak overpressure
outsized for object size

Catastrophic structural
failure at single point at a
high dynamic pressure

Best analogy:
Weak concrete made
only of sand and small
gravel

Takeaways

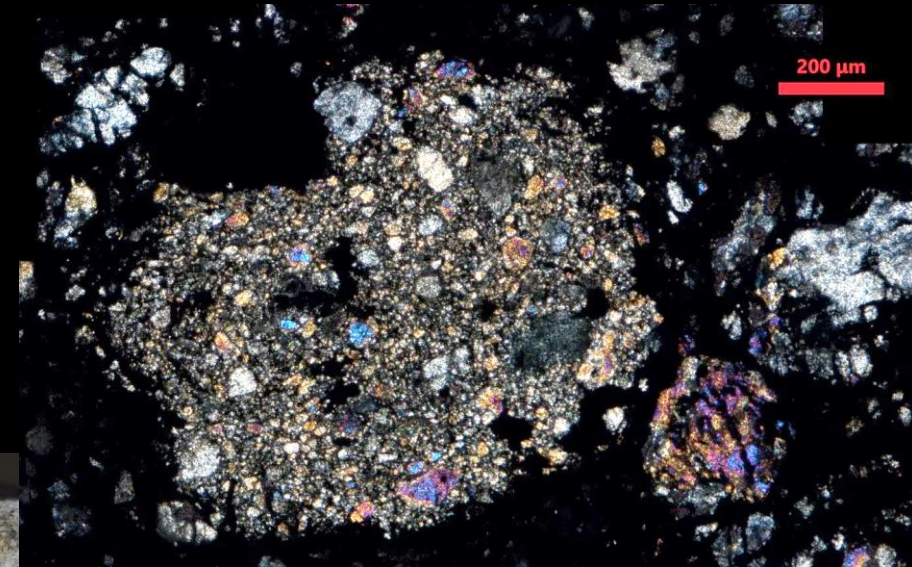
- Currently unknown: Structural composition of impactors
- Developed: Open-source method for calibrating casual fireball videos and images
- Accurate structural model = High-resolution video + dynamics
- Novo Mesto superbolide - a new “bunker buster” class
 - Deposits ~90% of energy in a single deep disruption

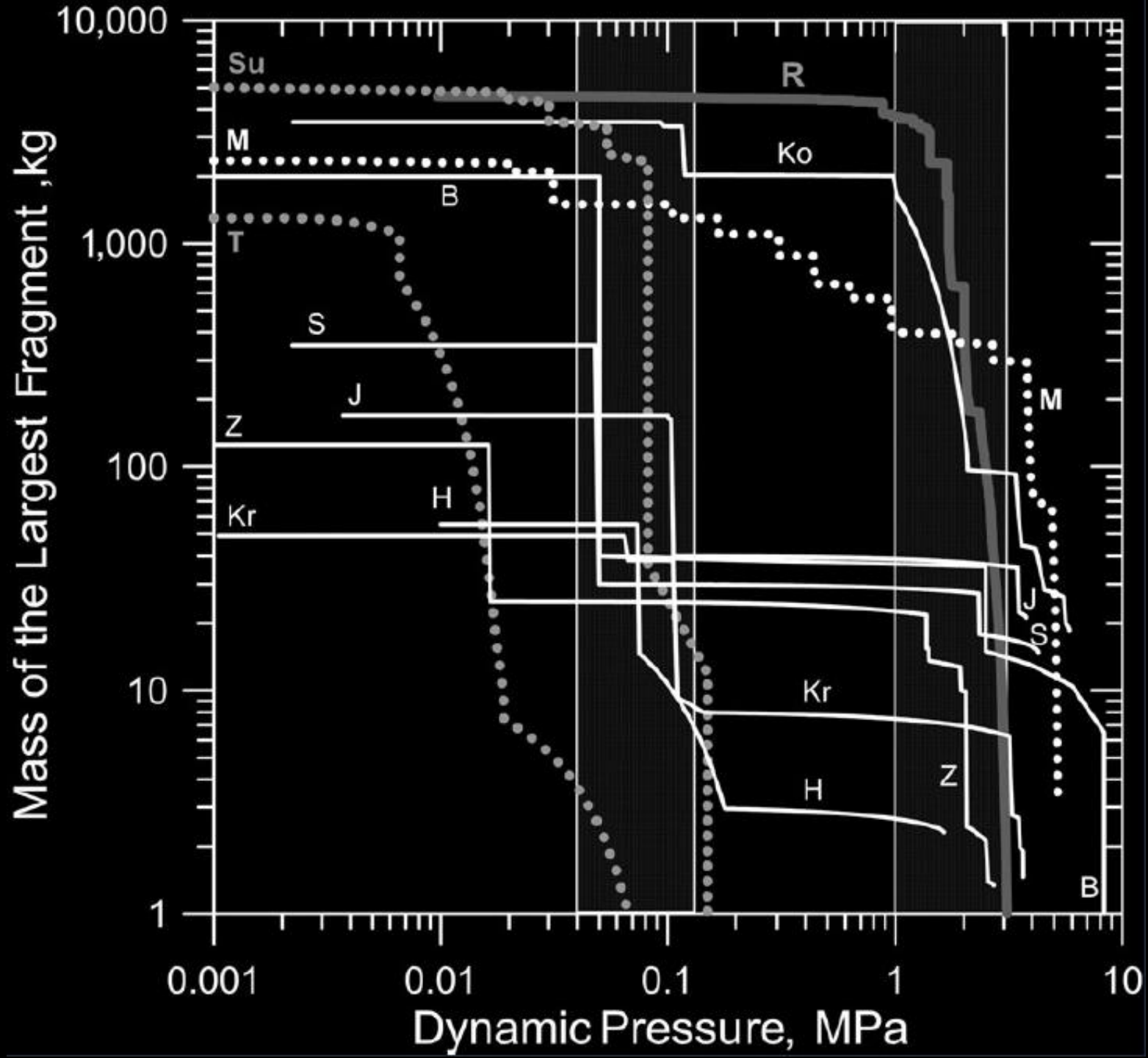


Backup slides

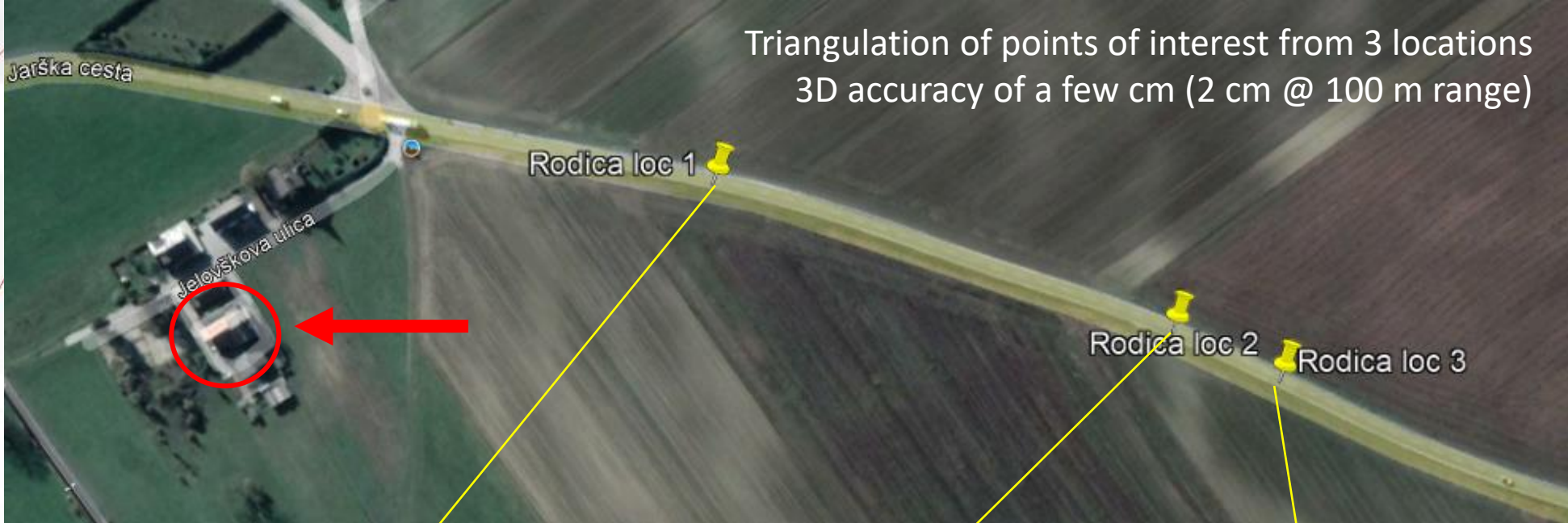
Meteorites

- Recovered by locals under the fireball's path within 2 weeks
- L5 chondrite, S5 shock stage, W0/1 weathering grade
- Brecciated and composed of several types of clasts in a fine-grained matrix
- Most chondrules are fractured, displaced





Borovička, J., Macke, R. J., Campbell-Brown, M. D.,
 Levasseur-Regourd, A. C., Rietmeijer, F. J., & Kohout, T.
 (2019). Physical and chemical properties of meteoroids.
Meteoroids: Sources of Meteors on Earth and Beyond, 37.



Triangulation of points of interest from 3 locations
3D accuracy of a few cm (2 cm @ 100 m range)

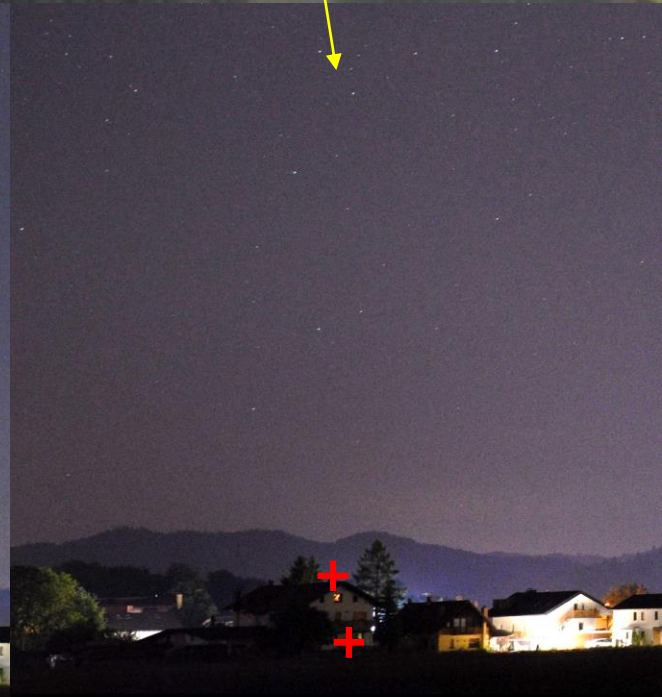
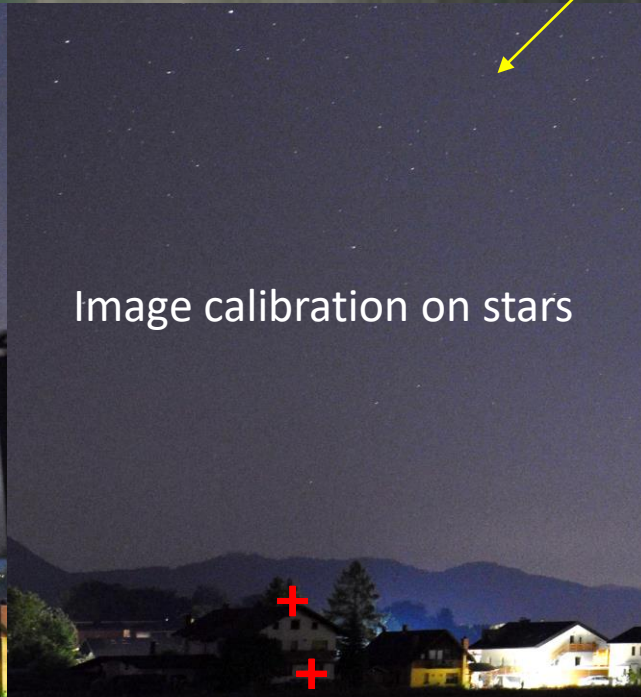
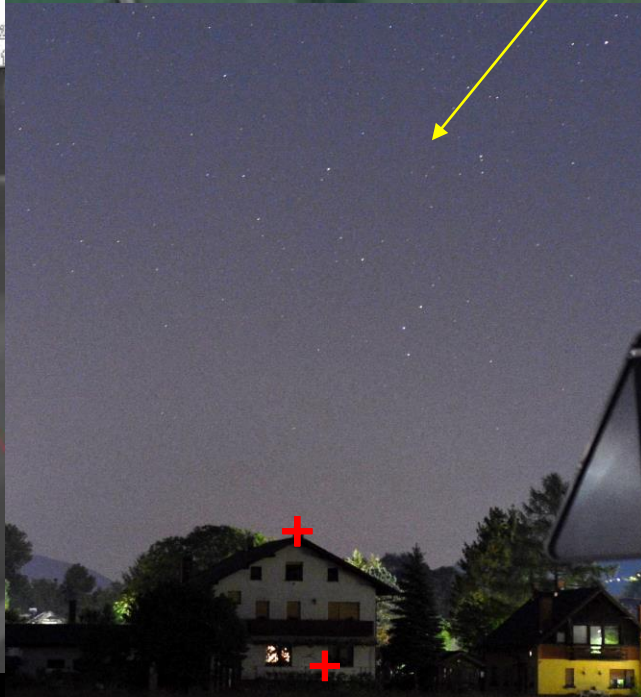


Image calibration on stars

Dashcam video

Loc 1

Loc 2

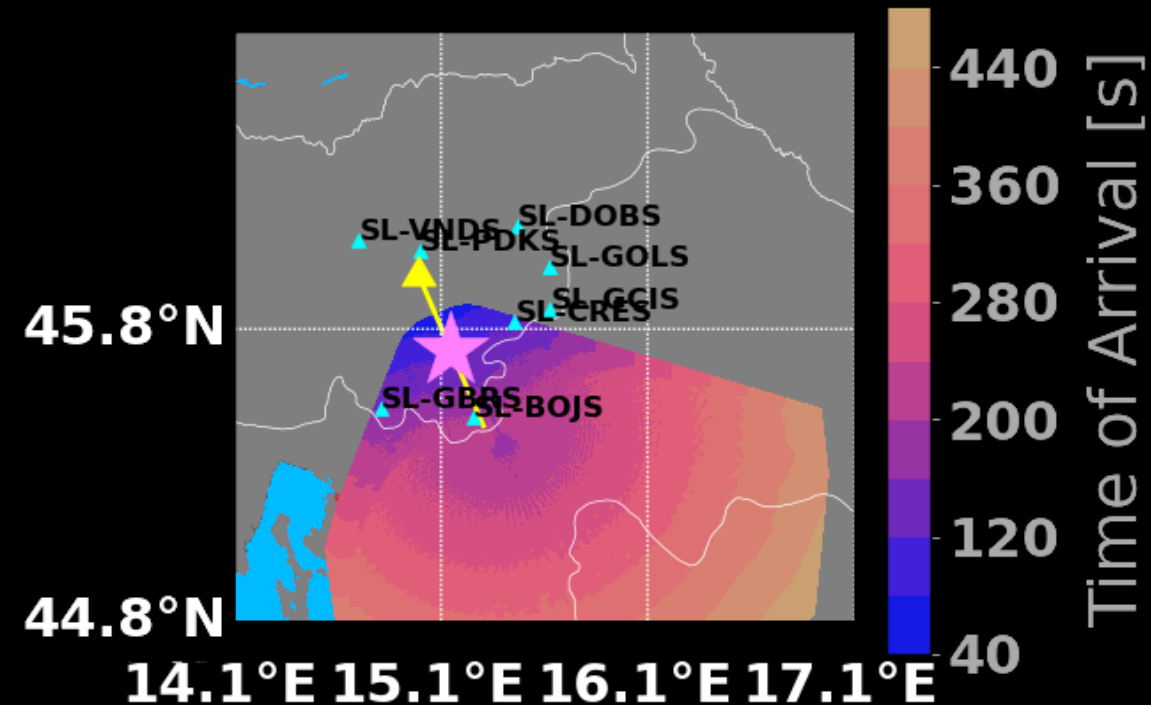
Loc 3

Ground-based observations

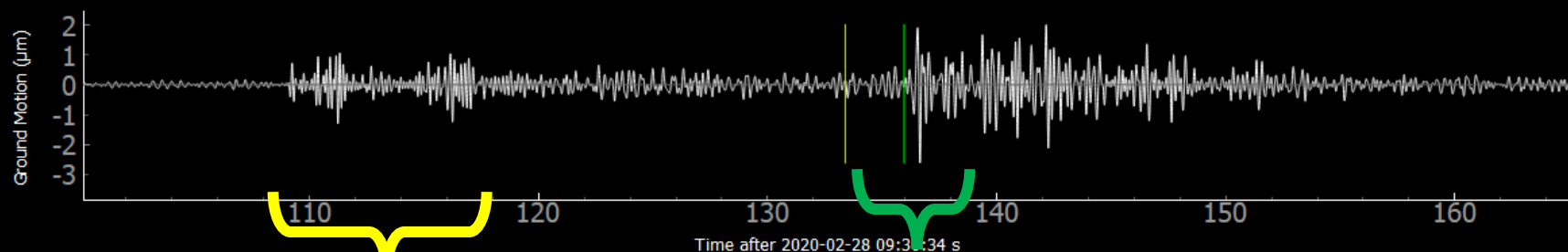
- Pros:
 - Dedicated fireball networks
 - Accurate dynamics
 - Resolving individual fragments and modes of fragmentation
- Cons:
 - Large objects rare
 - Large objects saturate sensors → Hard to invert an unsaturated energy deposition
 - Only casual data often available

Seismic data

- Fragmentation (point source) heard by 8 seismic stations
- Ballistic (cylindrical) wave heard by SL-BOJS
- Arrivals 10 s of seconds earlier may be from seismic propagation (precursor waves)



Contour shows expected arrival times from the ballistic source



Precursor?

Consistent with Fragmentation at 35 km