DART



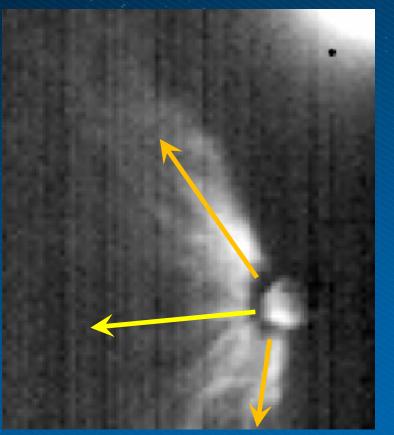
3D Characterization of the Ejecta Produced by the DART Impact

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#166

Measuring Momentum Enhancement

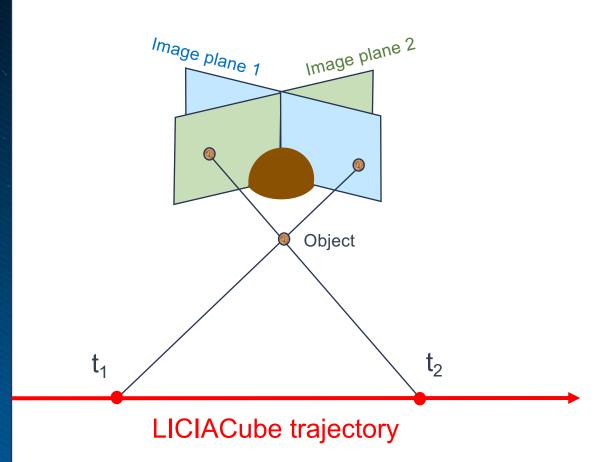
- The primary result of the DART impact: $3 < \beta < 4$
 - Measurement governed by observing changes in orbital period
 - Reflects only excess momentum parallel to the orbital velocity
- LICIACube and HST observations show significant amounts of ejecta emitted in other directions
 - How might this additional momentum affect the system?
- DART investigation team are working to characterize the ejecta field
 - Posters modeling the shape of the ejecta cone:
 - Hirabayashi et al. "DART-driven ejecta cone geometry measurement from Hubble Space Telescope and LICIACube
 - Deshapriya et al. "Constraining the ejecta cone geometry following DART impact on Dimorphos using LICIACube data"



LICIACube LUKE image, Impact +172 s

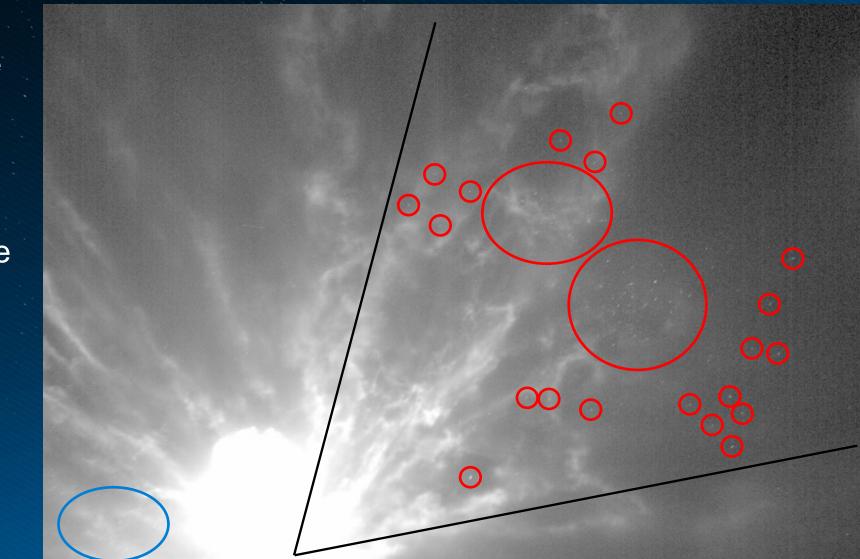
Ejecta Measurements in Three Dimensions

- Use parallax induced by LICIACube's motion to solve for positions of features in LUKE images
- Changes in the apparent position WRT Didymos define the object's location in 3-D space
- Multiple viewpoints can improve the accuracy of the position measurement and constrain the velocity



Initial Work: Individual Boulders

- 100+ "point sources"
 - Mostly clustered to the upper right direction
 - A few to lower left
 - Visible in multiple frames
- Presumably these are individual boulders ejected in the impact
- Clustering suggests that the boulders were ejected in preferred directions





- Determine physical properties of the boulders
 - Use photometry to derive size and "shape" of the boulders
 - Use parallax to derive location in space
 - Derive ejection velocity from parallax or distance/time
- Compute their contribution to momentum enhancement

Boulder Sizes

Use photometry from calibrated LUKE images to derive the boulder sizes

Derive the object size from measured Irradiance

 $R = \{ 4 (I_{\text{meas}} / I_{\text{Sun}}) (r \Delta)^2 / (p f(\alpha)) \}^{1/2}$

- Absolute calibration is not yet available
 - Bootstrap brightness off calibrated DRACO images of Didymos
- Brightest boulder → ~2 m radius
- Brightnesses of most boulders $\rightarrow \leq 1 \text{ m radius}$
 - Lower limit defined by the instrument sensitivity/contrast
- Variations in lightcurves suggest boulders are irregular in shape



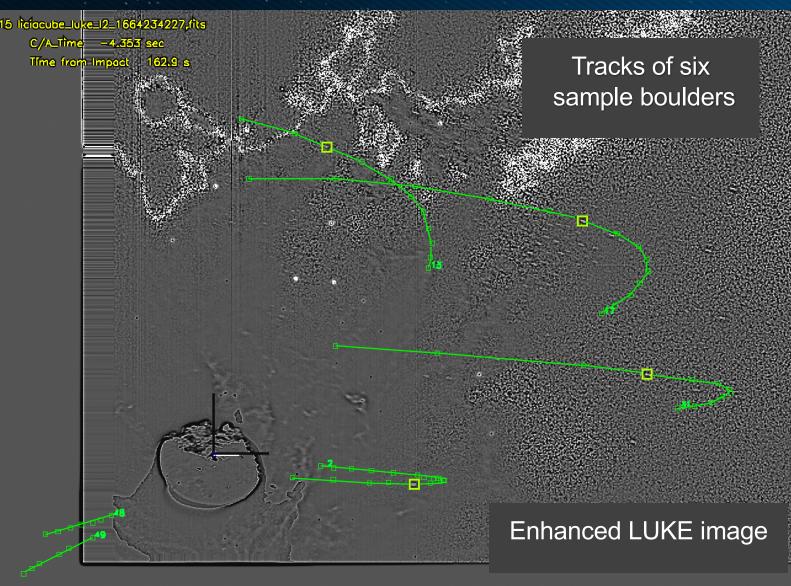
Example boulder sizes

To date, 91 boulders have been tracked

 Another 20-30 can be seen in long exposures but are too faint to track

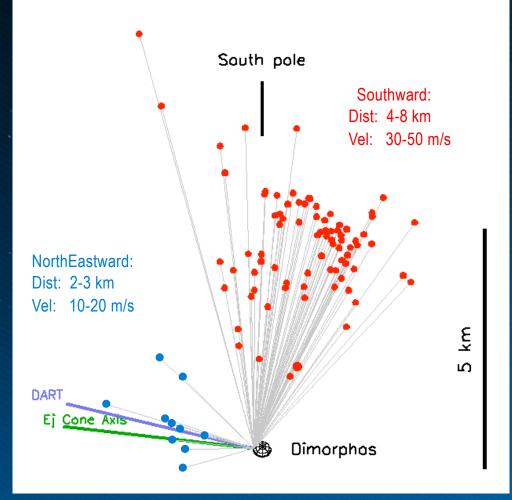
Paths form characteristic horseshoe shapes

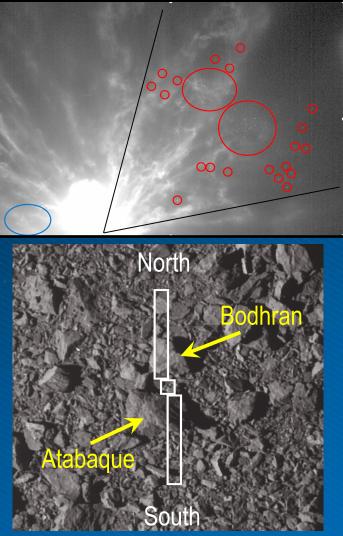
Tracking Boulders



Preliminary Results – Spatial Distribution

- Two populations
 - Large cluster ejected
 Southward
 - Small cluster ejected NorthEastward
- High velocities for meter-sized boulders
- Preferred directions + experimental results
 - Possibly the remains of boulders Atabaque (S) and Bodhran (NE) ?





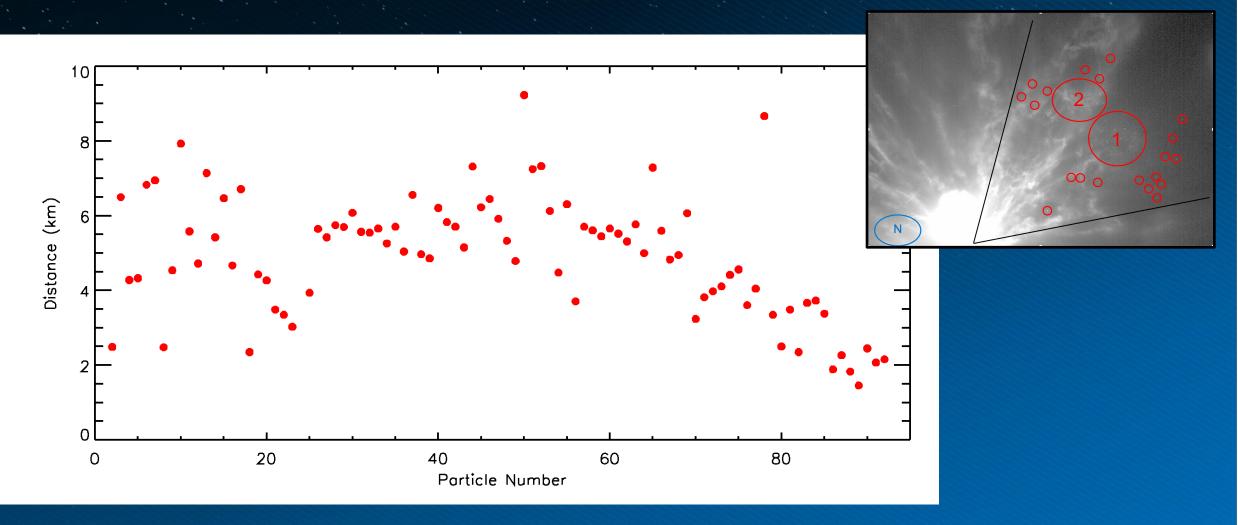
Note the different orientation

Summary

- Results based on tracking of 91 boulders
 - Boulders typically ≤1m in radius
 - Generally divided into two populations / directions
 - Most boulders clustered in Dimorphos' South polar direction
 - A few clustered to the NorthEast direction
- Momentum from measured boulders is likely comparable to DART momentum
 - Primarily to the South, perpendicular to DART's path
 - What are the effects of these components?
 - Dimorphos' rotation, orbital inclination and eccentricity
 - What about boulders in other directions?
- Next up: derive locations of features in the ejecta clouds

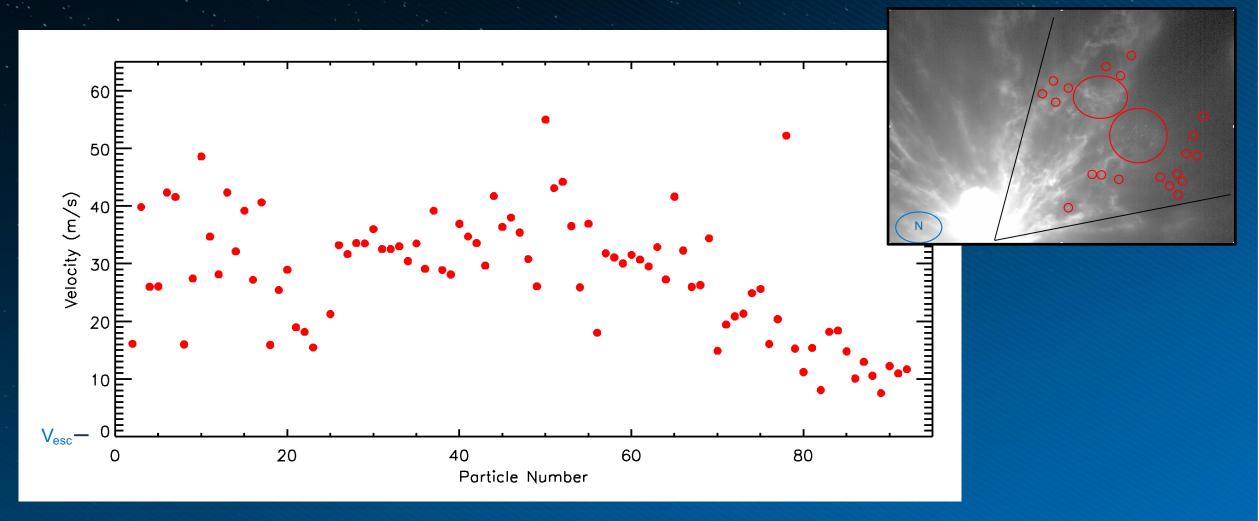


Preliminary Results - Distance



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Preliminary Results - Velocity



Velocity = Distance / 167 sec