



The ESA Hera mission: Detailed Investigation of the NASA DART Impact Outcome and Characterization of the Binary Asteroid Didymos

Dr. Patrick Michel

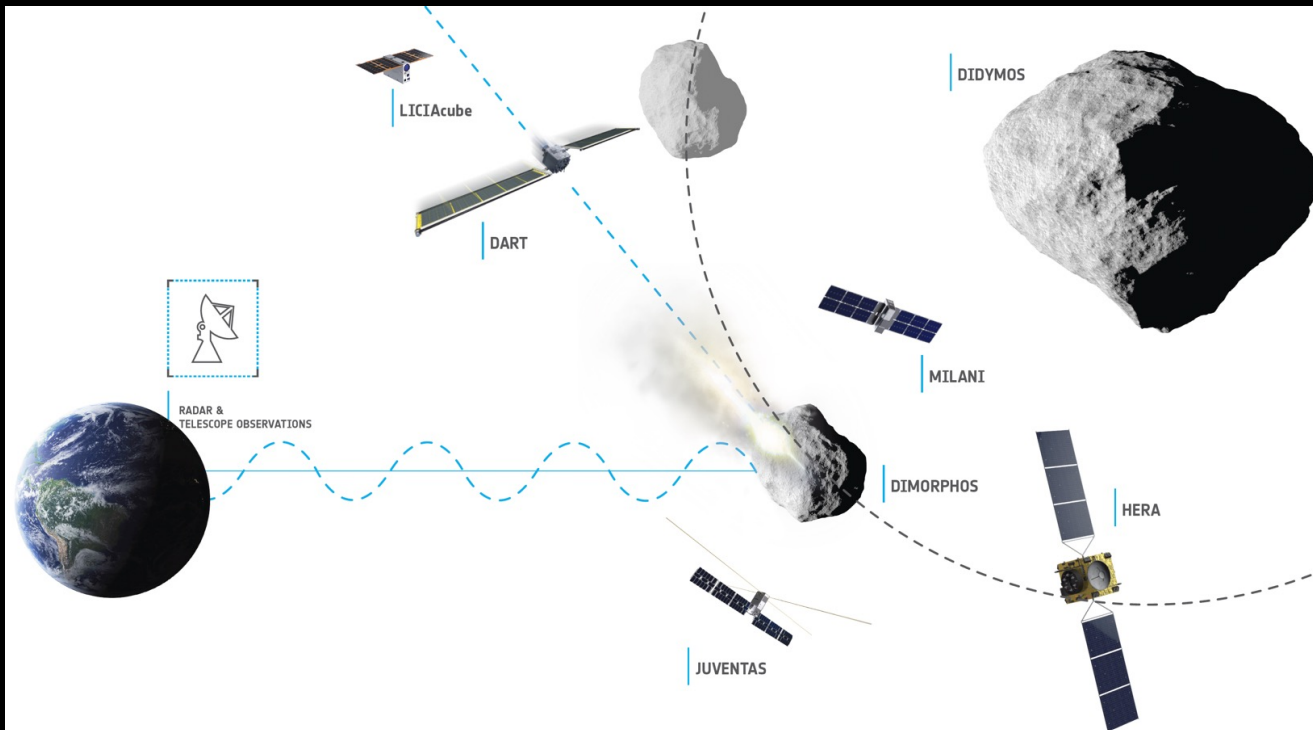
On behalf of the Hera Science Team

Hera mission Principal Investigator

Univ. Côte d'Azur, Obs. Côte d'Azur, CNRS, Lagrange Lab.

Michel, P. et al. 2022. Planet. Sci. J. 3, 160
(PSJ Special Issue DART/Hera pre-arrival)

AIDA collaboration for planetary defense



Successful impact on September 26th
at 23:14:24 UTC

LICIACube successful imaging of the
ejecta



Big questions! Didymos & Dimorphos shapes



Observations of the two components on all sides are missing!!



Strong implications for binary formation

Dimorphos



What is its
mass?
Is it an
aggregate?



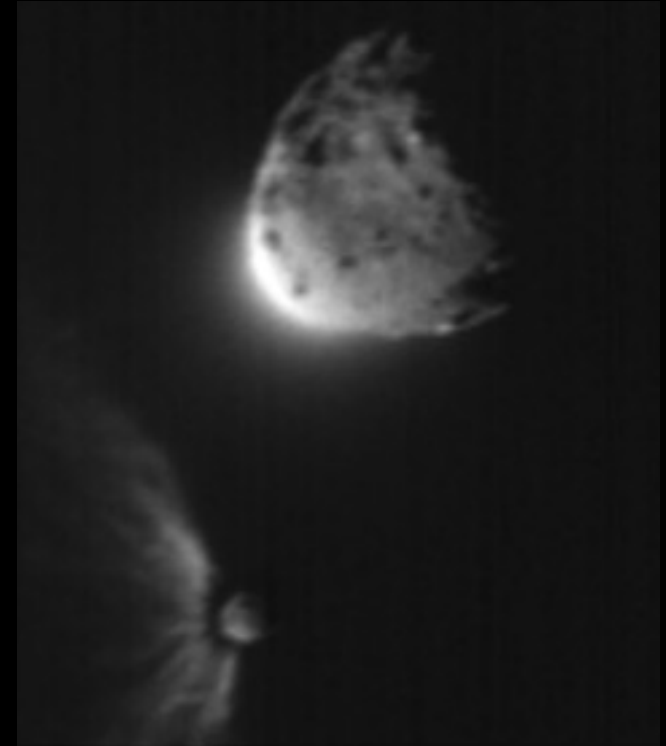
Credit: NASA/Johns Hopkins APL

Strong implications for binary formation, small body geophysics & impact physics



What does Dimorphos look like after impact?

Seeing the « final » Dimorphos

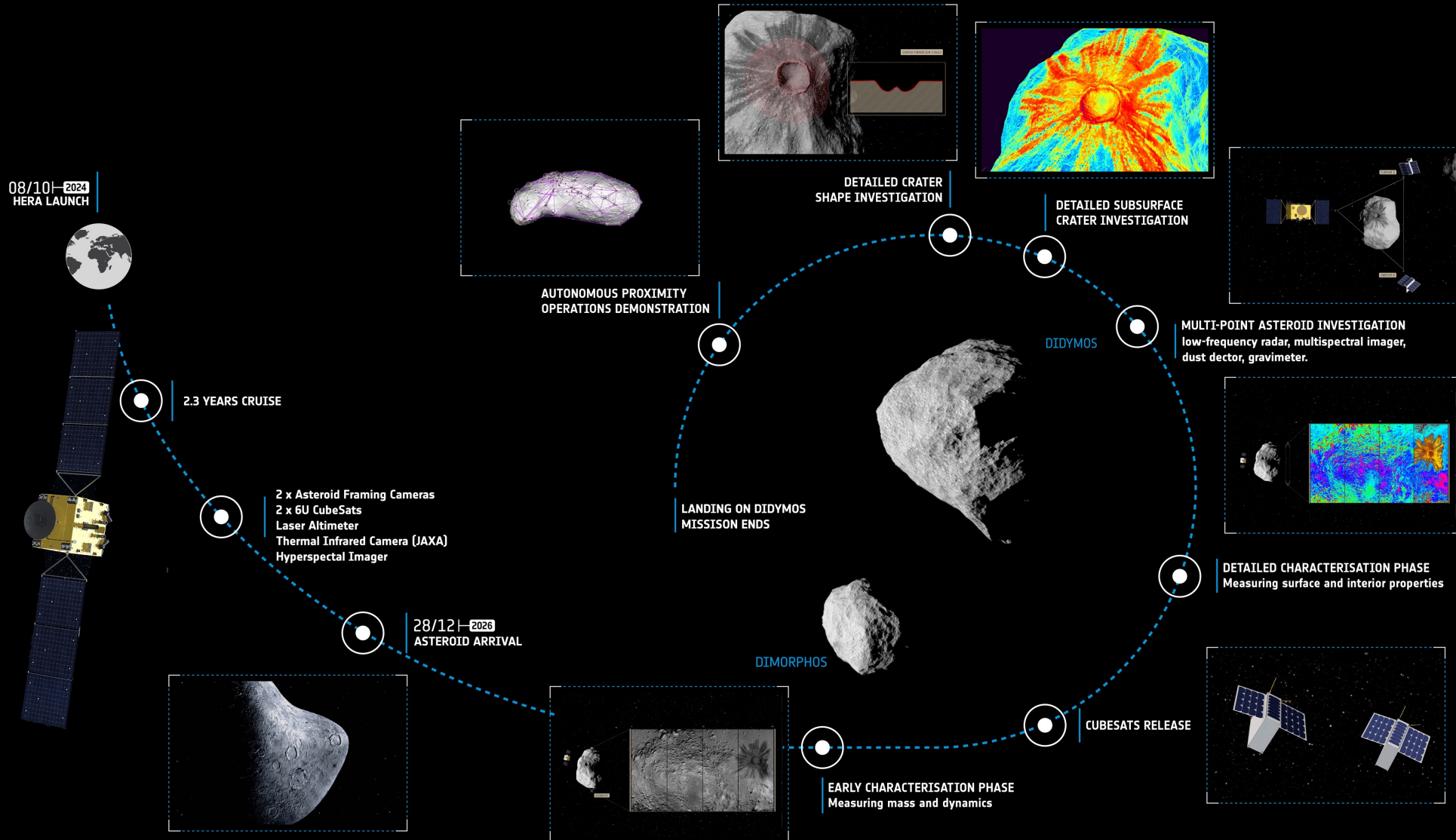


Credit: NASA/ESA/STScI/Hubble

Credit: ASI/NASA

Only way to interpret DART impact, to understand the response of such small rocks in space & to validate impact models

Hera mission



- First mission to rendezvous with a binary asteroid and smallest asteroid ever visited
- First radar tomography of an asteroid
- First full-scale cratering physics experiment investigation
- First deep-space CubeSat for very close asteroid inspection

Hera spacecraft

Launch: 8 Oct 2024

Satellite integration completed in 2023

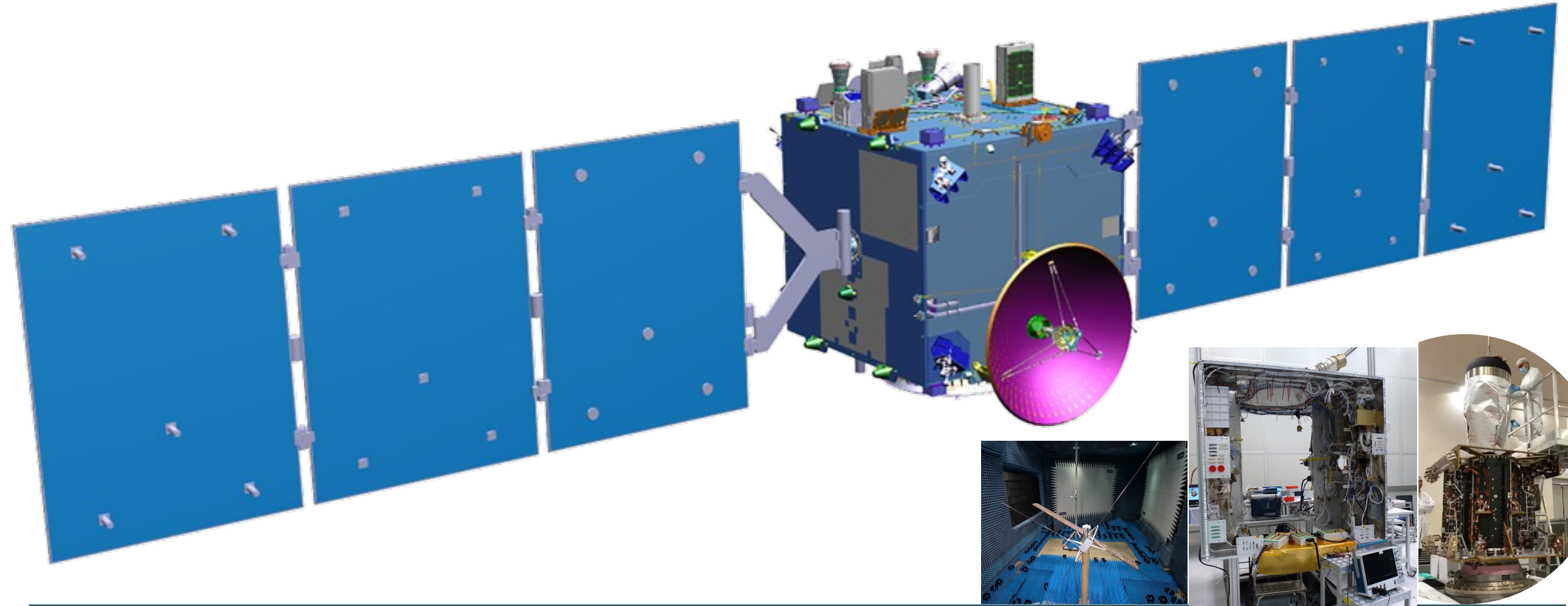
Tests at ESTEC in 2023/24

Mass: 1215 kg

Power: 826 W

Size: 2m × 11m × 2m

Countries: 17 EU + Japan





DART & Hera: First fully documented impact experiment on a 160 m-size asteroid at asteroid collision speed



Target properties



- Cohesive strength - **not known**
- **Mass - not known** → momentum transfer efficiency
- Bulk density / porosity - **not known**
- Internal structure - **not known**

Hera

Impact conditions



- Impact velocity - **known**
- Impact angle - **known**
- Impactor mass/shape - **known**

+ detailed impact outcome

Hera mission objectives



CORE asteroid investigation requirements

- Mass of Dimorphos
- Global properties (volume, linear scale, density)
- Size distribution of surface material
- Dynamical properties of the Didymos system
- Shape and volume of DART impact crater
- Size distribution of excavated material



OPPORTUNITY asteroid investigation requirements

- Surface strength
- Interior structure of Dimorphos
- Composition of Dimorphos
- Transport of impact ejecta from Dimorphos to Didymos

Hera Payloads



Core Payload

Asteroid Framing Cameras (AFC)



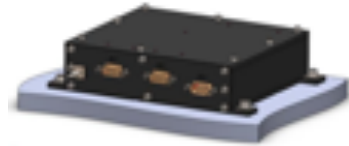
Visible camera for GNC and asteroid science

Qty 2 (Nom + Red)



Opportunity Payloads

Planetary ALTimeter (PALT)



Small Monitoring Camera (SMC)



Deep Space Deployer (DSD)



Low velocity deployer for CubeSats

Inter-Satellite Link (ISL)



S-band comms between Hera and CubeSats, with ranging and radio science

Image Processing Unit (IPU)

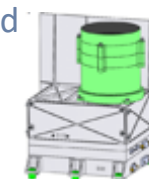


Experimental aid for on-board autonomous GNC

Thermal InfraRed Imager (TIRI)



Thermal infrared camera with filter wheel



Hyperscout-H



2D snapshot hyperspectral imager



CubeSat "Juventas"



CubeSat with low-frequency radar and gravimeter, incl. radio science with ISL



CubeSat "Milani"




CubeSat with multispectral imager and dust analyzer payloads




Hera instruments



AFC – Asteroid Framing Camera 
Redundant visible cameras used for both spacecraft guidance and science purposes

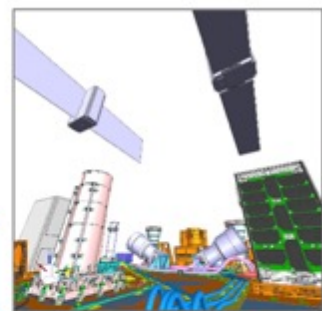
HyperScout 
Multispectral imager

TIRI 
Thermal InfraRed Imager


Milani CubeSat 
Multispectral imager
Retroreflectors
Dust analyzer
Radio-science

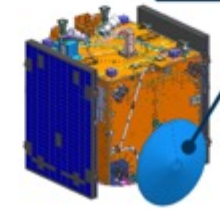
Juventas CubeSat 
Radar, Gravimeter, Radio-science

SMC 
Monitoring Camera



Radio Science Experiment 

PALT – Planetary Altimeter 
Laser-range finder for science & navigation



First detailed investigation of a binary asteroid

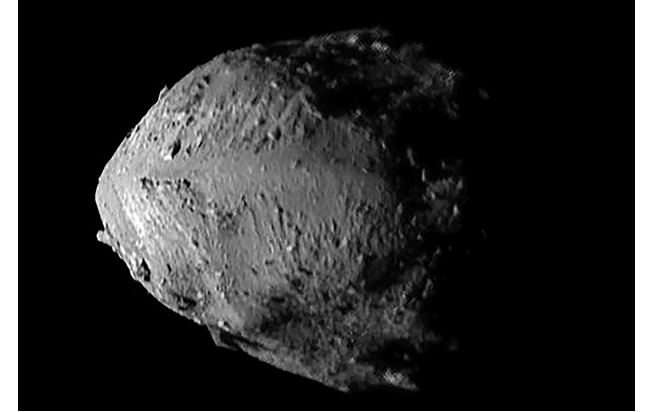


15% of NEAs are binaries

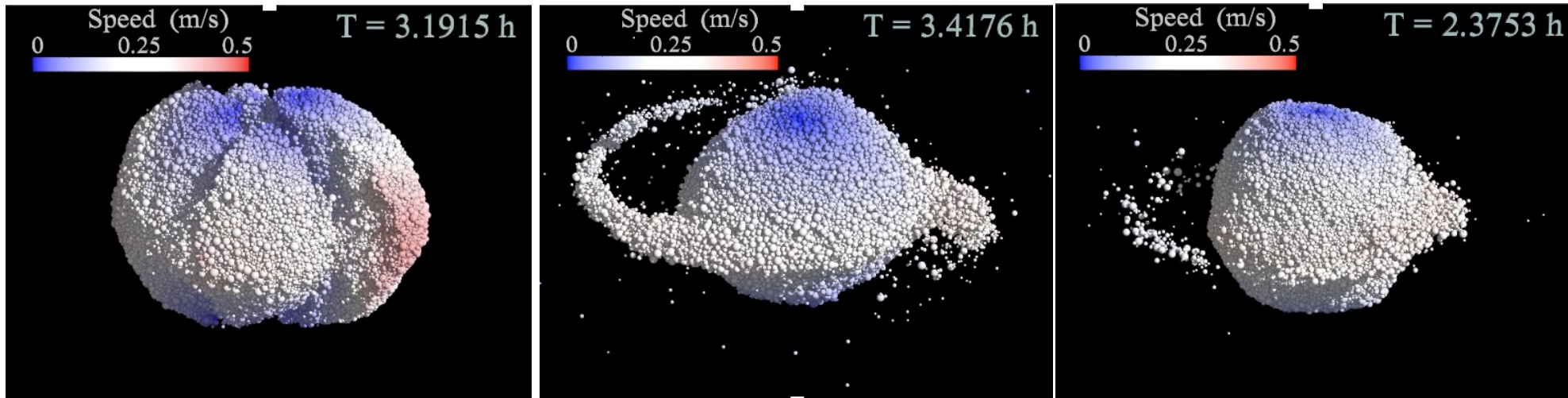
- Are Didymos and Dimorphos rubble piles?
- Why does Dimorphos have an oblate spheroidal shape?

Structural processes and evolutions

- First asteroid visited that may be at the limit of structural stability (**Didymos spin period is 2.26 hr**)
- Insights on the geophysics of fast spinning tops



Credit: NASA, APL



Zhang, Michel, Richardson, et al. 2021, Icarus 362

Structural stability and spinup outcome depends on cohesion, at the level of 1 Pa!!

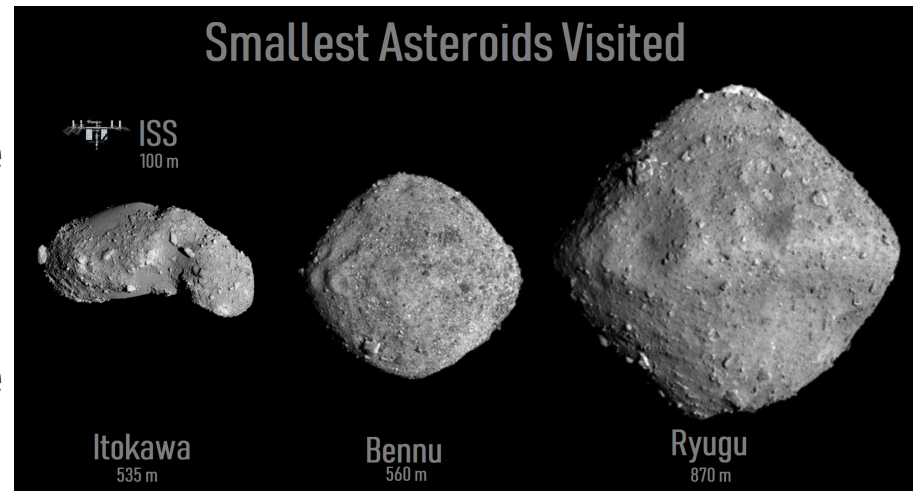
First geophysical data on a 160 m-size object



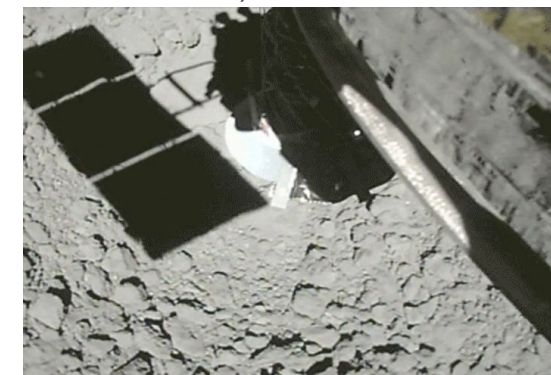
- H2 and O-REx : 1st detailed characterization and surface response of 2 primitive asteroids in different gravity conditions



- Ryugu is 900 meter wide
- Bennu is 500 meter wide



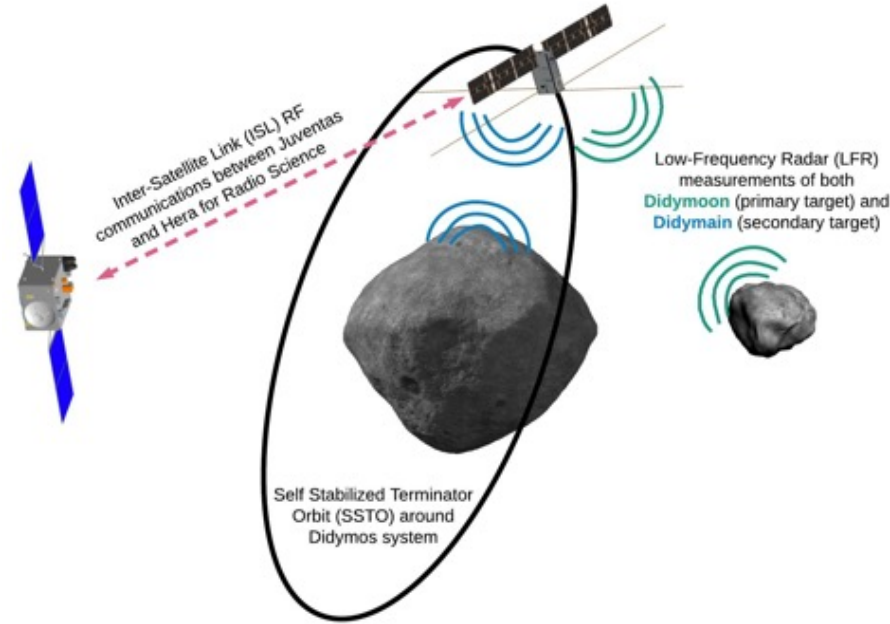
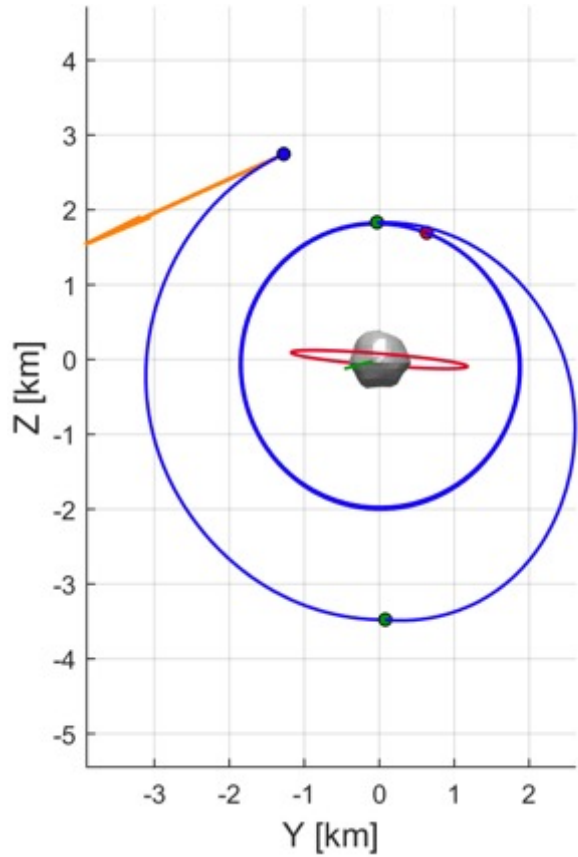
NASA, Univ. Arizona



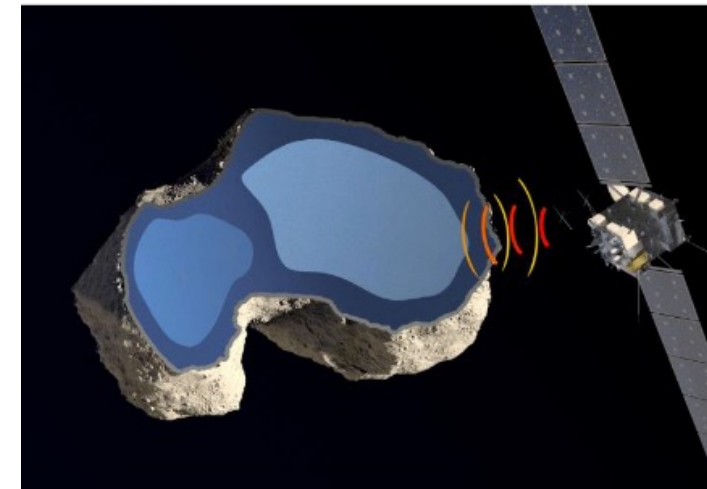
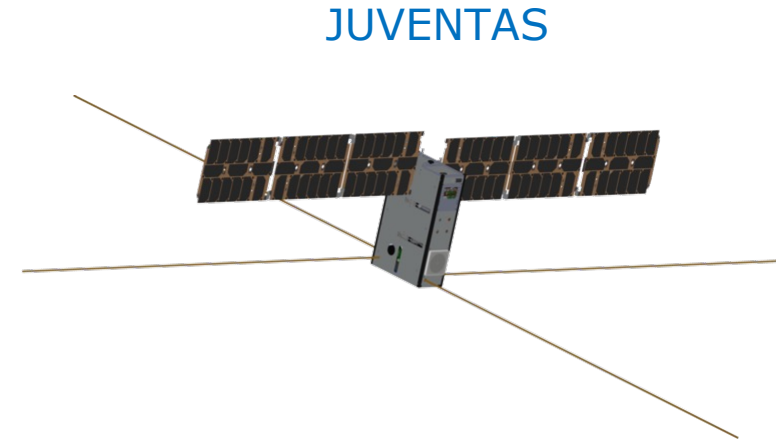
JAXA, Univ. Tokyo and Co

- Dimorphos is ~ 3 times less wide than Bennu
- Possibility to understand how some processes scale with gravity down to the low-g of Dimorphos

First internal probing of an asteroid



Asteroid interior probing by radar tomography (a first)



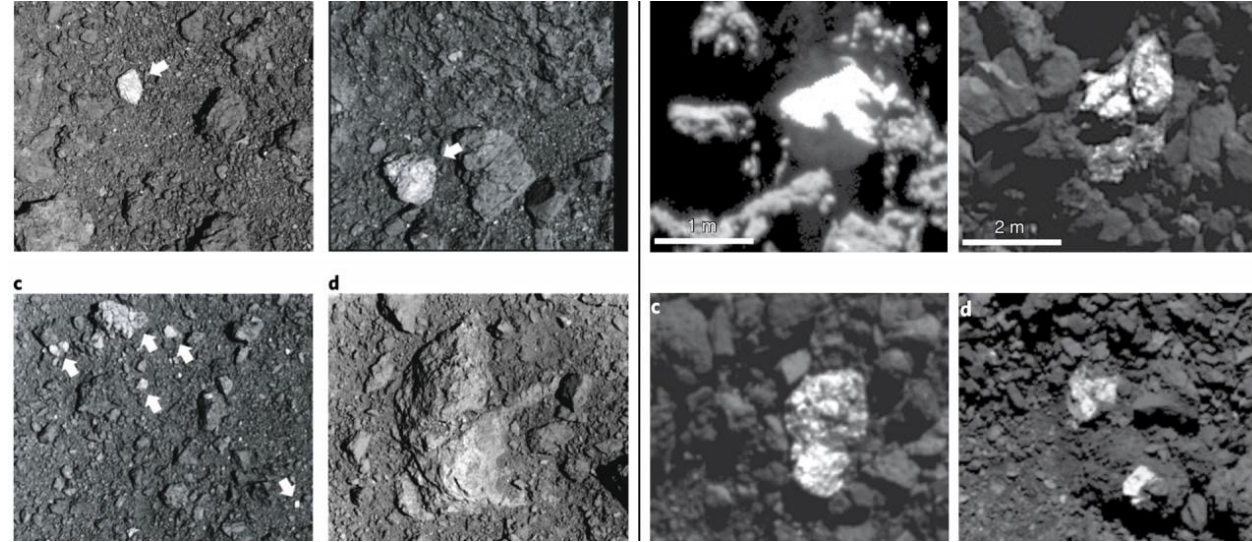
What is the level of heterogeneity in the interior of Dimorphos?

First mineralogical study of a 160m-size objects and its subsurface revealed by the DART impact

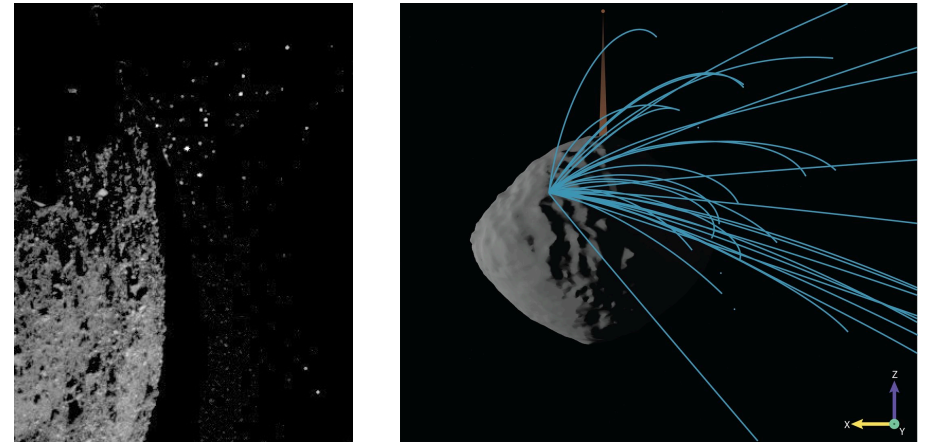


Milani Cubesat will investigate in details the mineralogy and possible surrounding dust

- Is the composition homogeneous?
- Is there dust in the system?
 - Exchange of material between Didymos and Dimorphos
 - Micrometeorite impacts

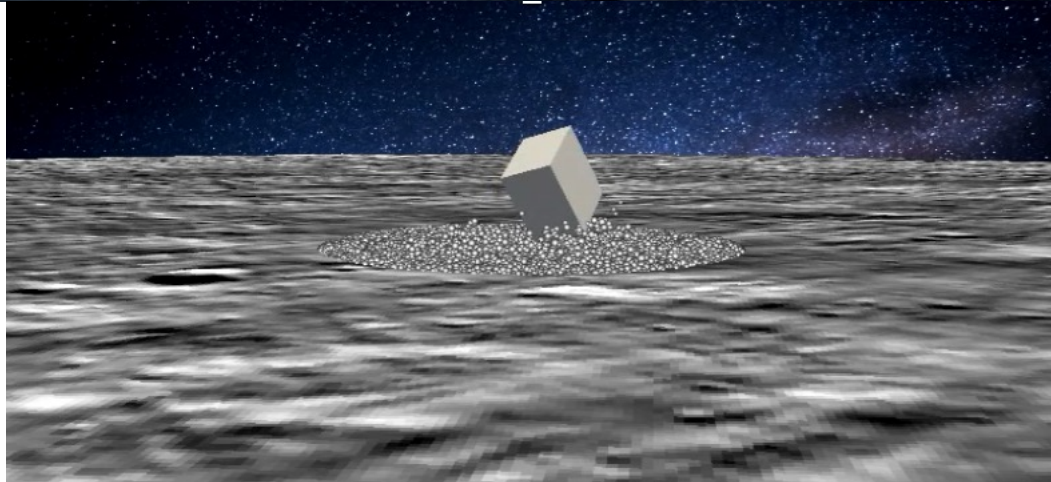


Exogeneous material on Ryugu & Bennu



Dust ejected from Bennu

First landing on a 160 m-size object



Cubesats Juventas & Milani

Landing on Dimorphos

- Bouncing properties
- Only way to determine the response of these bodies in their low gravity environment
- Implications for science, planetary defense and mining

Close-ops timeline



1. Early Characterization Phase (ECP)

- Distance ~20-30 km (entire system in FOV of AFC)
- Ground-controlled SC \Rightarrow ensure observations of Didymos system

2. Payload Deployment Phase (PDP)

- Release of CubeSats from passively safe trajectory similar to ECP

3. Detailed Characterization Phase (DCP)

- Distance ~9-20 km
- Autonomous navigation based on AFC only (Didymos LOS navigation)
- CubeSats operations (multi-point measurements)

4. Close Characterization Phase (COP)

- Distance ~4-22 km (Didymos and/or Dimorphos LOS navigation)
- Dimorphos high-resolution imaging and full characterization of DART's impact crater

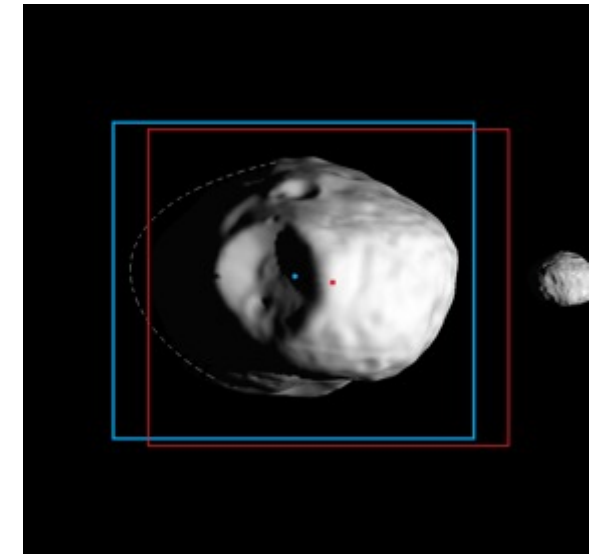
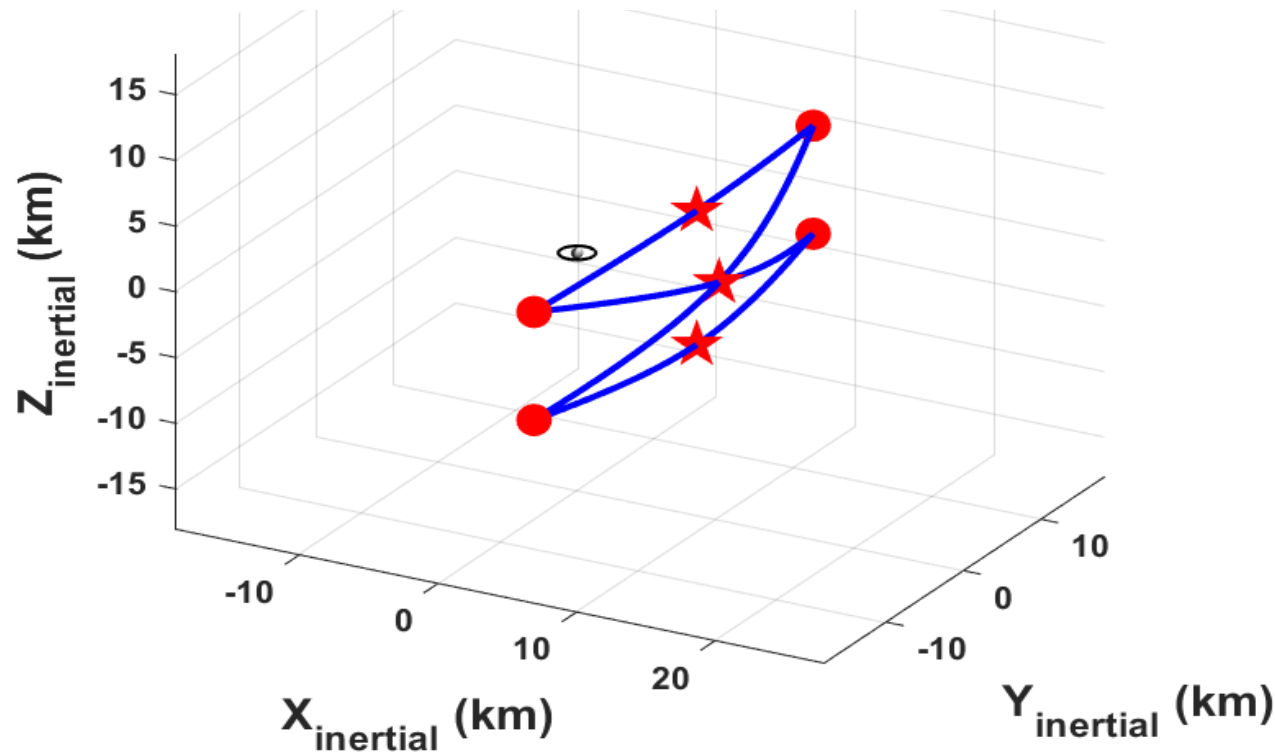
5. Experimental Phase (EXP)

- Very low altitude fly-bys ~1km (autonomous delta-V to lower pericentre)
 - PALT + feature tracking below ~2km distance
-

HERA Detailed Characterization Phase (DCP)



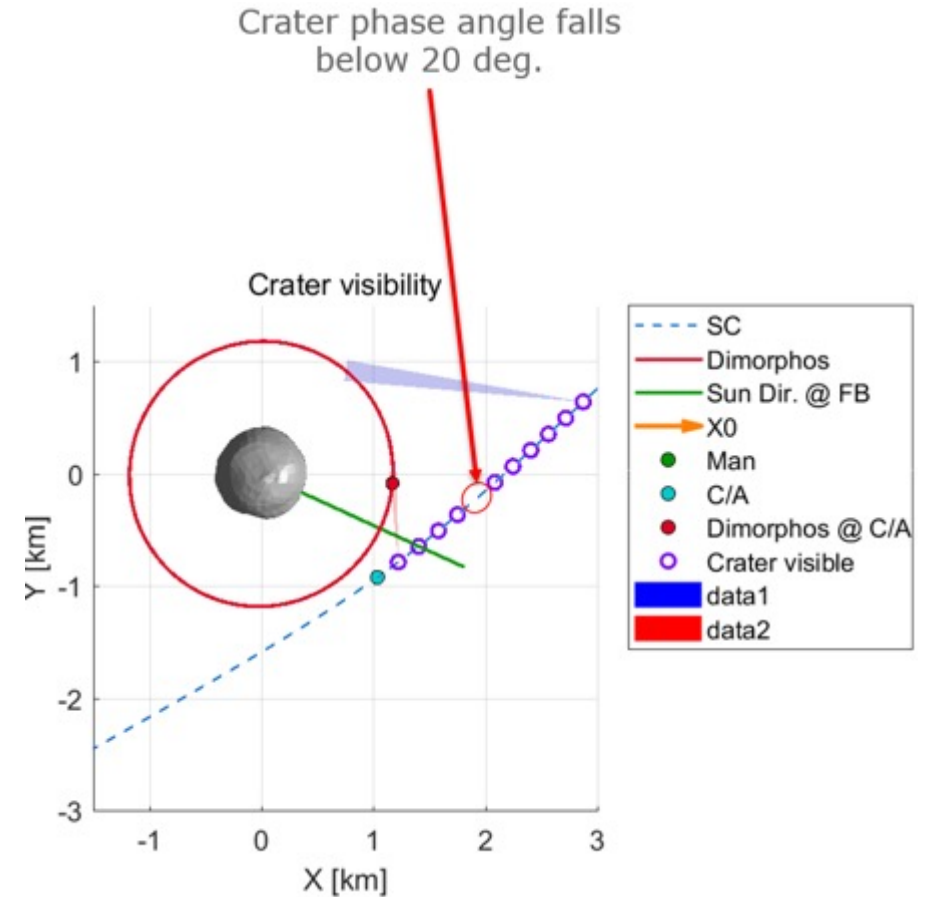
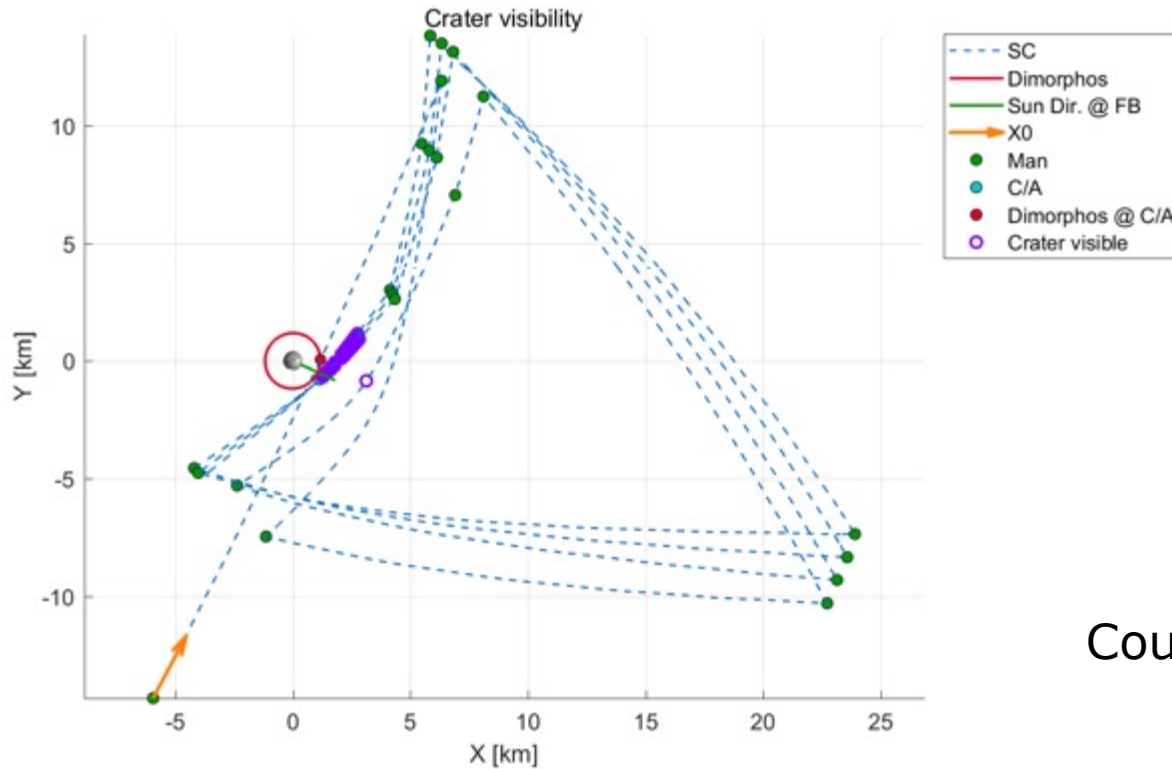
- Change trajectory shape to reach lower distances keeping constraints
- Autonomous Pointing \Rightarrow autonomous navigation
 - Pericenter around 10km to have full Didymos in FOV



HERA Experimental Phase (EXP)



- Pericentre is very close to Dimorphos orbit to reach crater imaging resolution
- Very-close fly-bys reaches night side and optical navigation is interrupted
- GNC must ensure that can take images back in the day side to recover

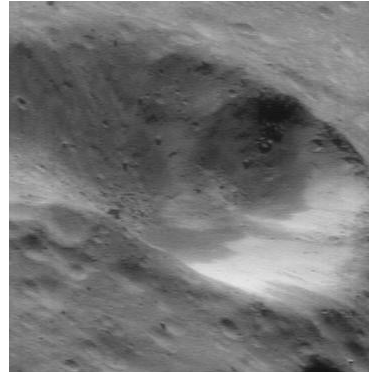
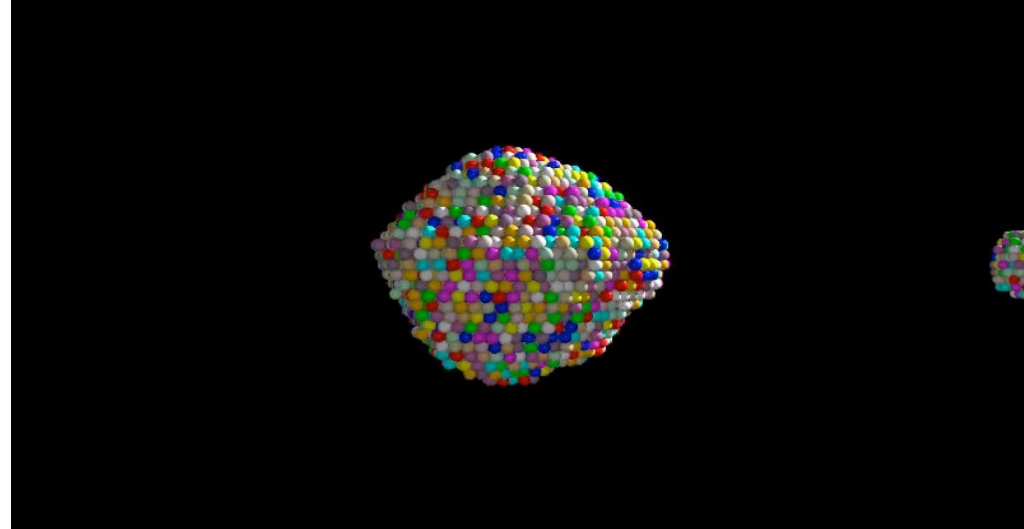


Courtesy GMV

Hera: a mission full of « firsts »

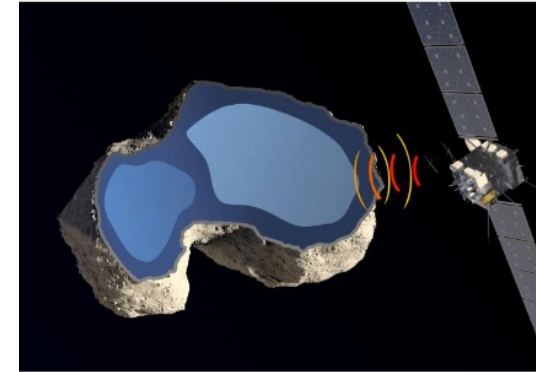


- **1st** rendezvous with a binary & smallest visited asteroid (Dimorphos)
- **1st** detailed outcome of an impact at a speed of 6 km/s
- **1st** Deep Space CubeSats & **close inspection of composition & internal, subsurface and surface structures**
- **1st** landing on an asteroid of 160 m in diameter



Credit: NASA/JPL/JHUAPL

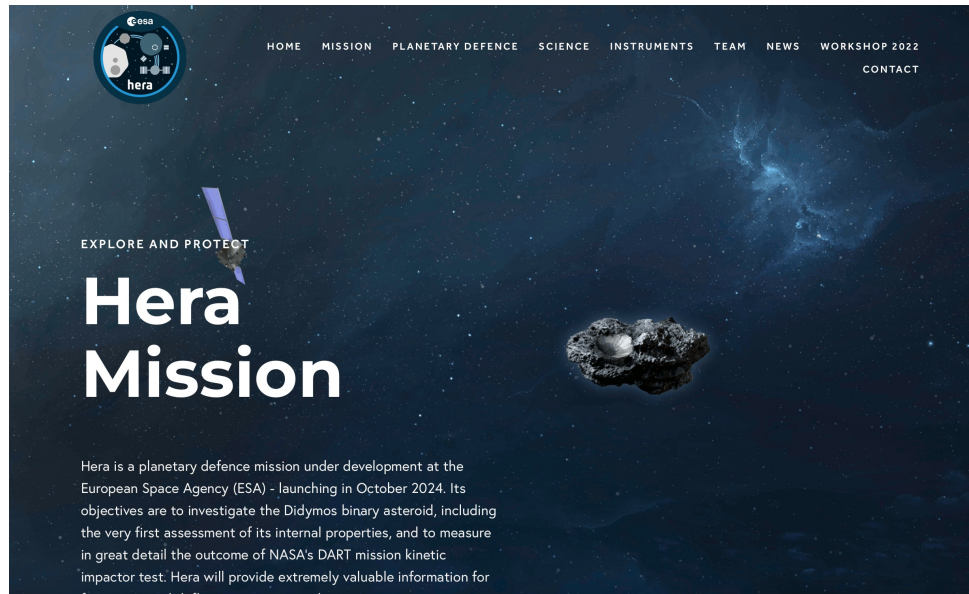
Analysis of the size and interior of a crater



Analysis of the interior of an asteroid

The structural analysis and measurement of the response to a surface interaction in low-gravity conditions are crucial knowledge for many purposes

Hera Community Website



Dr. Patrick Michel

Mission PI
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Dr. Michael Kueppers

ESA Project Scientist
ESA-ESAC, Spain

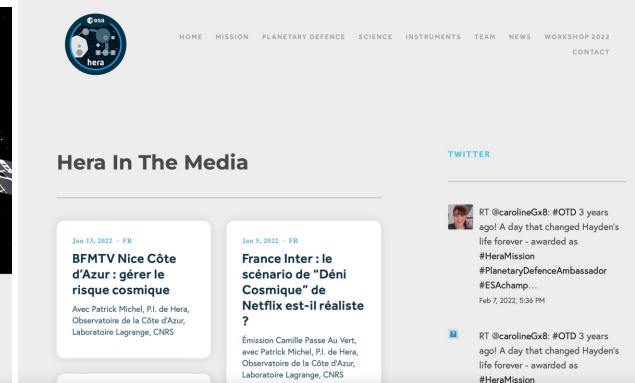
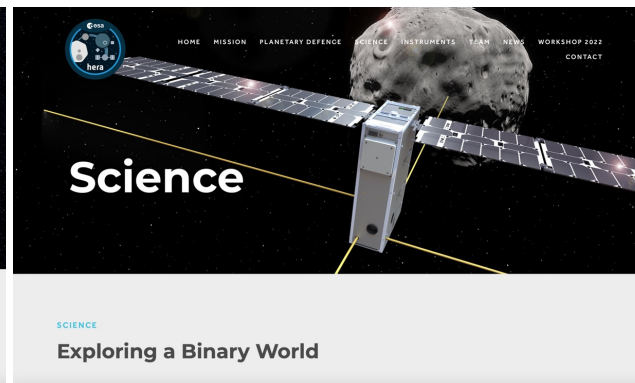
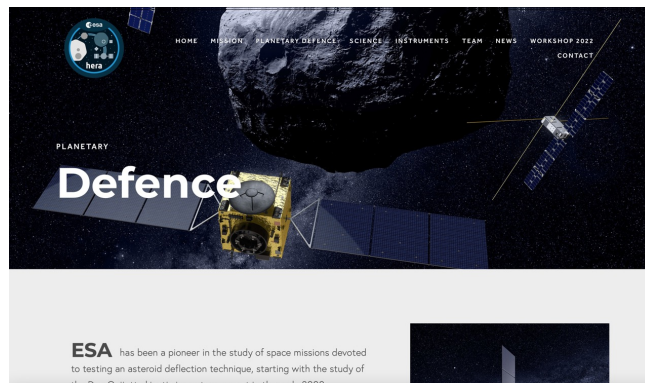


Ian Carnelli

ESA Mission Manager
ESA-ESTEC, The Netherlands

Hera Science Team

The Hera Science Team includes the ESA Project Scientist and the mission PI, who, together with the Advisory Board, compose the Science Management Board (SMB). It also includes the Investigation Team, which is composed of the Mission PI, the Advisory Board, the Working Group Chairs and some instrument representatives. There are also instrument PIs who are responsible for decisions about delivery and safe operation of their instrument, and an Operation Group with a Chair appointed by ESA for the operational aspects of the mission.



<https://www.heramission.space>

Conclusions



- Hera spacecraft development progressing at fast pace, **on schedule to launch in October 2024 onboard Falcon-9**
- Baseline trajectories achieve multiple geometries and observation conditions of the binary asteroid considering all constraints
 - With the assumption of tidally locked Dimorphos, almost half of Dimorphos surface can be imaged with a resolution of 40 cm or less
 - Rotation around Sun direction can allow deviations from hypotheses
- Great international team and synergy with DART promising breakthroughs in our knowledge