



# SIZE AND ALBEDO DISTRIBUTIONS OF NEAR-EARTH ASTEROIDS OBSERVED BY NEOWISE

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# NEOWISE Introduction and Survey Statistics

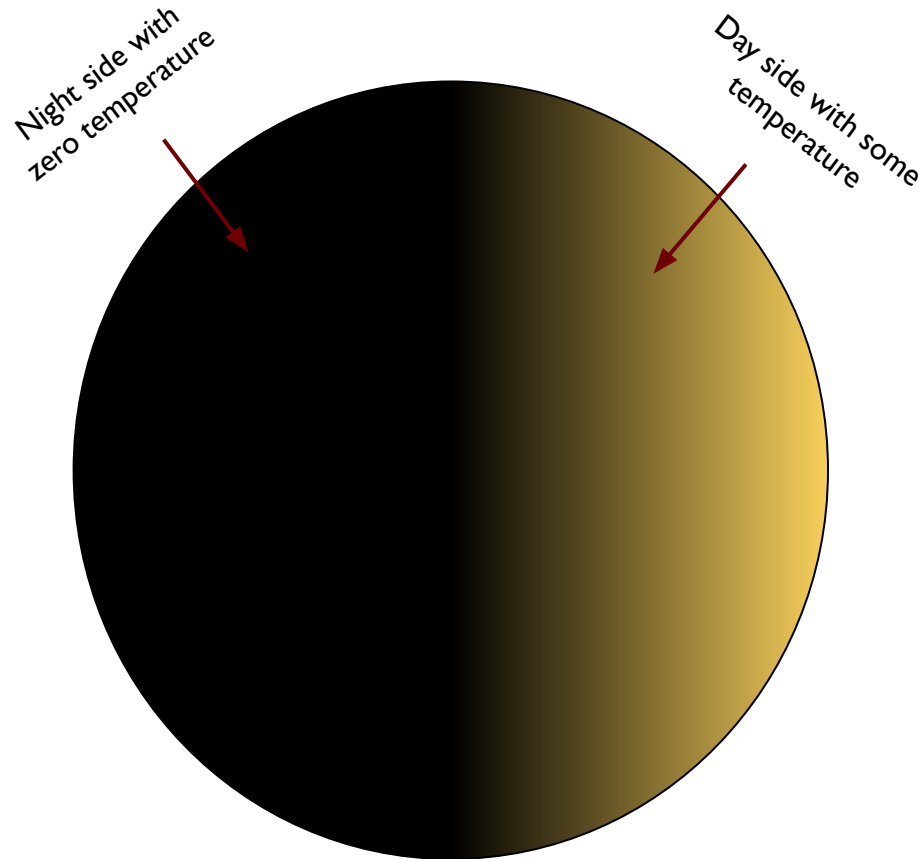


- Near-Earth Object Wide-field Infrared Survey Explorer or NEOWISE (Mainzer et al. 2014)
  - Extended part of the WISE mission (Wright et al. 2010)
- Two-band (3.4 and 4.6 micron) all-sky thermal infrared survey
- Since its reactivation in 2013, NEOWISE has detected tens of thousands of minor planets in the solar system
- NEOWISE's 9<sup>th</sup> data released on March 2023!
- Dataset overview
  - Total NEO detections: 5686
  - Total unique NEOs: 3705
  - Total number of thermally selected fits > 2198

NEOWISE spacecraft pre-launch  
Credit: NEOWISE/NASA/JPL



# NEATM model



**NEATM temperature approximations**  
(this is a general example)

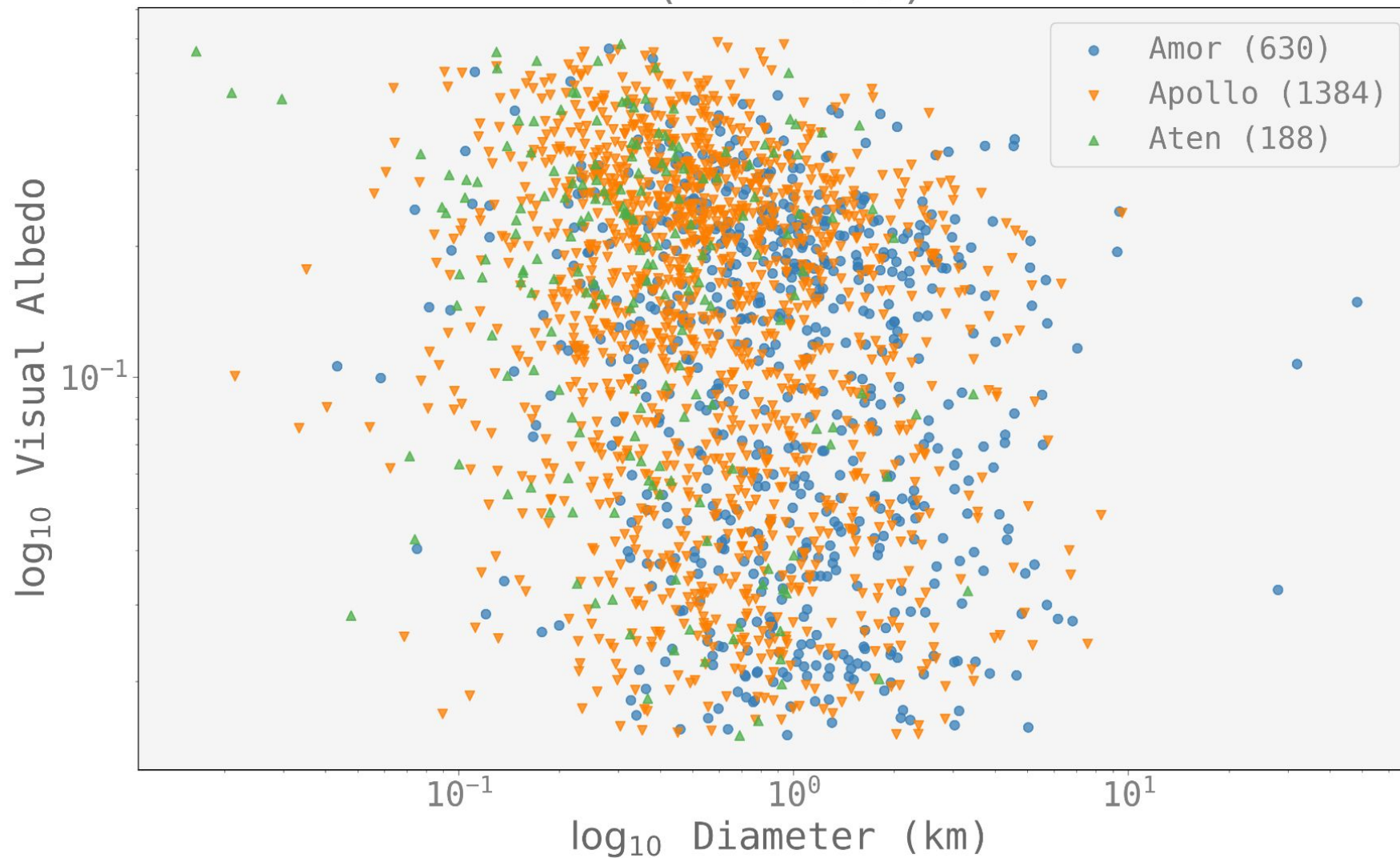
- Near-Earth Asteroid Thermal Model (Harris 1998)
- Asteroid = idealized sphere
- Beaming parameter to account for surface roughness, thermal conductivity, period, etc.
- Computationally efficient
  - Scalar minimization
- We fit —
  - Diameter and albedo for 2-band data
  - Diameter, albedo, and beaming for >2-band data



# Results



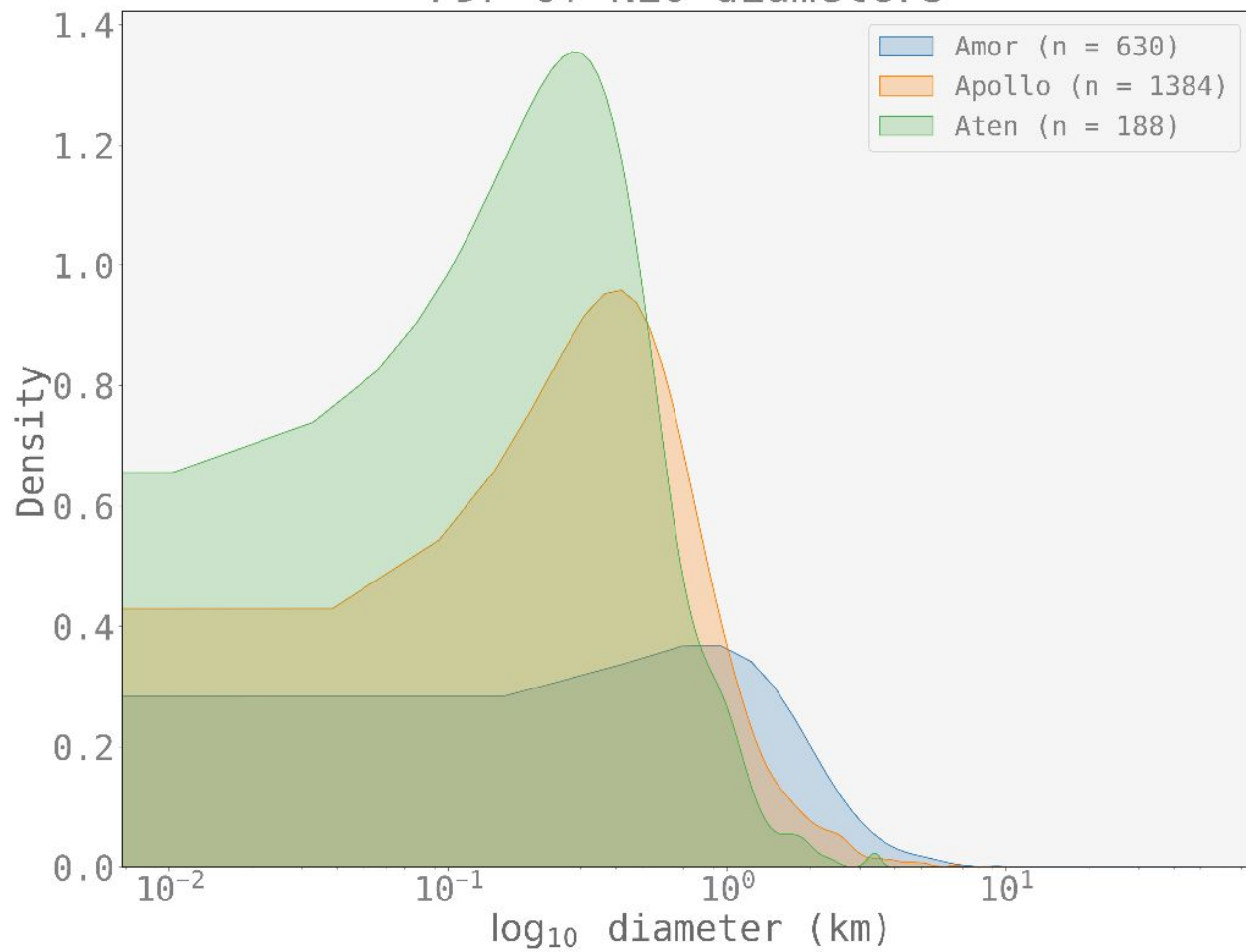
Diameter vs. Albedo  
(n = 2198)



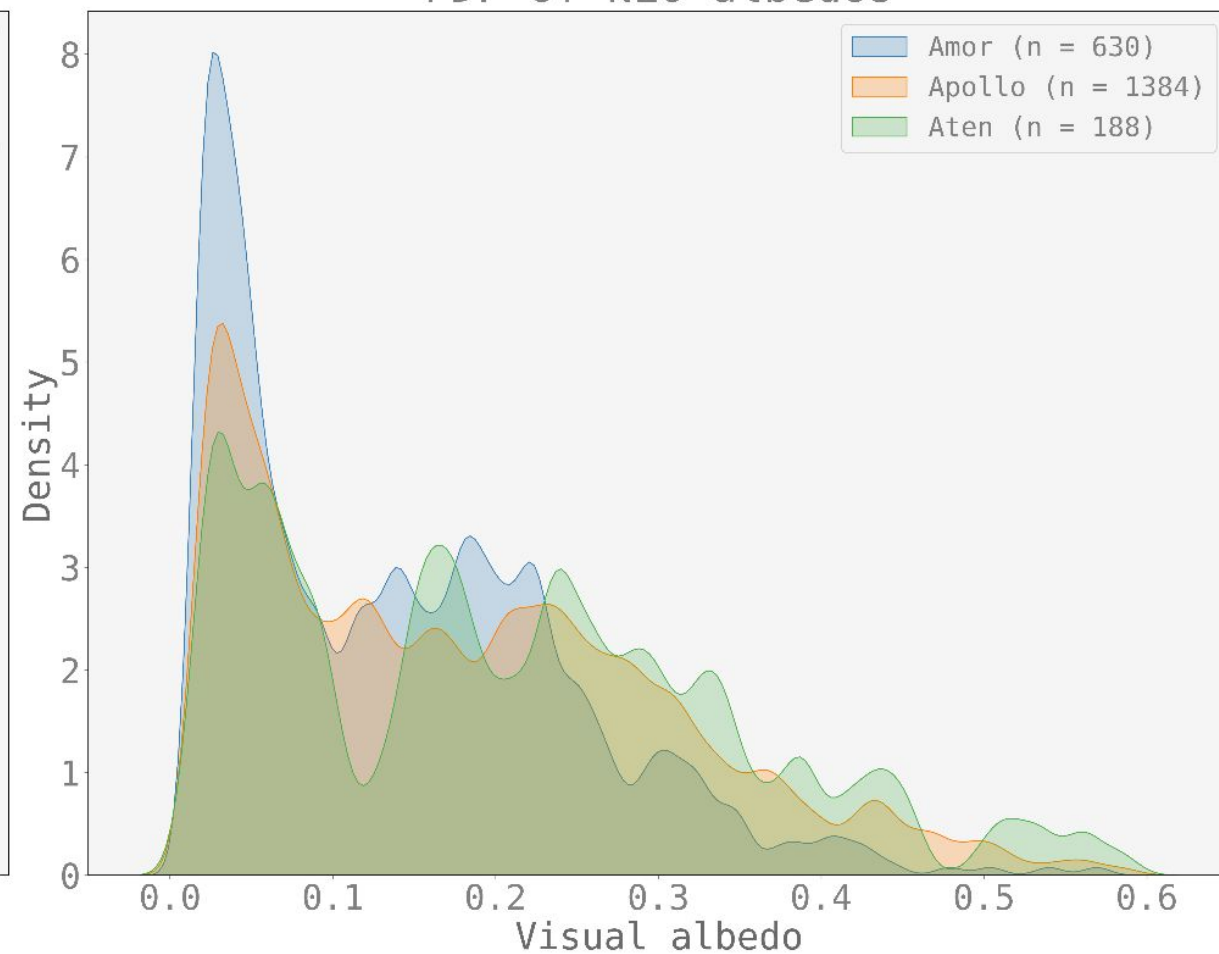
# Results



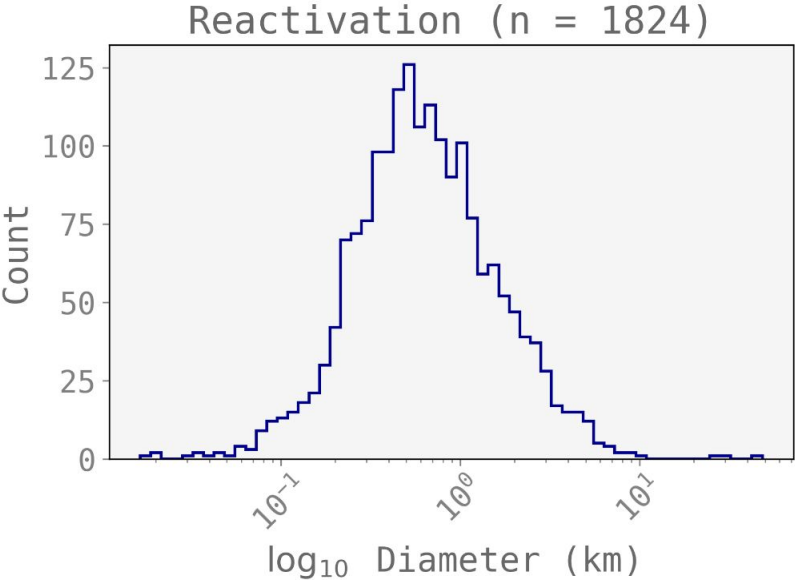
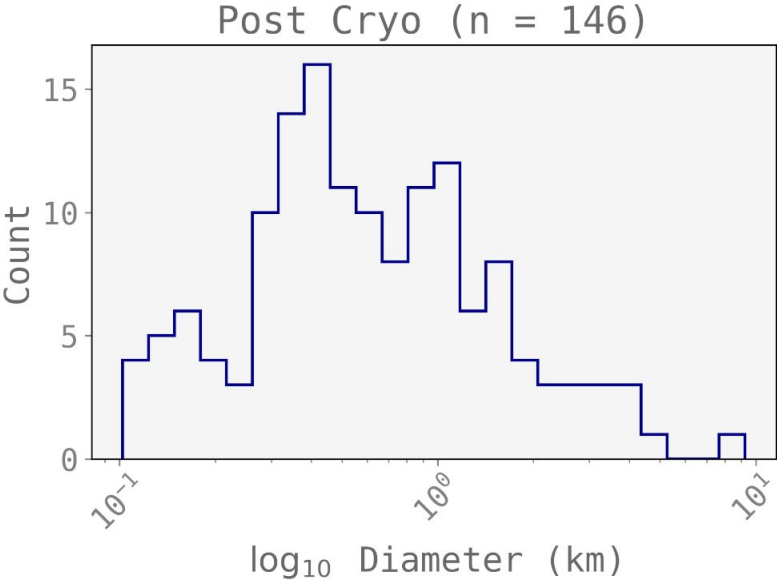
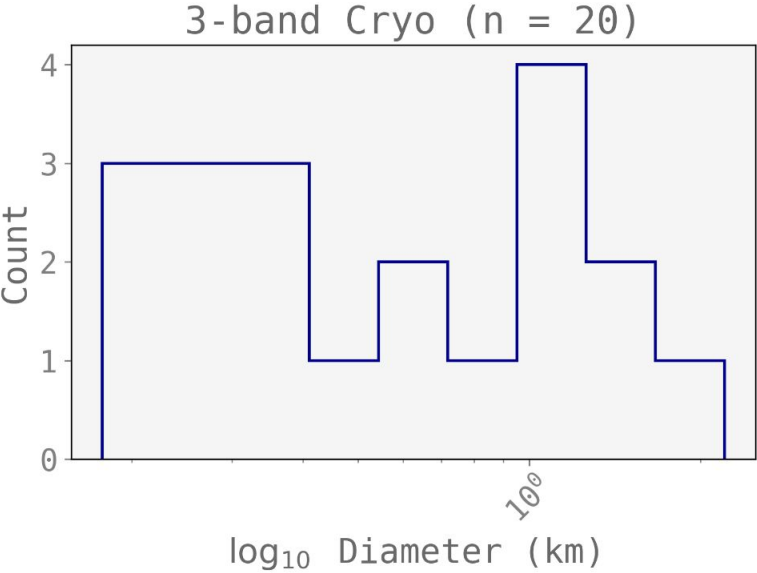
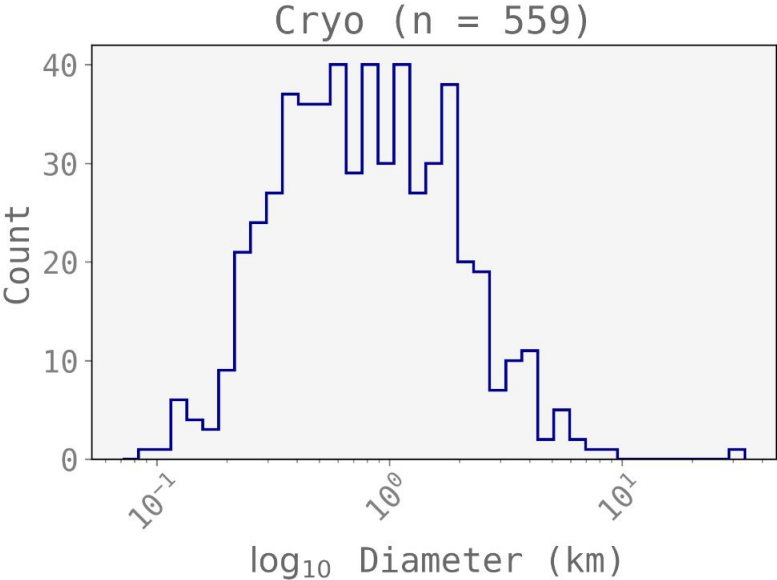
PDF of NEO diameters



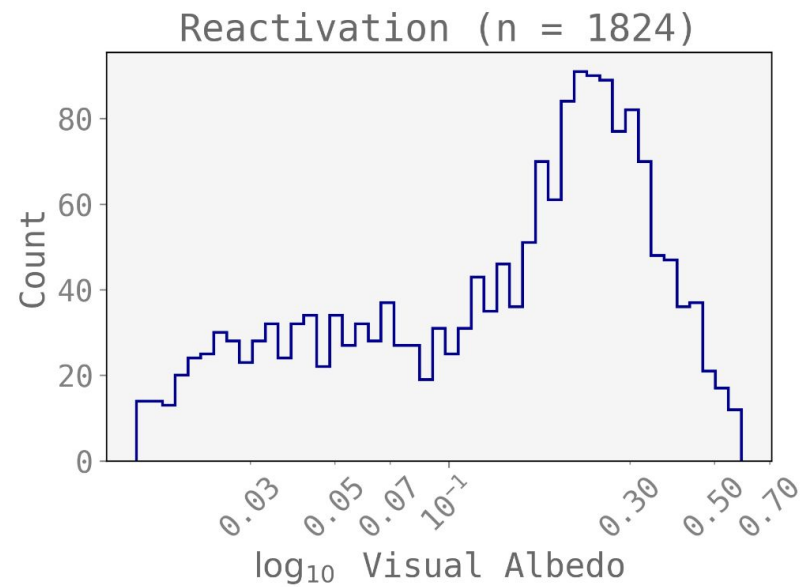
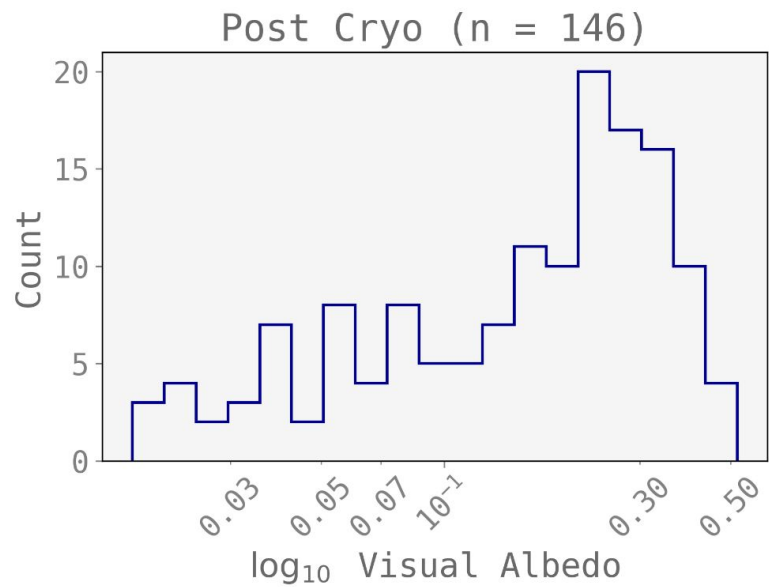
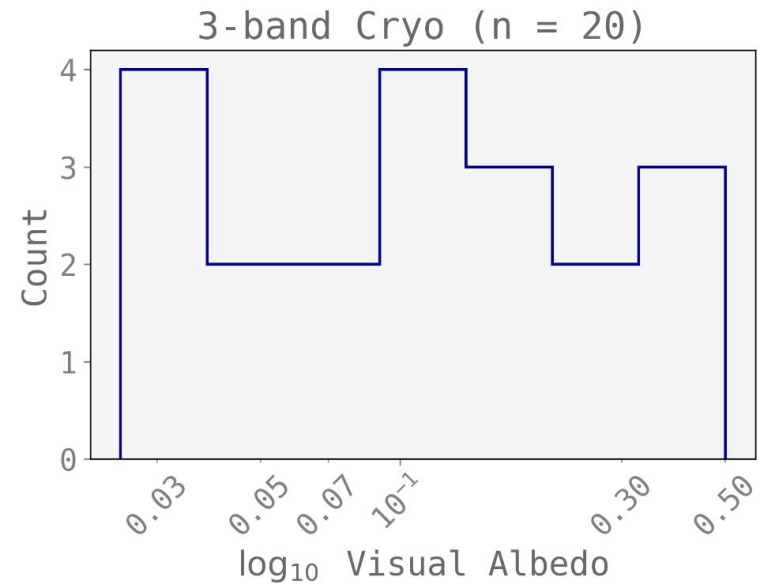
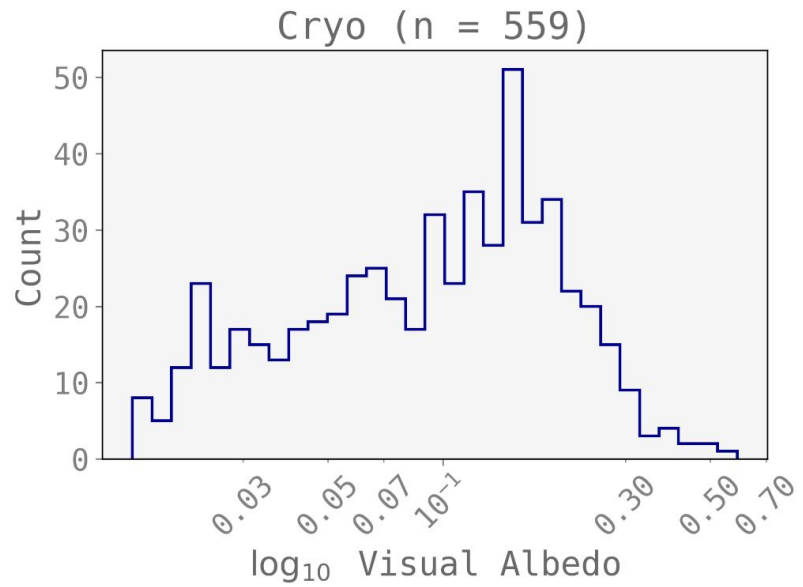
PDF of NEO albedos



# Results



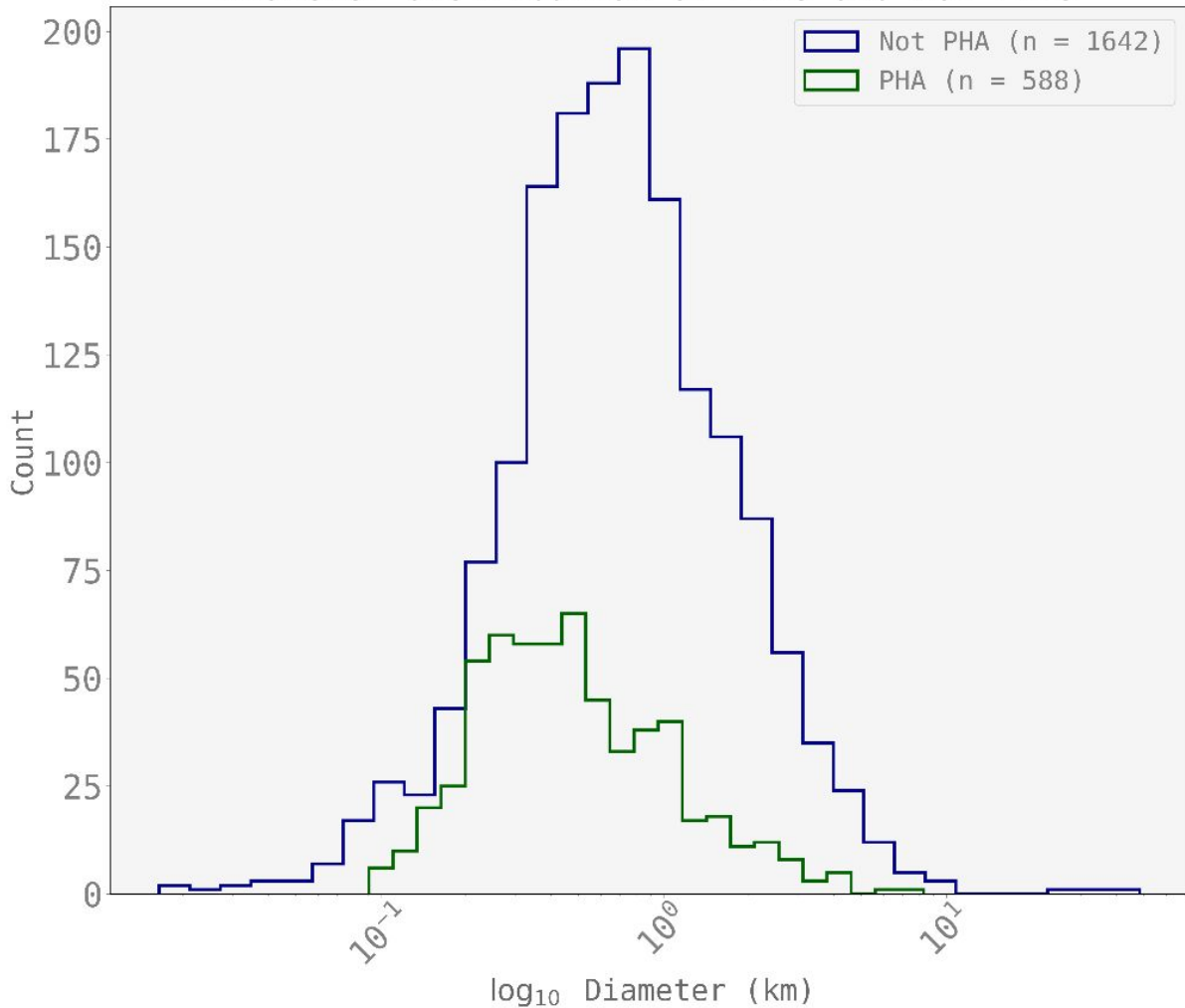
# Results



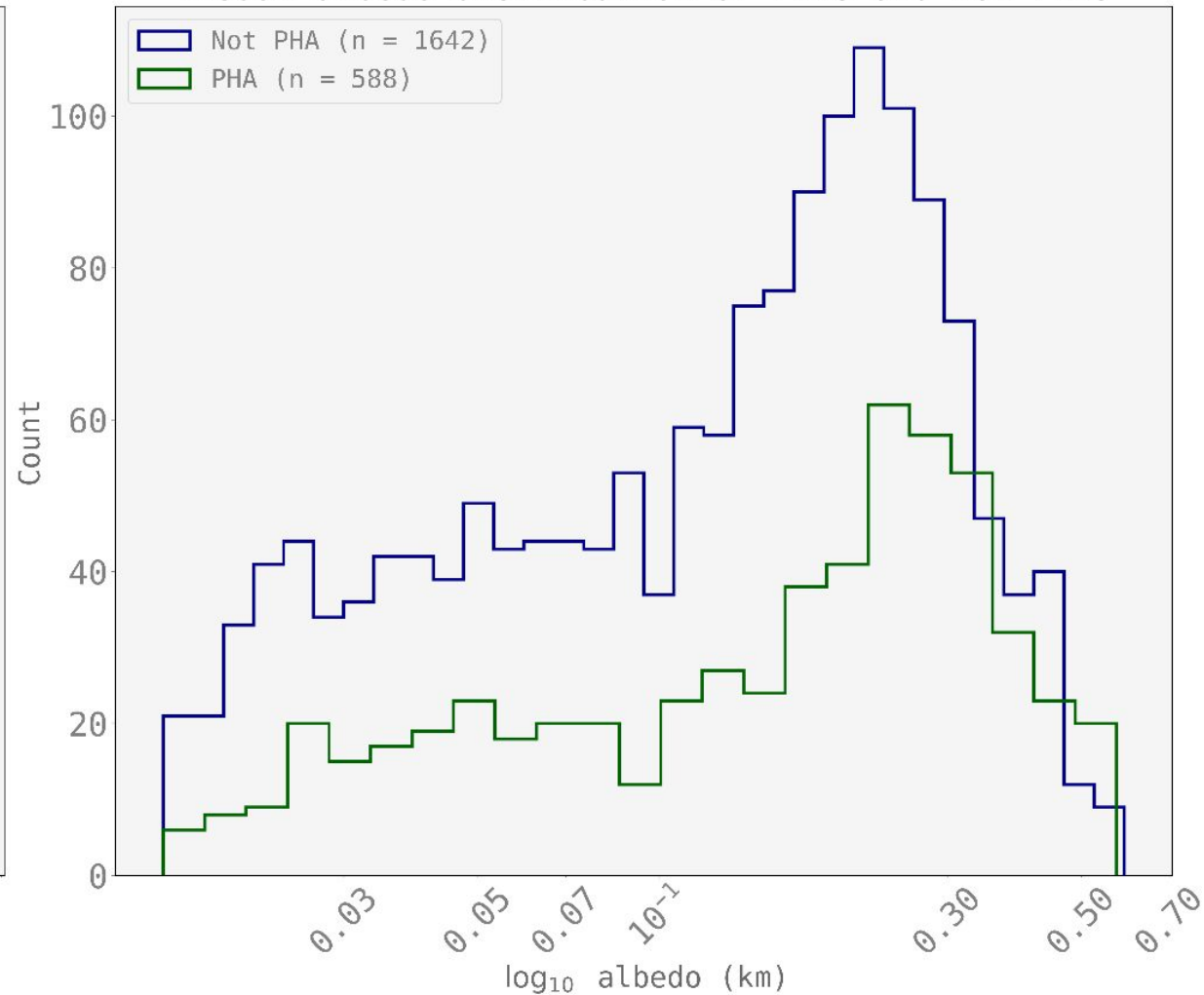
# Results



Diameter distribution of PHAs and not PHAs



Visual albedo distribution of PHAs and not PHAs



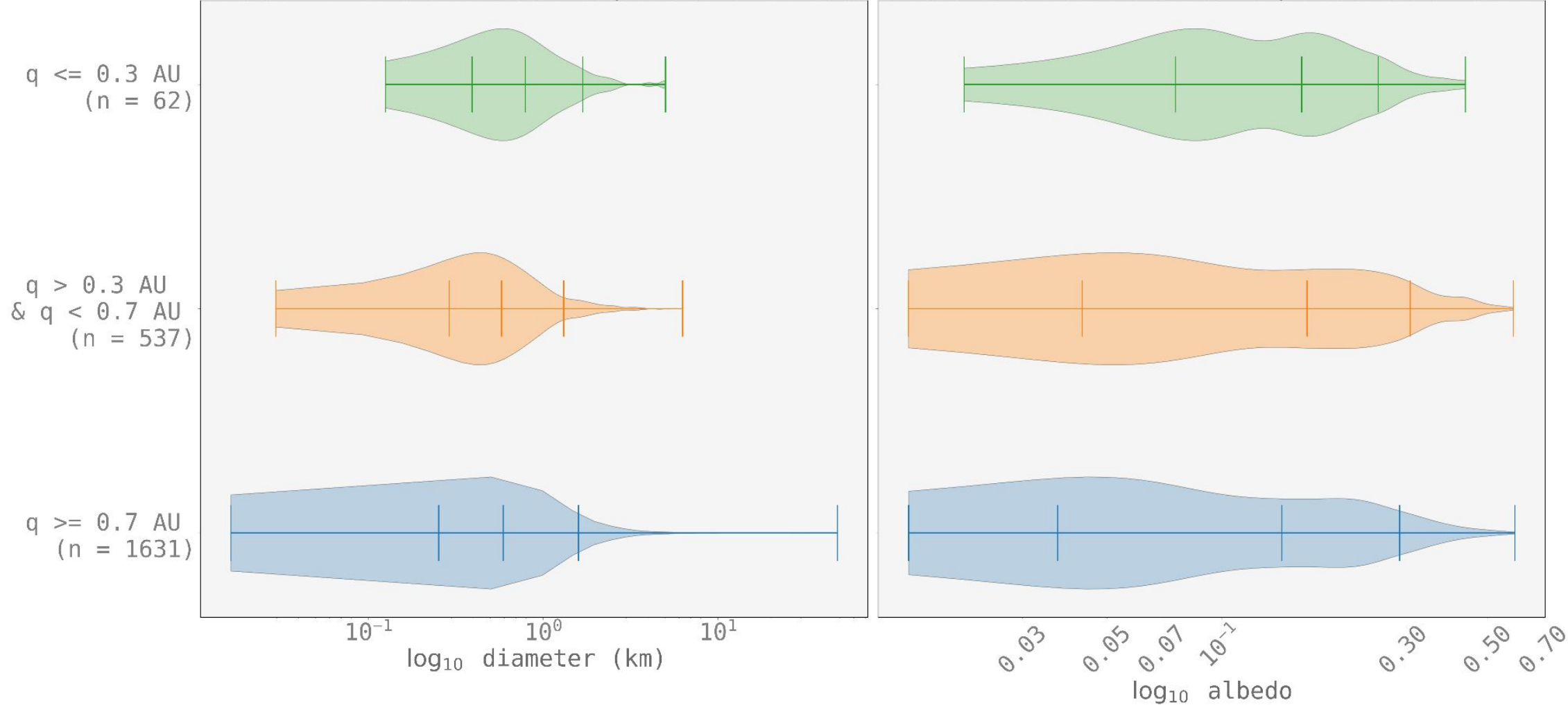


# Results



Diameter distribution based on perihelion distance

Albedo distribution based on perihelion distance

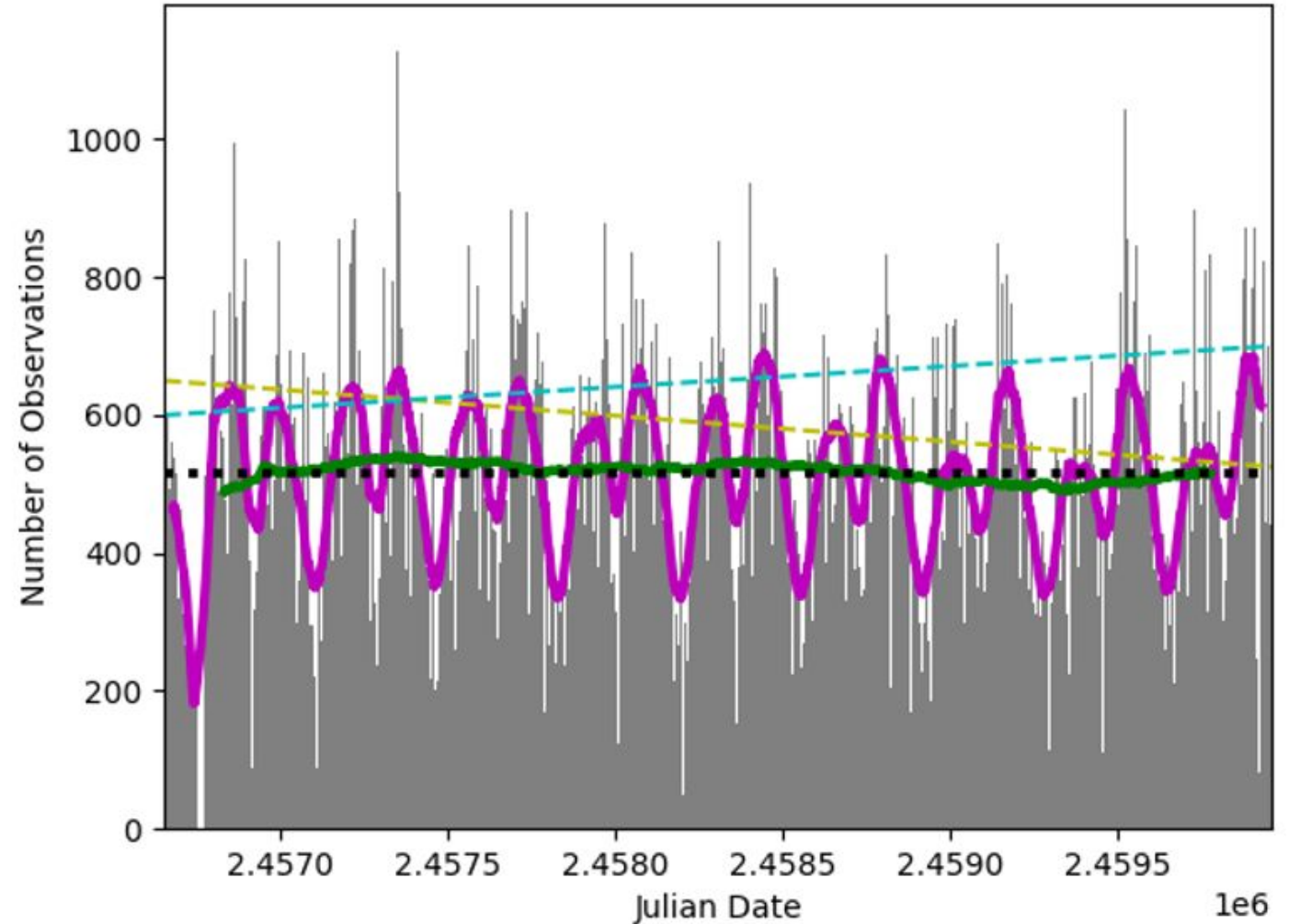


# Future of NEOWISE



- Data quality remains good
- Solar activity is increasing due to the solar cycle → more atmospheric drag
- Re-entry date ~ May, 2025
- Thermal and pointing control will remain within bounds for another year

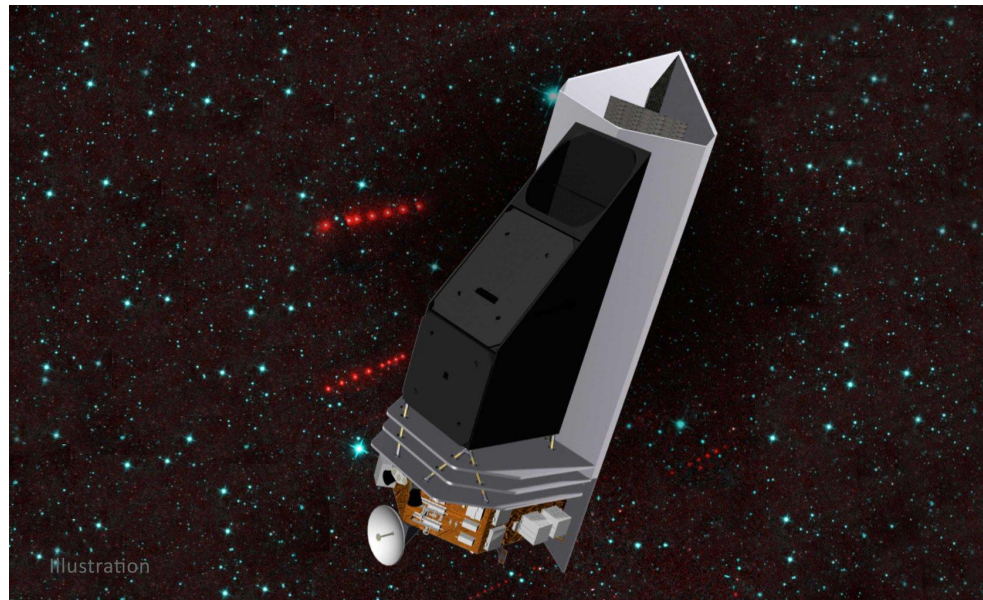
Daily asteroid detection rate



# Current and future prospects



- We aim to refine the NEATM characterized NEOs
  - Refitting fast rotating asteroids
  - Comparison using non-IR constrained data
  - Synthetic population testing
  - For objects with multi-epoch observations → MCMC-based TPM
- Preparing the data analysis pipeline for NEOS





# References

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Mainzer, A., Bauer, J., Cutri, R. M., et al. 2014, *ApJ*, 792, 30, doi: 10.1088/0004-637X/792/1/30

Wright, E. L., Eisenhardt, P. R. M., Mainzer, A. K., et al. 2010, *AJ*, 140, 1868, doi: 10.1088/0004-6256/140/6/1868

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