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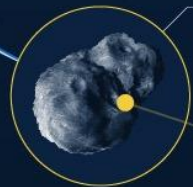
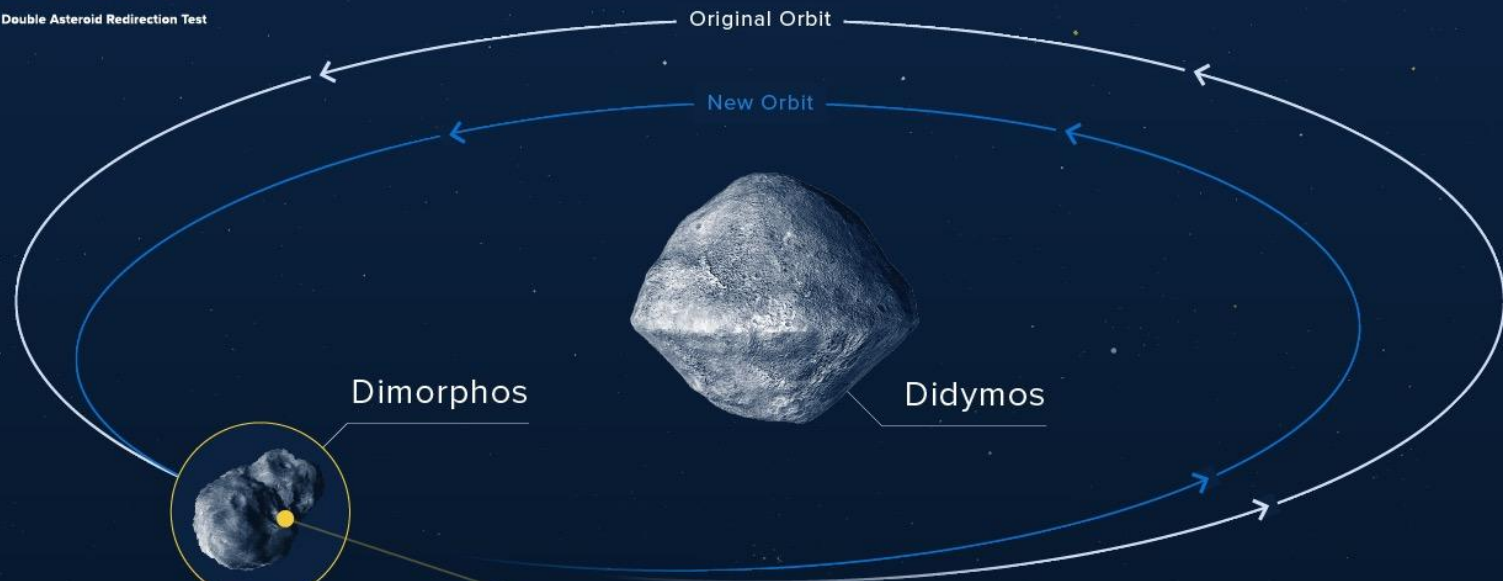
 DART



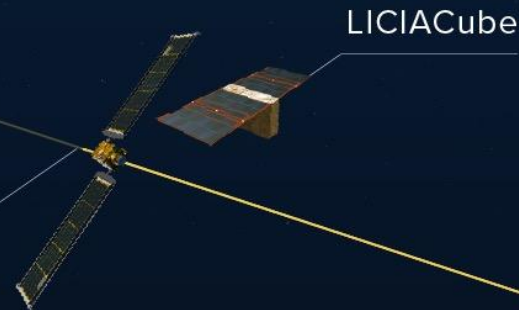
# Deflecting rubble-pile asteroids: Lessons learned from the DART impact on Dimorphos

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**IMPACT**



Earth-based  
observations

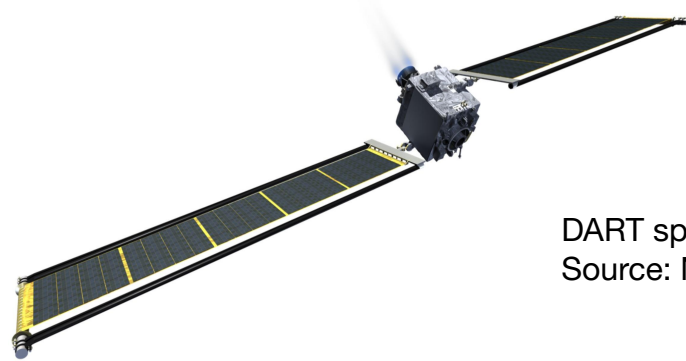
# We didn't know Dimorphos's material properties

## Target properties



- ❑ Cohesive strength – not known
- ❑ Bulk density/porosity – not known
- ❑ Internal structure – not known

## Impact conditions



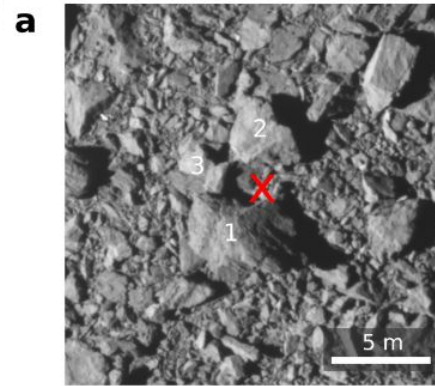
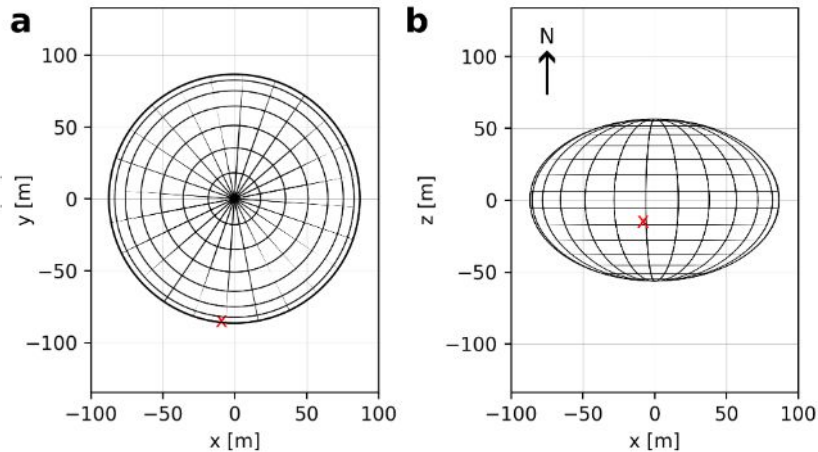
DART spacecraft.  
Source: NASA.

- ❑ Impact velocity – known
- ❑ Impact angle – known
- ❑ Impactor mass/shape – known

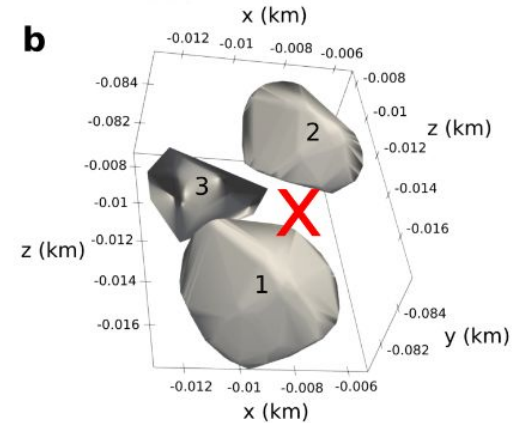
# We used Bern SPH to simulate the DART impact on rubble-pile targets

We used Dimorphos's shape and the impact site characterisation

Projectile: 579.4 kg, at 6.1449 km/s



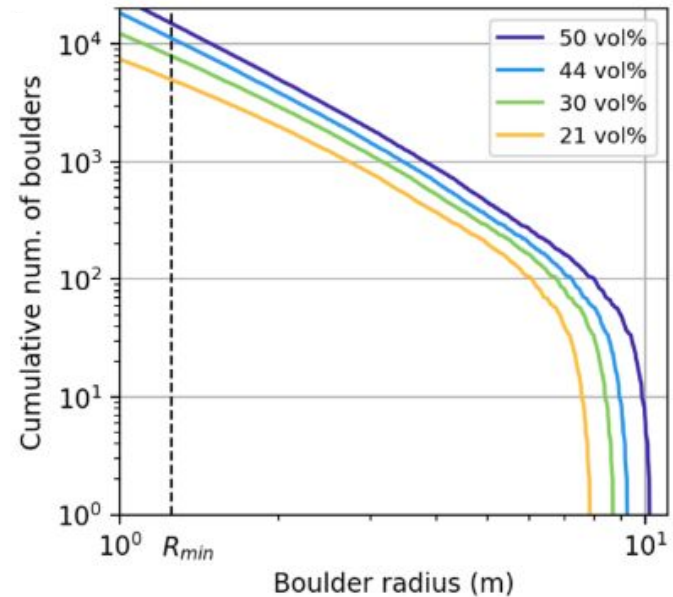
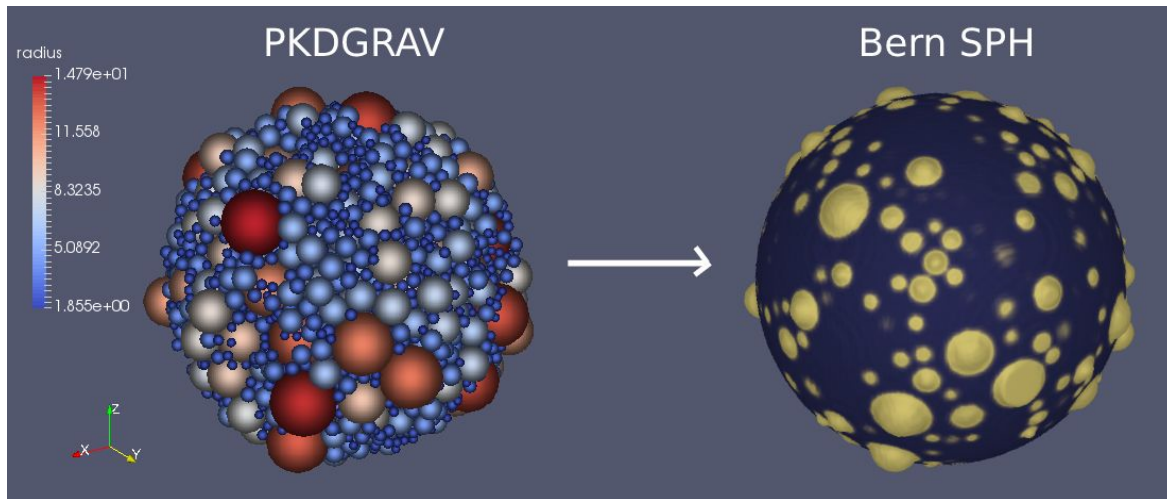
- 1 - Atabaque Saxum
- 2 - Bodhran Saxum
- 3 - Caccavella Saxum



Daly et al., 2023 (Nature)



# We used pkdgrav to generate rubble-pile shapes as inputs for Bern SPH



# We used Bern SPH to simulate the DART impact

Boulder tensile strength,  $Y_T = 10$  MPa

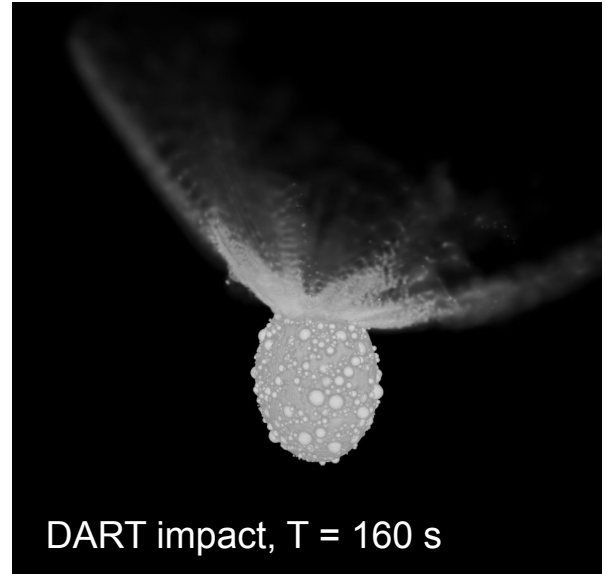
Boulder porosity, 10%

We systematically varied target material properties:

Boulder volume fraction, 0 – 50 vol%

Matrix cohesion,  $Y_0 = 0 - 50$  Pa

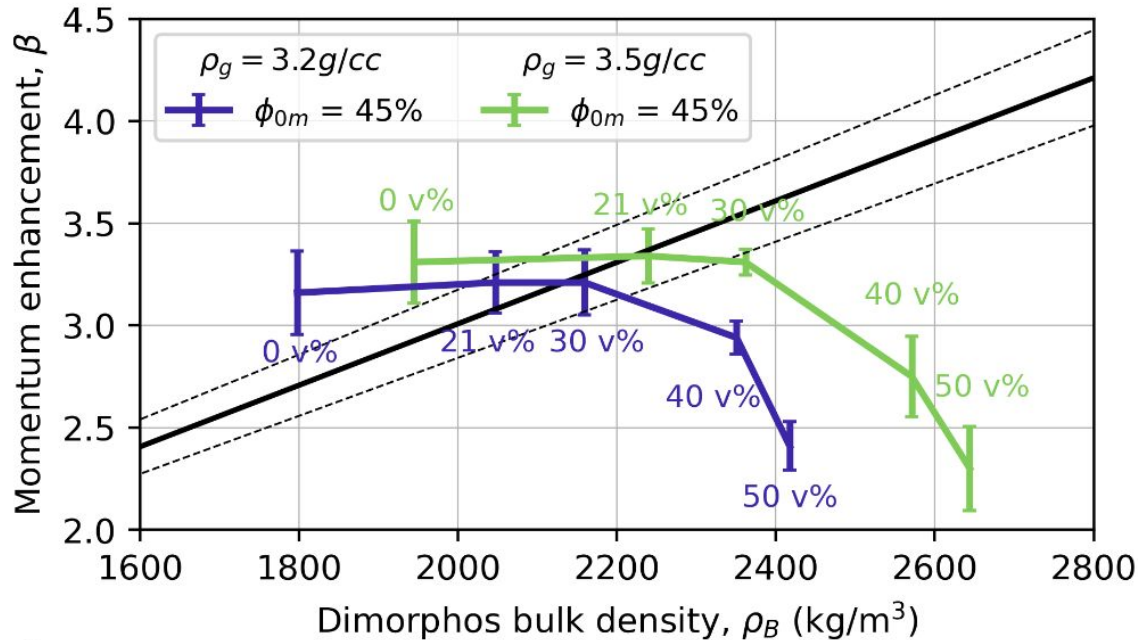
Matrix porosity, 45%



We track the impact outcome to up to 2 hours after the impact and compare the outcome with the observations.

# Finding 1: The surface boulder packing is low (<40 vol%)

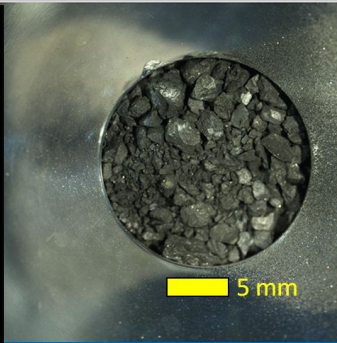
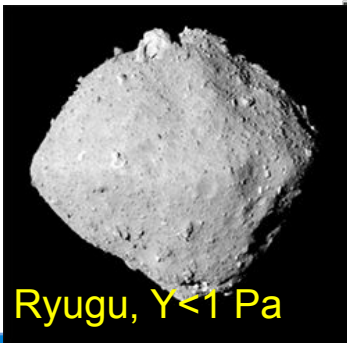
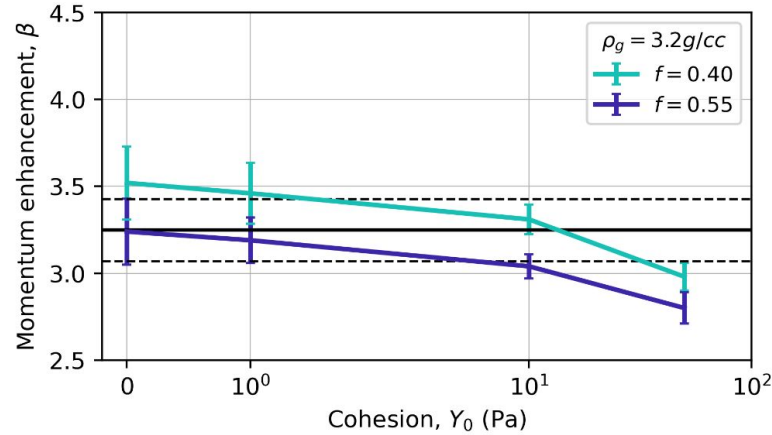
(of boulders larger than 2.5 m)



Constant volume  $\sim 0.00181 \text{ km}^3$

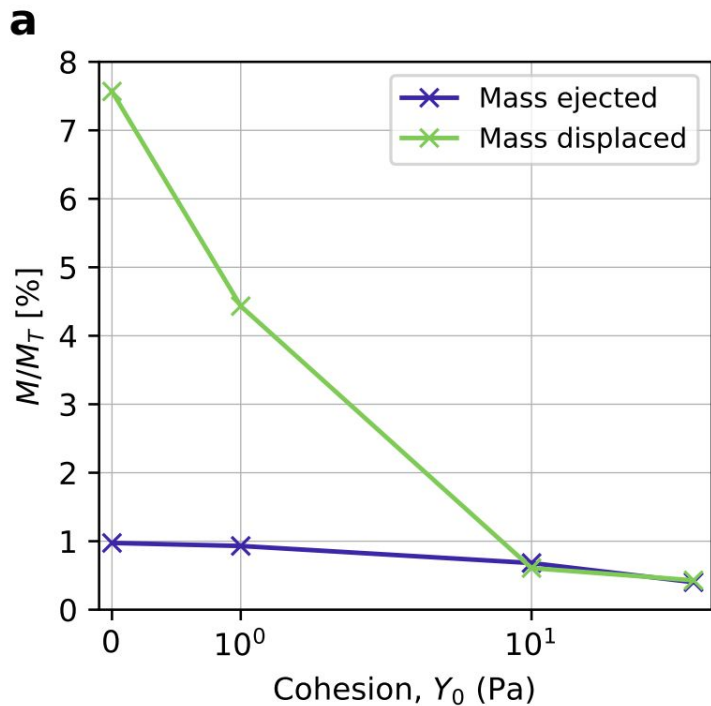


# Finding 2: The surface/shallow sub-surface cohesion is less than 50 Pa. Best fit, $Y < 1\text{ Pa}$





# Finding 3: More than 0.3% and up to 1% of Dimorphos's mass was ejected – consistent with early observations



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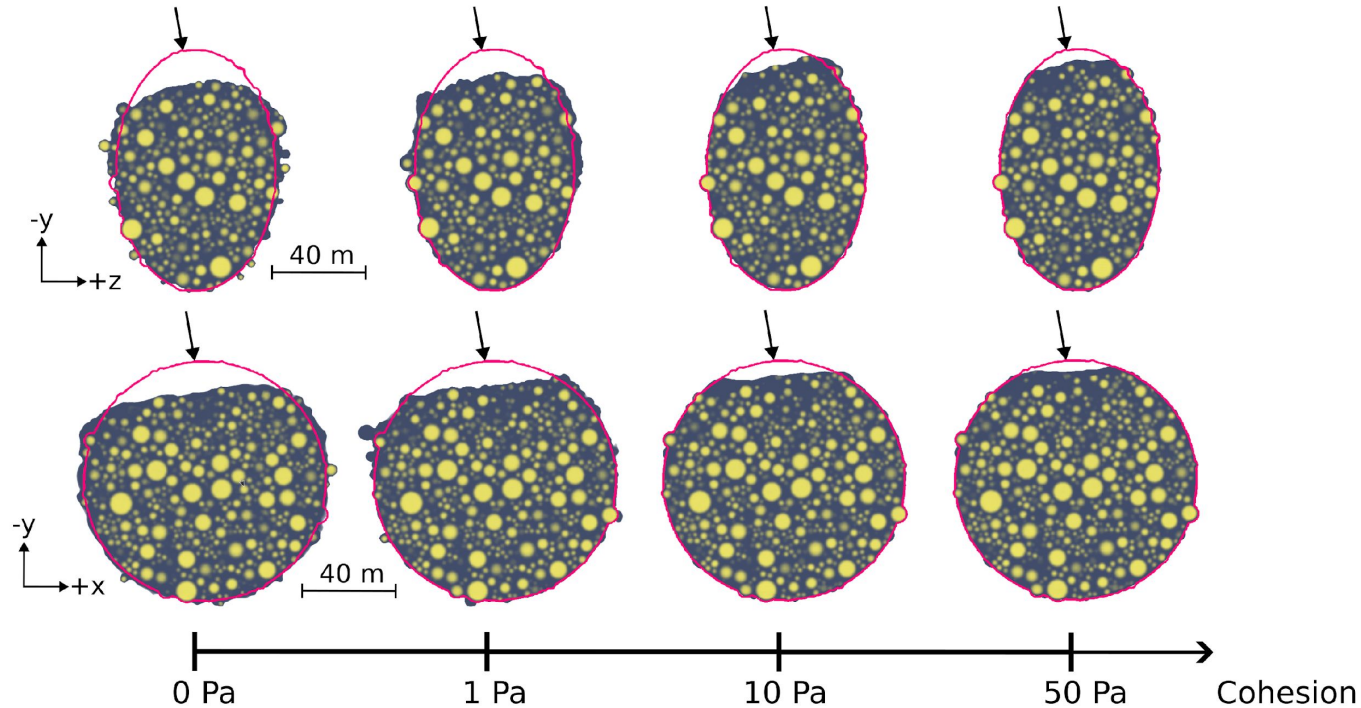
Article | [Published: 01 March 2023](#)

## Light Curves and Colors of the Ejecta from Dimorphos after the DART Impact

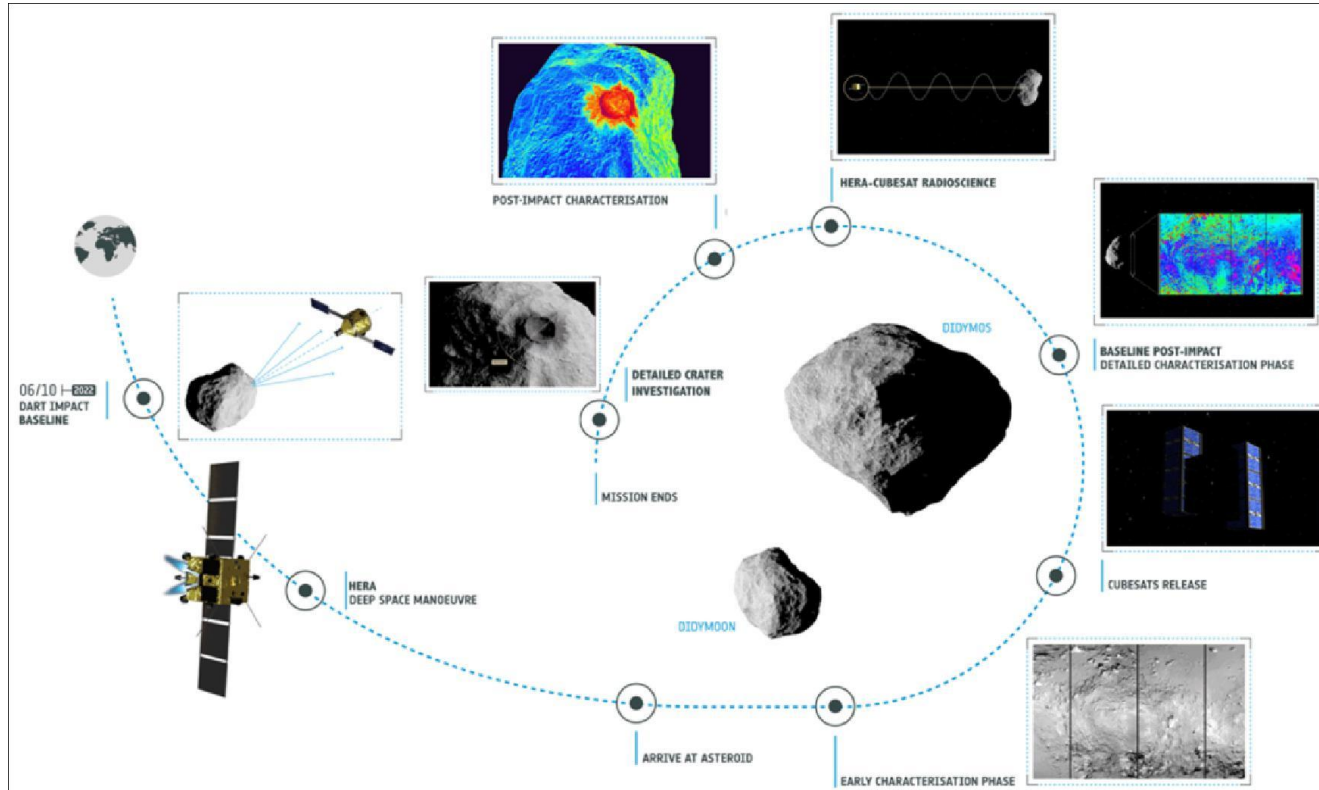
[Ariel Graykowski](#) [✉](#), [Ryan A. Lambert](#), [Franck Marchis](#), [Dorian Cazeneuve](#), [Paul A. Dalba](#), [Thomas M. Esposito](#), [Daniel O'Conner Peluso](#), [Lauren A. Sgro](#), [Guillaume Bclairard](#), [Antonin Borot](#), [Arnaud Malvache](#), [Laurent Marfisi](#), [Tyler M. Powell](#), [Patrice Huet](#), [Matthieu Limagne](#), [Bruno Payet](#), [Colin Clarke](#), [Susan Murabana](#), [Daniel Chu Owen](#), [Ronald Wasilwa](#), [Keiichi Fukui](#), [Tateki Goto](#), [Bruno Guillet](#), [Patrick Huth](#), ...  
[Ian M. Transom](#) [+ Show authors](#)



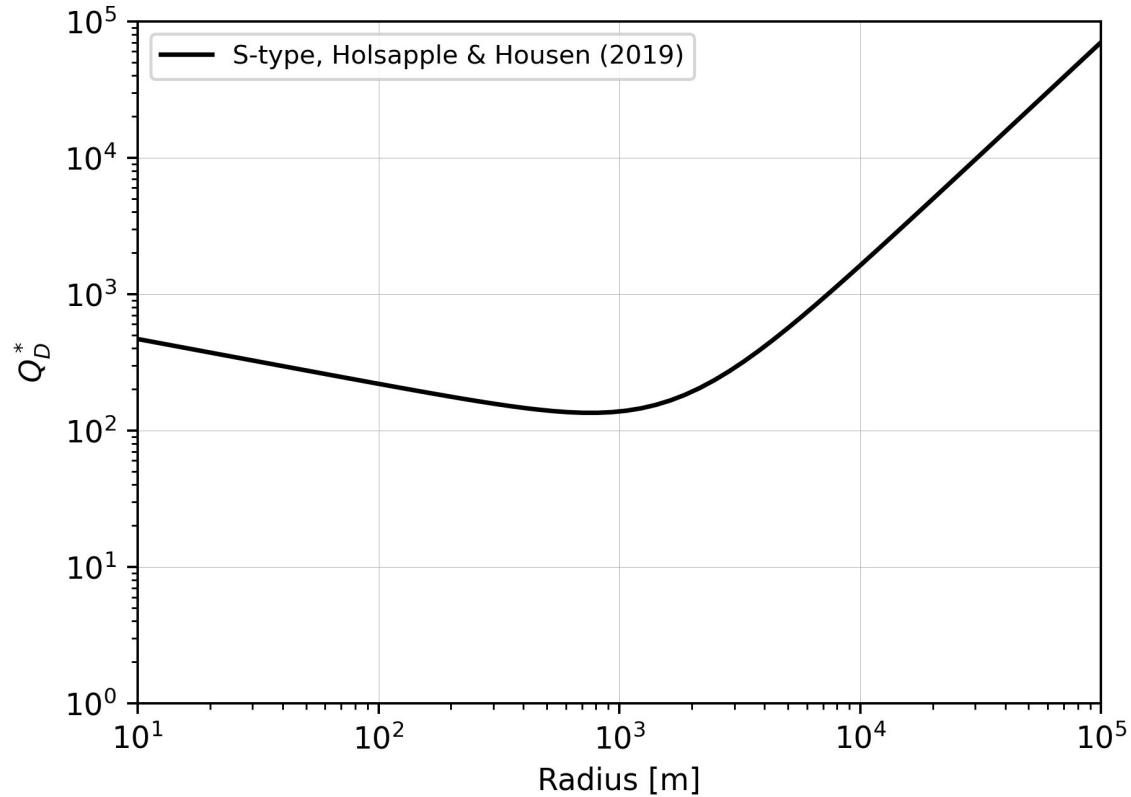
# DART impact occurred in the sub-catastrophic regime and may have caused the global deformation of Dimorphos



# Our models provide testable predictions that will be studied by ESA's Hera mission

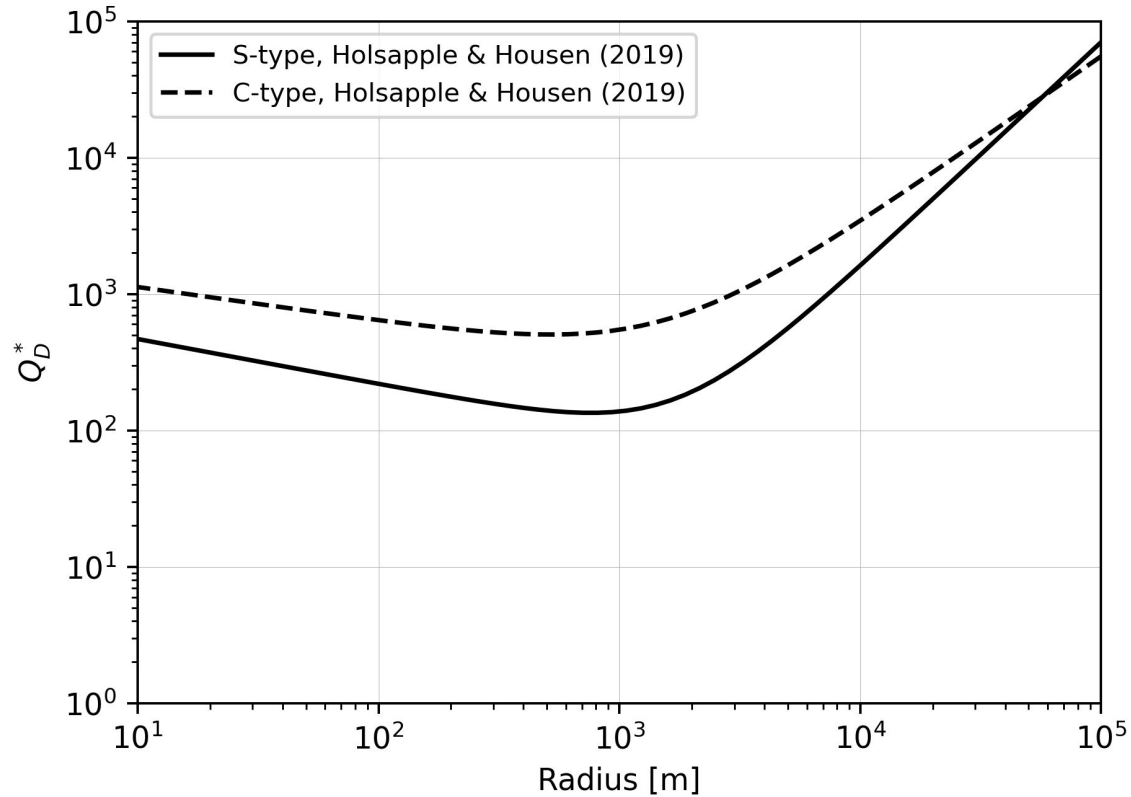


# Catastrophic catastrophic disruption threshold for asteroids



-  $Q_D^*$  curve for rocky S-type asteroids, at 5.5 km/s and  $45^\circ$

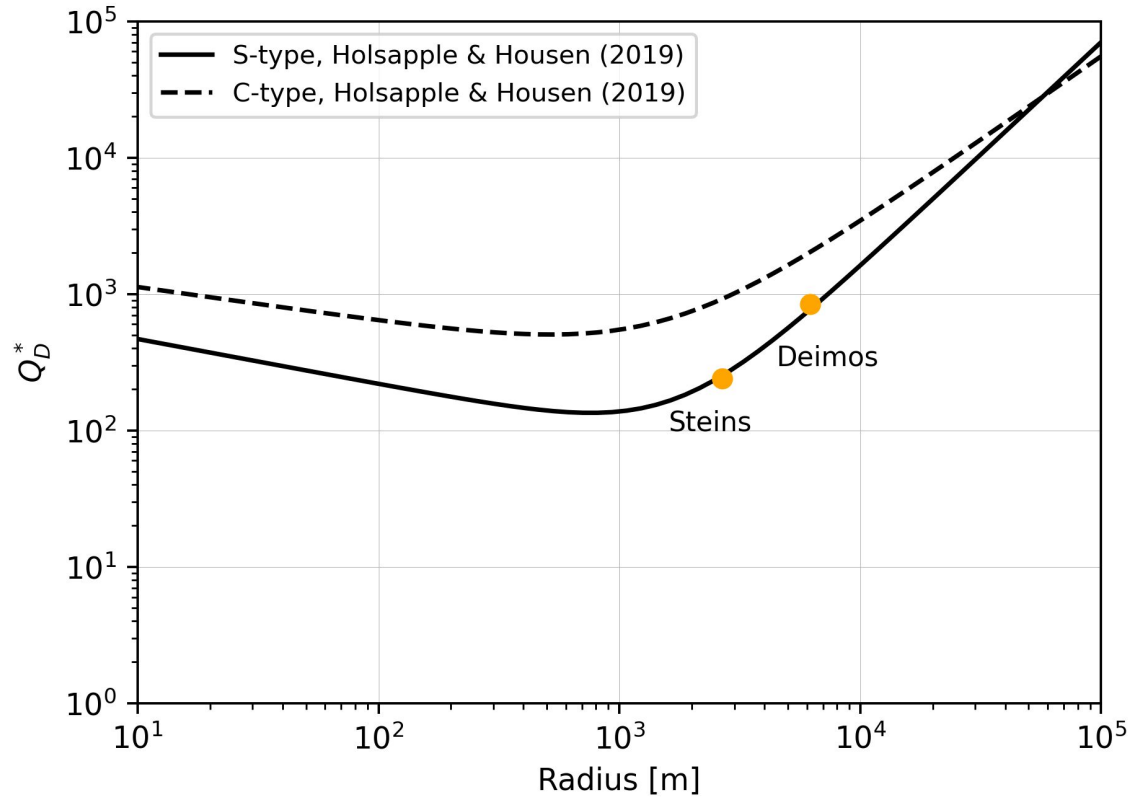
# Catastrophic catastrophic disruption threshold for asteroids



- $Q_D^*$  curve for rocky S-type asteroids, at 5.5 km/s and  $45^\circ$
- $Q_D^*$  curve for porous C-type asteroids, at 5.5 km/s and  $45^\circ$

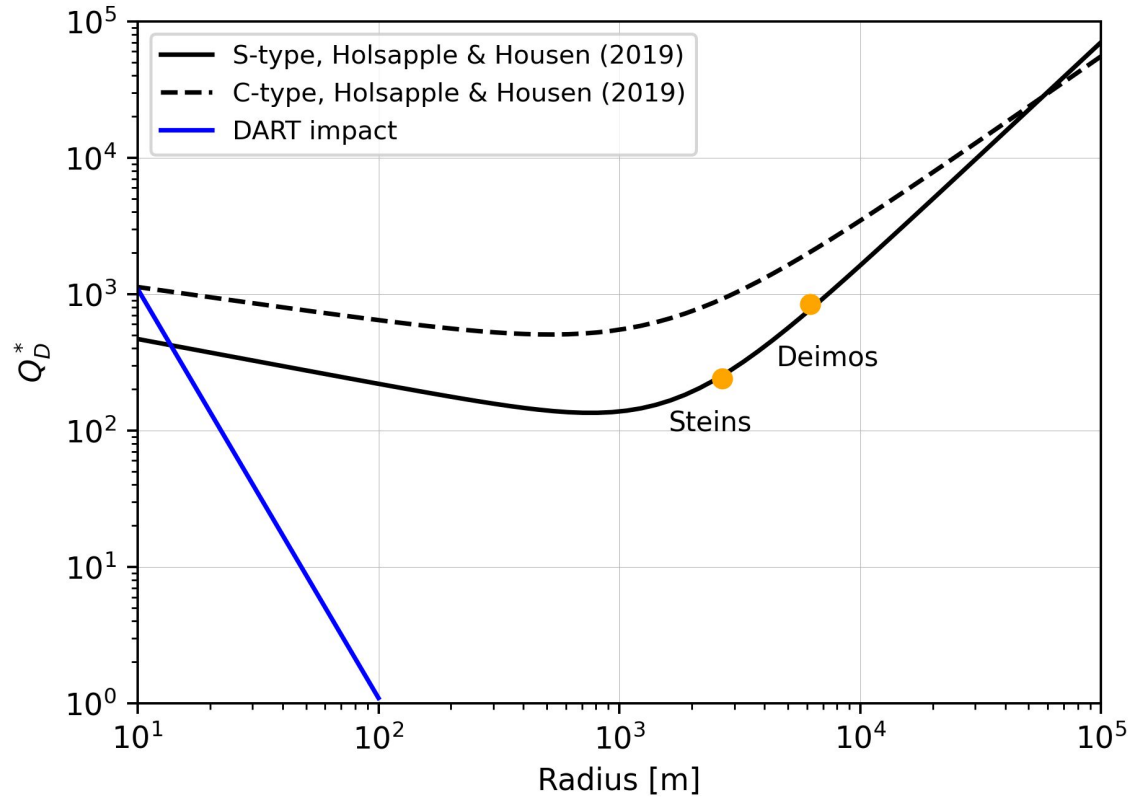


# Catastrophic catastrophic disruption threshold for asteroids



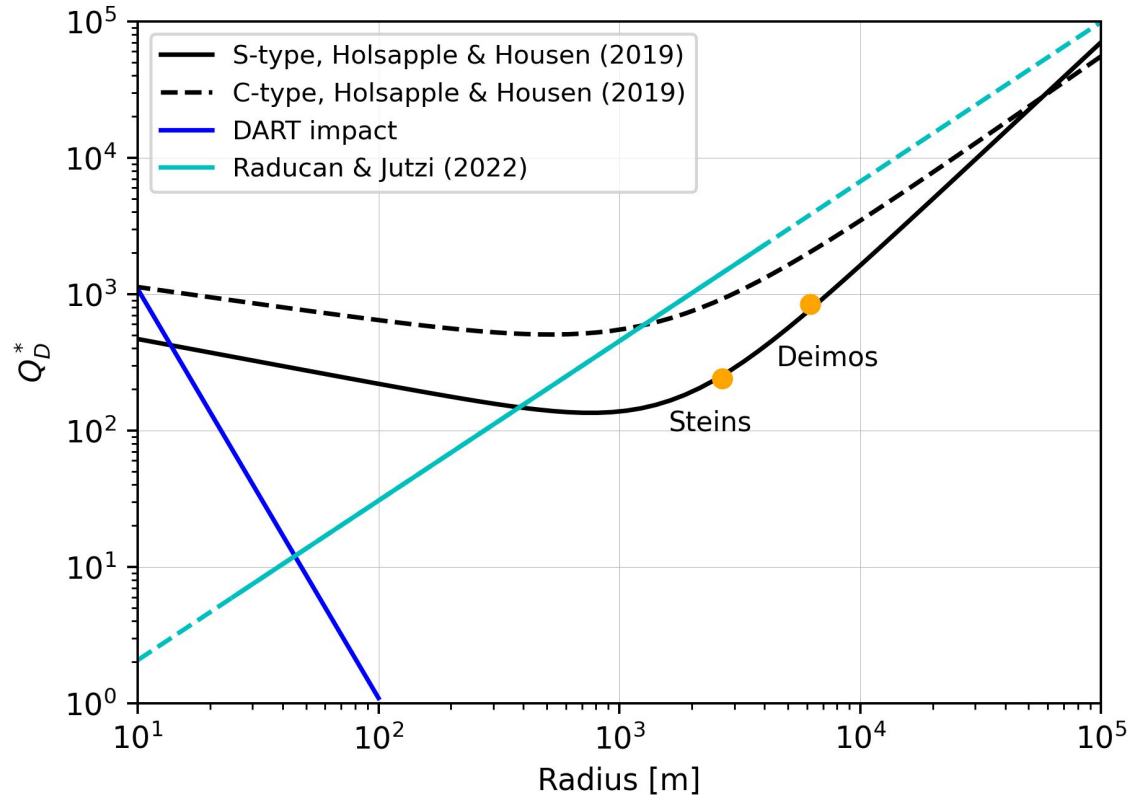
- $Q_D^*$  curve for rocky S-type asteroids, at 5.5 km/s and  $45^\circ$
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# Catastrophic catastrophic disruption threshold for asteroids



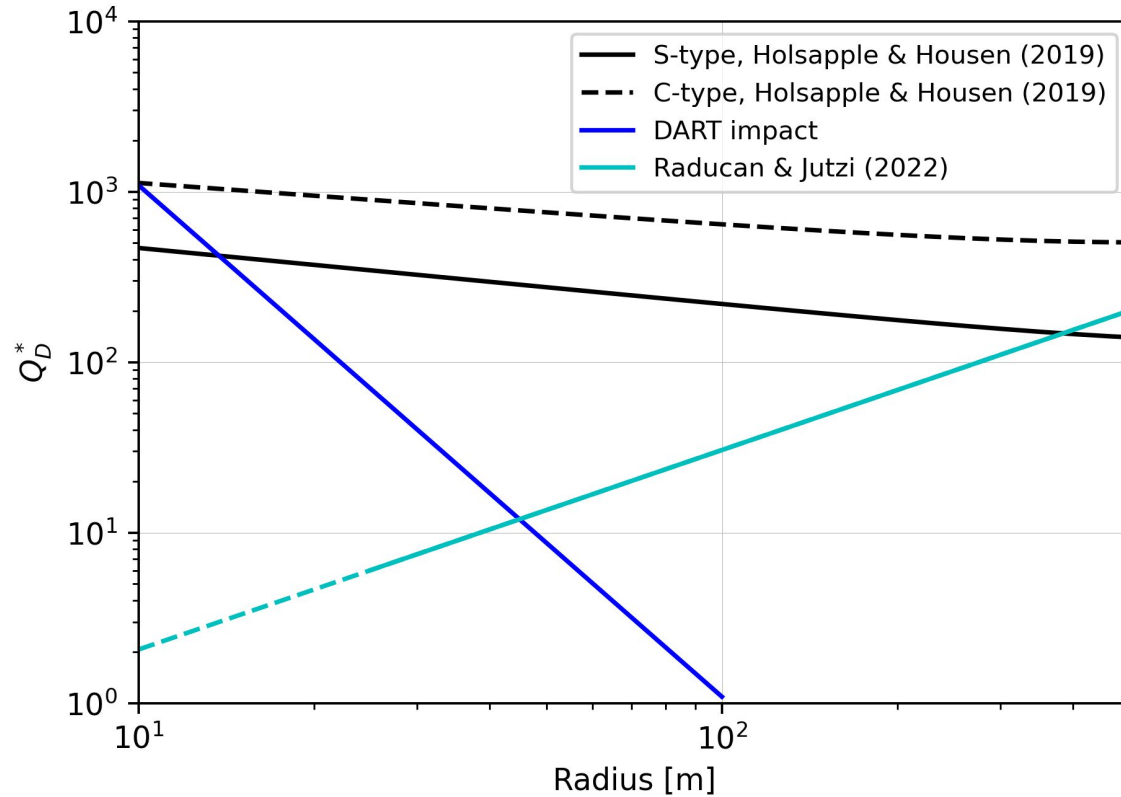
- $Q_D^*$  curve for rocky S-type asteroids, at 5.5 km/s and  $45^\circ$
- $Q_D^*$  curve for porous C-type asteroids, at 5.5 km/s and  $45^\circ$
- Specific impact energy for DART

# Catastrophic catastrophic disruption threshold for asteroids

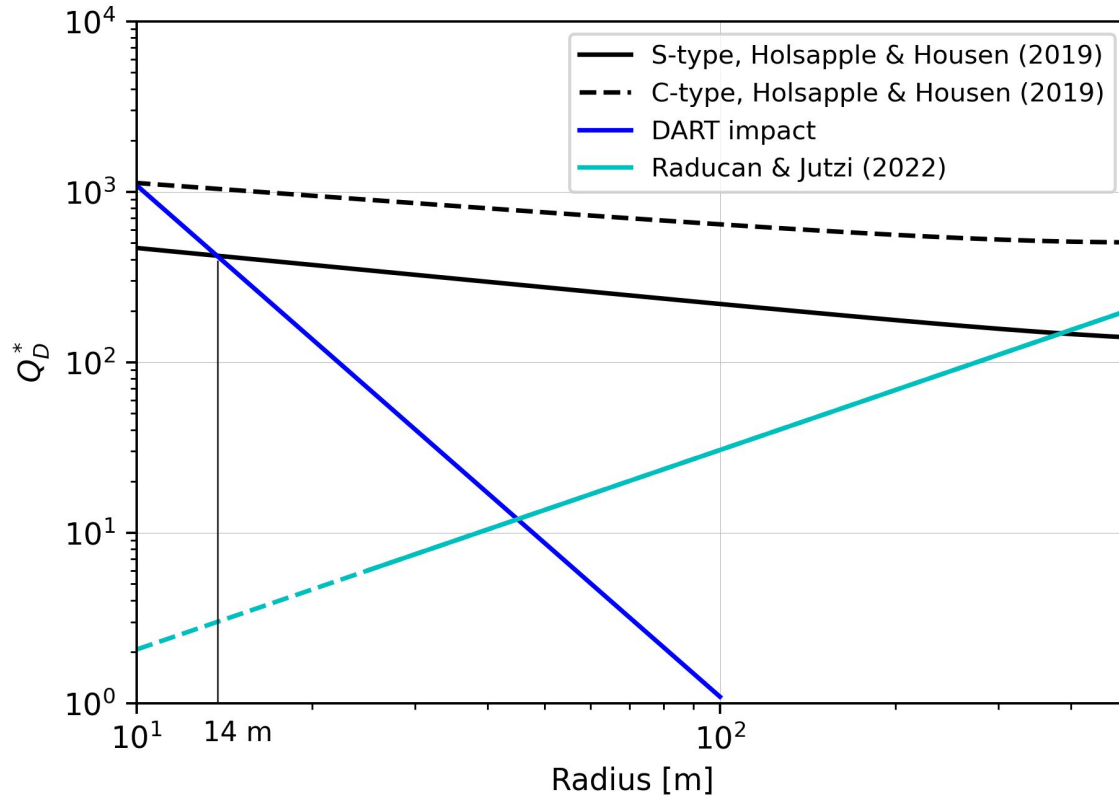


- $Q_D^*$  curve for rocky S-type asteroids, at 5.5 km/s and  $45^\circ$
- $Q_D^*$  curve for porous C-type asteroids, at 5.5 km/s and  $45^\circ$
- Specific impact energy for DART
- $Q_D^*$  curve for homogeneous, cohesionless asteroids

# Catastrophic catastrophic disruption threshold for asteroids



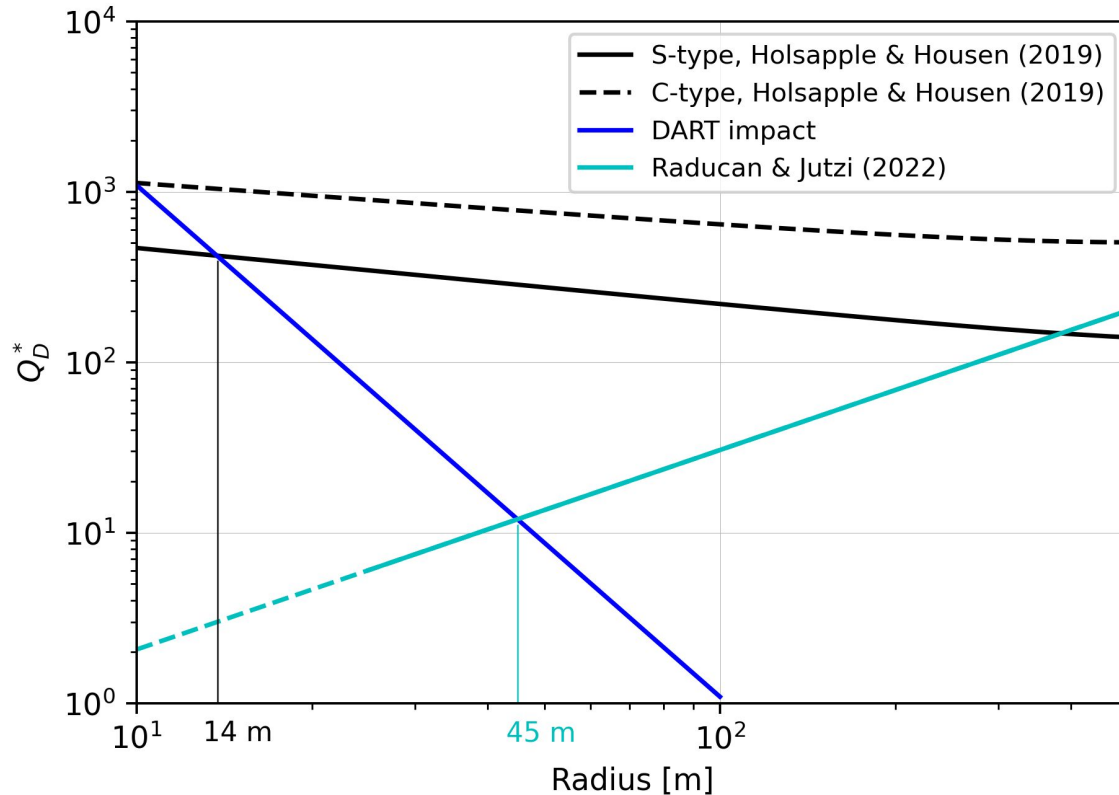
# Catastrophic catastrophic disruption threshold for asteroids



- A DART-like spacecraft would disrupt a <28 m (diameter) S-type monolith

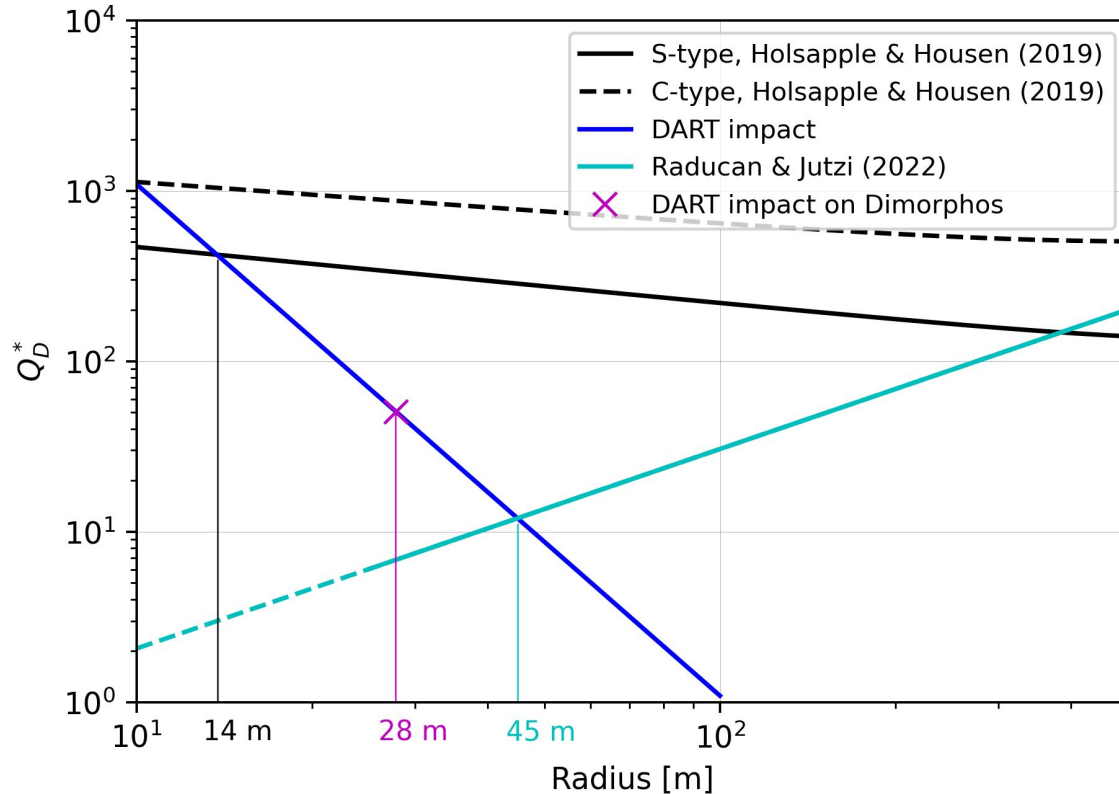


# Catastrophic catastrophic disruption threshold for asteroids



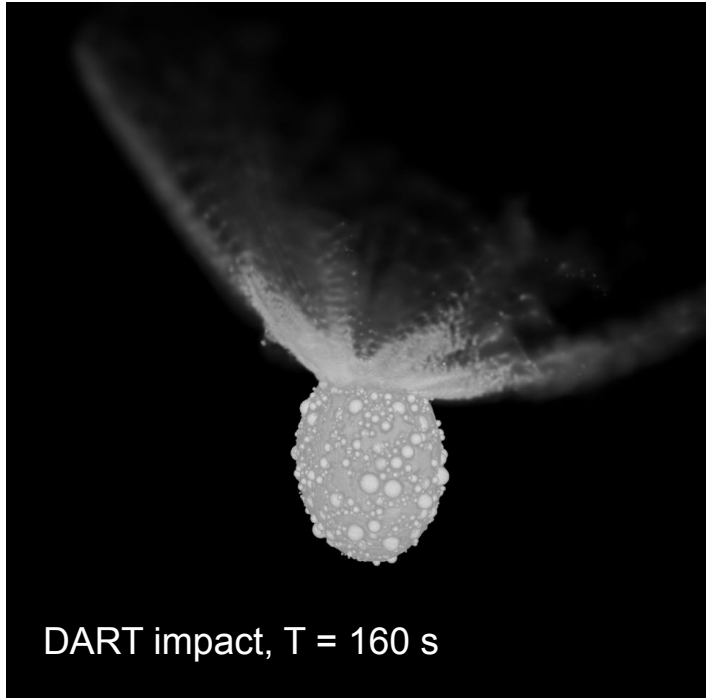
- A DART-like spacecraft would disrupt a <28 m (diameter) S-type monolith
- A DART-like spacecraft would disrupt a <90 m cohesionless, homogeneous target

# A DART spacecraft would catastrophically disrupt a rubble-pile asteroid smaller than ~56 m across

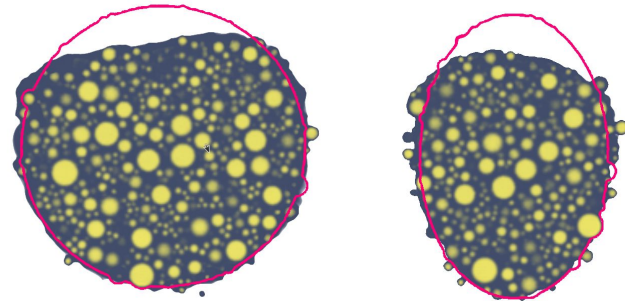


- A DART-like spacecraft would disrupt a 28 m (diameter) S-type monolith
- A DART-like spacecraft would disrupt a <90 m cohesionless, homogeneous target
- A DART-like spacecraft would disrupt a <56 m Dimorphos

# Conclusions



- Dimorphos is a rubble-pile, with a low surface cohesion and a low boulder packing (less than ~40%)
- DART likely caused the deformation and resurfacing of Dimorphos, as opposed to a well defined impact crater
- Our models provide testable predictions that will be studied by ESA's Hera mission in late 2026.
- **DO NOT** send a DART-sized spacecraft to an asteroid smaller than ~56 m across



Dimorphos,  
T = 2 hours