The ESA Hera mission: planetary defense and science return

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hera

(O) OBSERVATOIRE

DE LA CÔTE D'AZUR

UNIVERSITÉ CÔTE D'AZUR

Patrick Michel

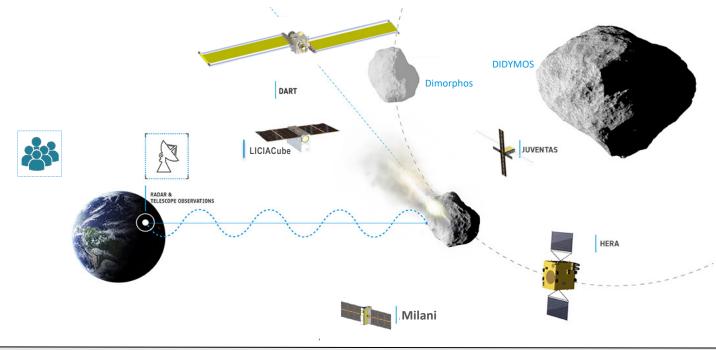
Université Côte d'Azur Observatoire de la Côte d'Azur CNRS, Lagrange Lab, Nice, France

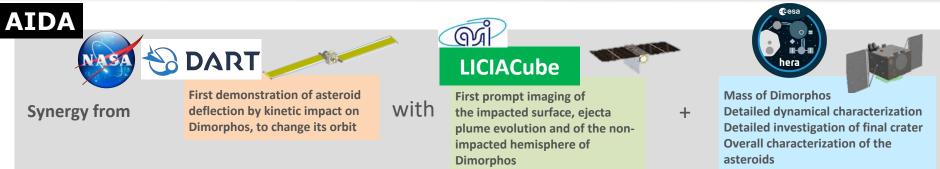
On behalf of the Hera Science Team

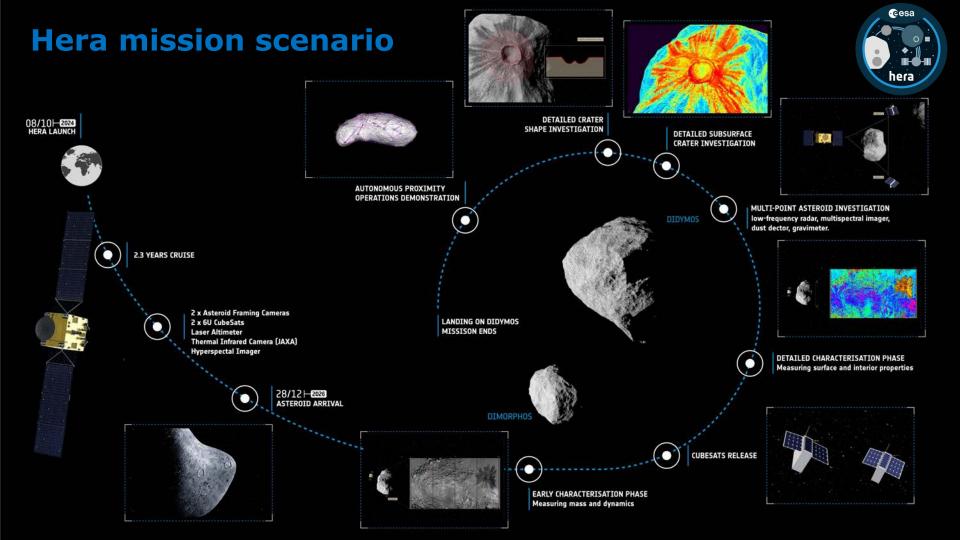
CNrs

7th IAA Planetary Defense Conference

AIDA international collaboration



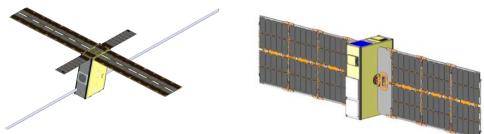


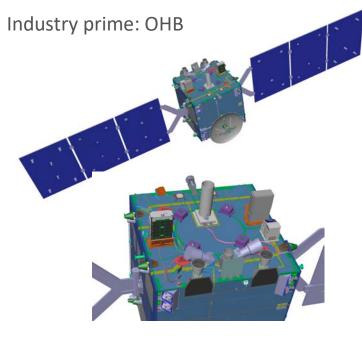


Hera spacecraft, payloads and 2 Cubesats

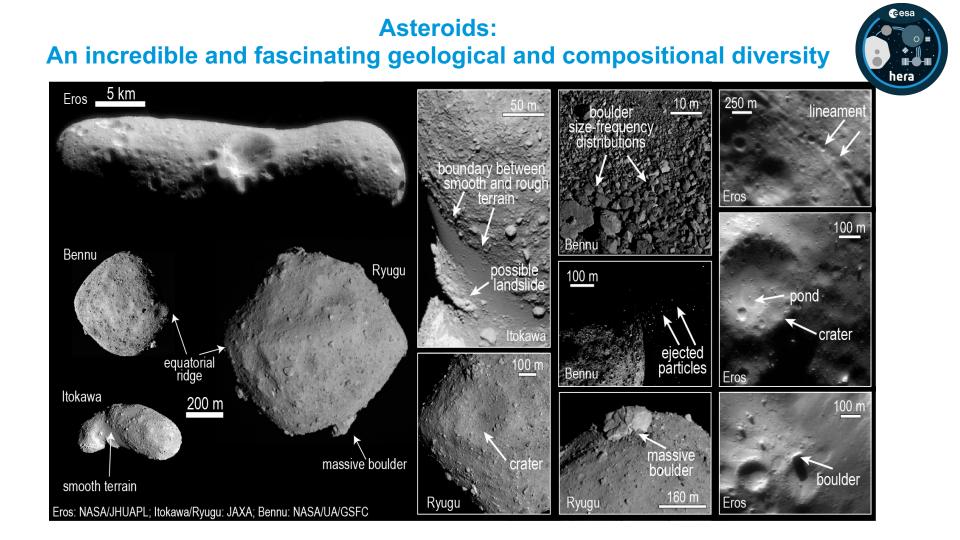


- Asteroid Framing Camera
- LIDAR (PALT)
- Thermal Infrared Imager (TIRI)
- Hyperspectral Imager (Hyperscout-H)
- Juventas Cubesat (internal structure, gravity field, surface properties)
- **Milani Cubesat** (Mineralogy, space weathering, dust detection, gravity field)

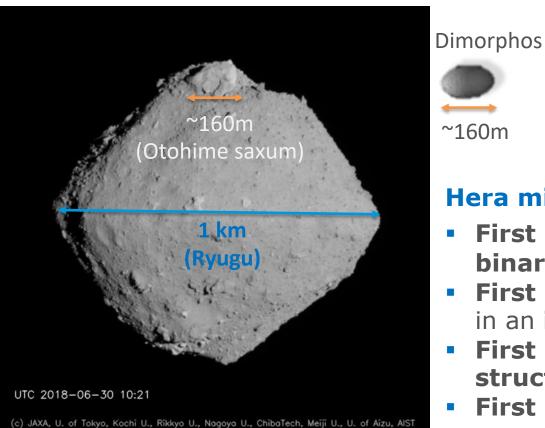




+ ISL and RadioScience



Dimorphos vs Ryugu (Hayabusa2)





~800m Hera mission firsts

~160m

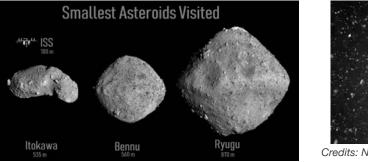
- First rendezvous mission to a binary asteroid
- **First** characterization of an asteroid in an **intriguing size range**
- First radar sounding (internal structure) of an asteroid
- First full scale cratering physics **experiment** (with DART)

Hera in the context of Hayabusa2 and OSIRIS-REx

 First detailed characterization and surface response of 2 primitive asteroids in different gravity conditions



- •Ryugu is 900 meter wide
- •Bennu is 500 meter wide

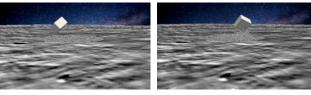




Credits: NASA/Goddard/University of Arizona

- Dimorphos is ~ 3 times less wide than Bennu
 - Another step in low gravity levels

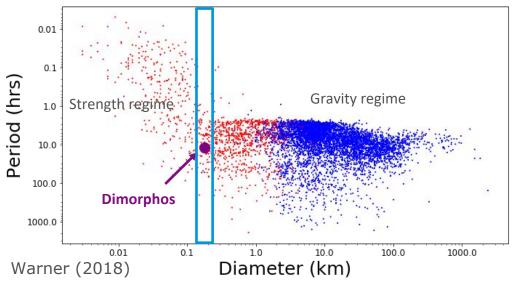
 Possibility to understand how some processes scale with gravity down to the low-g of Dimorphos



Dimorphos: in a very interesting size range



First detailed characterisation (**including interior**) of an asteroid at the transition between strength (cohesion) and gravity-dominated bodies



New knowledge in our understanding of asteroid geophysics and impact response (with DART)



Didymos and the prevalence of « spinning tops »





Credit: JAXA, University of Tokyo, et al.

Single Asteroid Bennu Howell et al. 2008, ACM Solog ACM Solog ACM Bennu Brosovic et al. 2011 Single Asteroid Binary Asteroid 1999 KW4 0399 KW4 0399 KW4 0390 KV5 2011 Becker et al. 2005 Date of al. 2005 Busch et al. 2005 Date of al. 2005 Busch et al. 2005 Date of al. 2005 Date o

Credit: NASA

- Ryugu and Bennu are top shapes possibly formed early by disruption/reaccumulation !!! (Michel, P. & Ballouz R.-L. et al. 2020)
- But both Ryugu and Bennu are single!
- What does a top shape binary primary look like?

First detailed investigation of a binary asteroid

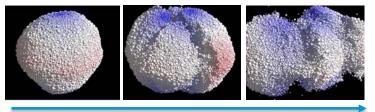
15% of NEAs are binaries

- YORP mass shedding?
- Are Didymos and Dimorphos rubble piles?

Key knowledge to understand Solar System accretion processes

- First asteroid visited near disruption limit
- Insights on formation or disruption of planets





Spin-up Zhang et al. 2021, Icarus 362



1998 QE2



2004 BL86



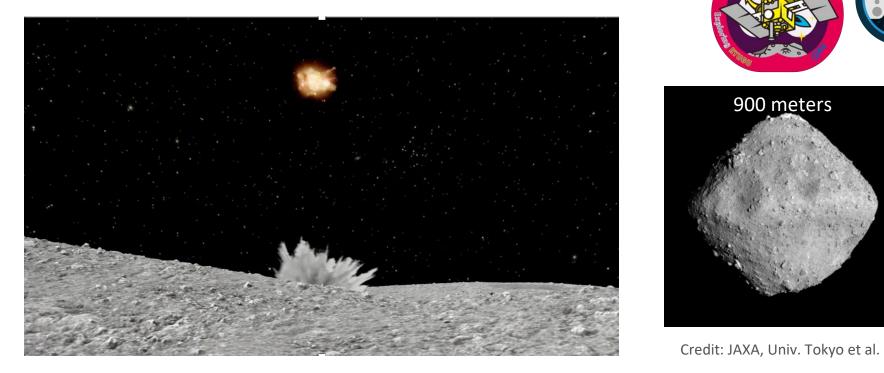
2014 JO25





Hera will measure the outcome of a cratering impact at actual asteroid scale

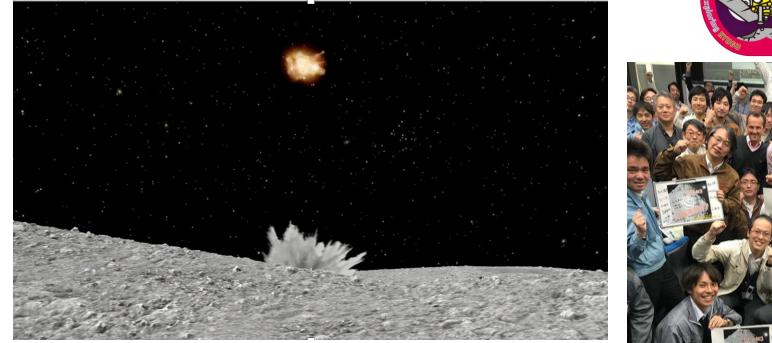
Cratering physics: Hayabusa2 versus Hera



Hayabusa2 Small Carry-on Impactor April 5, 2019

Arakawa et al. 2020, Science

Cratering physics: Hayabusa2 versus Hera



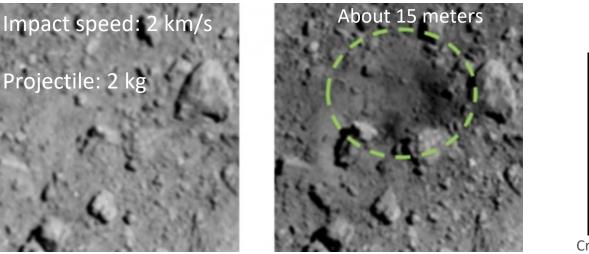
Hayabusa2 Small Carry-on Impactor April 5, 2019

Arakawa et al. 2020, Science

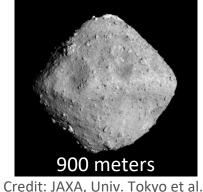


Cratering physics: Hayabusa2 versus Hera

Arakawa et al. 2020, Science



HAYABUSA2

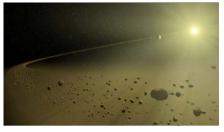


The large size of the crater can only be explain if the surface has no cohesion, which is a surprize! **Strong implications on surface age**

What about a 165 m-size body? And how does cratering scale with impact speed (from 2 to 6 km/s)?

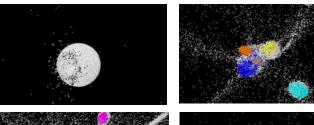
Understanding collisions in space

Planetesimal formation



Collisional accretion

EARLY PHASES





Michel & Richardson 2013

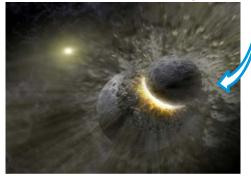
Collisional disruption

LATE PHASES

Planet formation

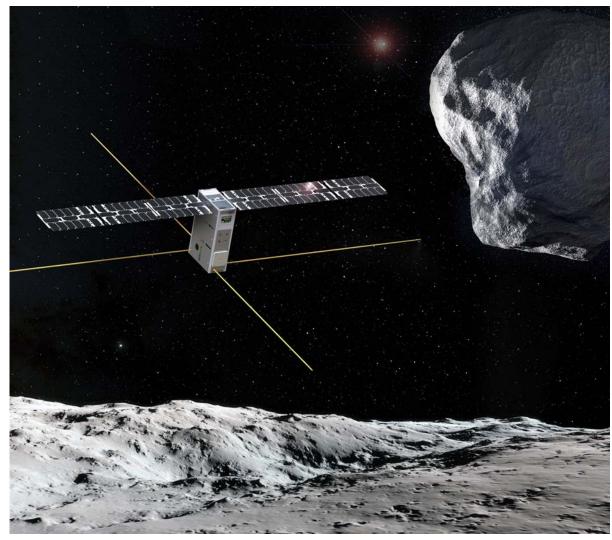


Giant impacts



Asteroid families







Juventas 6U Cubesat

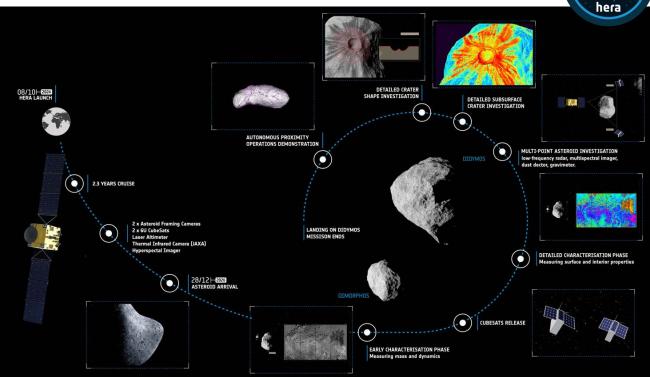
Landing on Dimorphos

- Bouncing properties
- Mechanical properties

Direct interaction with the surface is the only way to determine its often counter-intuitive response!

Hera: a mission of "firsts"

- First rendezvous with a binary asteroid and smallest asteroid ever visited
- First detailed measurements of asteroid cratering physics in the impact speed regime of asteroid collisions
- First deep-space Cubesats for very close asteroid inspection and first internal structure probing
- First Cubesat landing on a 160 m-size asteroid



Hera: planetary defense and science

Asteroids are fascinating:



- Sources of high interest of different communities (science, planetary defense, mining)
- Whichever is the primary objective to explore them (here, planetary defense), all those communities will benefit from the gained knowledge

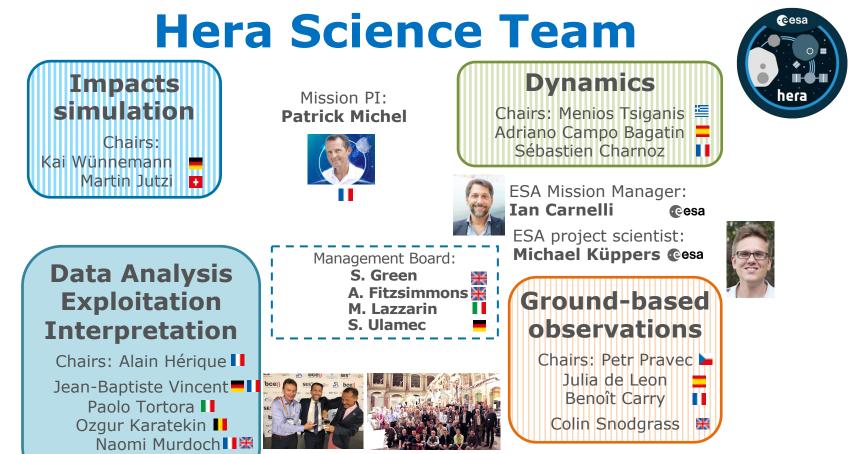
Hera is fascinating:

- First binary asteroid rendezvous and full characterisation, including subsurface/interior properties, and documentation of an impact outcome with NASA DART
- European contribution to planetary defense, concretizing ESA's pioneering initiative in the early 2000s starting with the Don Quijotte concept
- Great team and community actively working for it, promising amazing discoveries!









- + Operation Group
 - + Instrument PIs
- + International (US/Japan) members

Common AIDA (DART +Hera) Working groups

THANK YOU

