

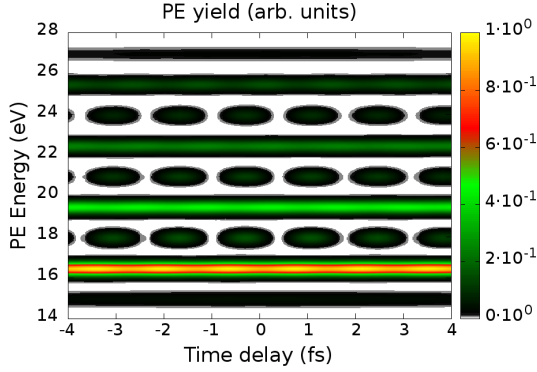
# RABBITT spectroscopy of transition-metal surface states

M. J. Ambrosio<sup>1</sup> and U. Thumm<sup>2</sup>

J. R. Macdonald Laboratory, Department of Physics, Kansas State University, Manhattan, Kansas 66506, USA.

**Synopsis** We distinguish contributions to the RABBITT phases from surface states of transition metals, due to Fresnel-reflection of the incident IR pulse, the high harmonics of the XUV pulse-train and the XUV photo release process.

We present numerical RABBITT (reconstruction of attosecond beating by interference of two-photon transitions) [1] spectra for photoemission from solids, the occupied surface states of Cu(111), Au(111), and Ag(111) (Fig. 1). We generate the transition-metal surface states based on an effective single-electron potential capable of reproducing the main characteristics of the substrate valence-band electronic surface [2].



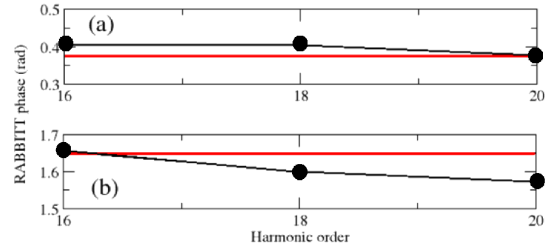
**Figure 1.** RABBITT spectrum from the Au(111) surface state.

Our numerical spectra show RABBITT phases to be strongly affected by the IR pulse Fresnel reflection at the solid-vacuum interface [3], which introduces a phase due to the photoelectron propagating through the overall (i.e., incident plus reflected) electric field. This Fresnel phase  $\phi_E$  largely determines the raw RABBITT phases

$$\phi_{2n}^{RAB} \equiv \phi_{2n-1}^+ - \phi_{2n+1}^- - \phi_{2n-1}^{(HH)} + \phi_{2n+1}^{(HH)} + 2\phi_E, \quad (1)$$

as shown in Fig. 2 for different incident directions and detection angles. Since it is well determined, the Fresnel phase contribution  $2\phi_E$  can be subtracted from the raw RABBITT phase  $\phi_{2n}^{RAB}$ , therefore leaving only the contributions from the HH phases  $\phi_{2n-1}^{(HH)}$  and  $\phi_{2n+1}^{(HH)}$ , of the XUV pulse train, and the scattering phases  $\phi_{2n-1}^+ - \phi_{2n+1}^-$  that are re-

lated to the ionization process.  $\phi_{2n-1}^+$  corresponds to the absorption of an XUV photon of energy  $(2n-1)\omega_{IR}$  plus the absorption of an IR photon, while  $\phi_{2n+1}^-$  stems from the absorption of an XUV photon with energy  $(2n+1)\omega_{IR}$  plus the release of an IR one. These scattering phase differences in our calculations are small, due to the very short propagation time of the photoelectron in the substrate.



**Figure 2.** RABBITT phases from the Ag(111) surface state (dots) and  $2\phi_E$  (red line). (a) Incidence angle  $30^\circ$ , with the XUV and IR polarizations tilted  $30^\circ$  from the incidence plane. Electron emission at  $28^\circ$  from the surface normal and  $18^\circ$  from the emission plane. (b) Incidence angle  $80^\circ$ , with the polarization axis tilted  $20^\circ$  from the incident plane, for normal emission.

Thus,

$$\phi_{2n}^{RAB} - 2\phi_E \approx -\phi_{2n-1}^{(HH)} + \phi_{2n+1}^{(HH)} \quad (2)$$

for the odd harmonics in the XUV pulse train can be used to estimate the unknown high harmonic phases  $\phi_{2n+1}^{(HH)}$ .

## Acknowledgements

We acknowledge support from NSF grant PHY-146 4417, the A. v. Humboldt foundation, and the Division of Chemical Sciences, Office of the Basic Energy Sciences, Office of Energy Research, US DoE.

## References

- [1] R. Locher et al. 2015 *Optica* **2** 405
- [2] E. V. Chulkov et al. 1999 *Surf. Sci.* **437** 330
- [3] M. J. Ambrosio et al. 1999 *Phys. Rev. A* **94** 063424

<sup>1</sup>E-mail: [mj\\_ambrosio@phys.ksu.edu](mailto:mj_ambrosio@phys.ksu.edu)

<sup>2</sup>E-mail: [thumm@phys.ksu.edu](mailto:thumm@phys.ksu.edu)